Observing Magnetized Shocks on OMEGA



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Acknowledgements

University of Michigan: <u>Joseph Levesque</u>, Carolyn Kuranz, Rachel Young, Mario Manuel (General Atomics), Sallee Klein, Gennady Fiksel, Matthew Trantham

Rice University: Patrick Hartigan, Andy Liao (LANL)

MIT: Chikang Li, Andrew Birkel (Proton Radiography)

LLE: Joseph Katz (Thomson Scattering)

This work is funded by the U.S. Department of Energy, through the NNSA-DS and SC-OFES Joint Program in High-Energy-Density Laboratory Plasmas, grant number DE-NA0002956, and the National Laser User Facility Program and William Marsh Rice University, grant number, R19071, and through the Laboratory for Laser Energetics, University of Rochester by the NNSA/OICF under Cooperative Agreement No. DE-NA0001944.

We have successfully observed magnetized bow shocks on OMEGA

- Campaign to create and observe astrophysically relevant laboratory-scale magnetized bow shocks
- Using a colliding plasma flow mechanism alongside MIFEDS we have achieved a β_{ram} regime in which magnetized shocks form
- We use proton radiography to measure the magnetic field topology
- We use the spatially resolved Imaging Thomson Scattering (ITS) diagnostic with 2ω probe beam to measure plasma properties across a shock
- By following reproducible features in the IAW spectra over multiple shots we find that the magnetic field affects the standoff distance of the shock from the wire

We want to create astrophysically relevant laboratory-scale magnetized bow shocks

- Magnetized bow shocks form when the incoming ram pressure of a plasma flow is equal to the magnetic pressure of the obstacle
- Define ram beta as ratio of ram pressure to magnetic pressure



$$egin{aligned} P_{ram} &=
ho u^2 & P_B &= rac{B_0^2}{2 \mu_0} \ & \ eta_{ram} &= rac{P_{ram}}{P_B} &= rac{
ho u^2}{B_0^2/2 \mu_0} pprox 1 \end{aligned}$$

Image credit: SOHO (NASA / ESA)

Reaching low- β_{ram} on OMEGA is challenging

- Previous attempts to achieve low- β_{ram} systems on OMEGA have been unsuccessful
 - MIFEDS capable of ~15 T maximum field for multi-mm-scale systems
 - Laser-produced plasmas have high ram pressure
- The ram pressure must be reduced for the current limitations



MIFEDS acts as the magnetized obstacle

- Straight, current-carrying wire is the magnetized obstacle
- Wire diameter: 0.762 mm
- Driven currents: 25 kA and 17 kA (or 13.5 T and 9 T max field at wire surface)



We use a multi-stage plasma source to reduce flow density and gradients

- Irradiate two counter-facing carbon foils
 - Diameter: 3.8 mm
 - Thickness: ~100 micron
- Collision redirects incoming flows outward from plane
- Expanding flow has lower density and velocity than constituents¹
- V ~ 100 km/s, ρ ~10⁻⁵ g/cc



¹ Liao et al. High Energy Density Physics 17 (2015)

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- This plasma acts as our solar wind



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We probe the system with proton radiography and ITS

- Imaging Thomson scattering (ITS) diagnostic measures scattered spectra
 - Centered 1.45 mm from wire
 - ~43° angle from primary flow axis
- Two probe configurations:
 - 20 J, 100 ps (no proton radiography)
 - 300 J, 1 ns (offset from proton driver)



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 - 20 J, 100 ps (no proton radiography)
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- D ³He proton source 1 cm from tcc
 - 3 and 15 MeV protons
- Protons flow antiparallel to wire current



We see a sharp density jump at 50 ns

- We fit the electron number density using EPW spectra
- Inferred plasma properties:
 - Unshocked: n_e=1 x 10¹⁸ cm⁻³
 - Shocked: n_e=12 x 10¹⁸ cm⁻³
 - Peak: n_e=24 x 10¹⁸ cm⁻³
 - Spike width ~ 0.1 mm
- Probe laser:
 - 20 J, 100 ps
- Detector:
 - 3 ns gate



We observe optical emission lines in the spectra



Optical emission spectra provides another measure of T_e

- Optical emission lines present in spectra
- PrismSPECT 0D, LTE line ratios match well for carbon at 5.5 eV
- This temperature should be of the unperturbed plasma
 - Line intensity comparable to Thomson scattered intensity due to long CCD gate
 - TS probe seems to greatly heat the plasma



The temperature measurements do not agree



Proton images show clear evidence of magnetic compression



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- Proton images are a deflection map of the magnetic topology
 - Dependent on:
 - proton velocity
 - magnetic field strength
 - magnetic field orientation
 - D³He proton source with CR39 detector²
 - 15 MeV image shown

Proton images also show the field evolution...

- $B_{max} = 9 T$, $E_p = 15 MeV$, Magnification = 15 X
- Dark bands away from wire indicate magnetic compression
 - We don't really know what the field structure across the shock should look like



... which changes with magnetic field strength

- $B_{max} = 13.5 T$, $E_p = 15 MeV$, Magnification = 15 X
- Shock feature appears farther from the wire at 60 ns



There is a clear difference in shock distance for the two field strengths

• The features move further out, but shock position is hard to decouple from deflectometry due to 3D effects



IAW spectra also shows sign of shock

- Observe simultaneous changes in density (EPW) and velocity (IAW)
- Shock coincides with a brief decrease in scattered energy



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The longer pulse duration changes spectral features

- The EPW spectra appears to be spread out
 - Emission lines no longer present
- The IAW spectra exhibit additional features coincident with shock
 - Not entirely sure what they represent
- We can use these features to test the effect of changing magnetic field



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There are interesting, reproducible features in the IAW spectra



We follow these features to measure how the magnetic field affects the standoff



Changing the magnetic field strength changes the standoff distance

		T
Field strength	Time (ns)	Shock Position (mm)
High	65	1.37
High	86	1.28
Low	66	1.30
Low	86	1.00

Table 1: IAW shock positions

- The high-field (13.5 T) shots have a greater standoff distance than the low-field (9 T) shots
- Additionally, the shock moves closer to the wire between 66 and 86 ns for the low-field case than the high-field

Changing the magnetic field strength changes the standoff distance... on both diagnostics



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Using ITS in tandem with proton radiography, we can infer magnetic field properties at the shock



- Generate synthetic proton images based on imposed magnetic field using ITS data
- Distance of shock from wire: 1.2 mm
- Estimated Shock depth (into page): 0.8 mm
 - Doesn't change center position much
- Assume no magnetic field behind shock
- The field jump at the shock is primary cause of the dark band(s)

MIFEDS acts as the magnetized obstacle

- Straight, current-carrying wire as the magnetized obstacle
 - Wire diameter: 0.762 mm
 - Driven currents: 25 kA and 17 kA (or 13.5 T and 9 T max field at wire surface)
- Multi-stage plasma source (collisional)
- Parameters at shock formation:
 - r_{shock} = 1.05 mm at t=50 ns
 - v ~ 100 km/s
 - n_e ~ 1e18 cm⁻³
 - T_e ~ 5.5 eV

