

Toroidal Equilibria with Transonic Poloidal Flow

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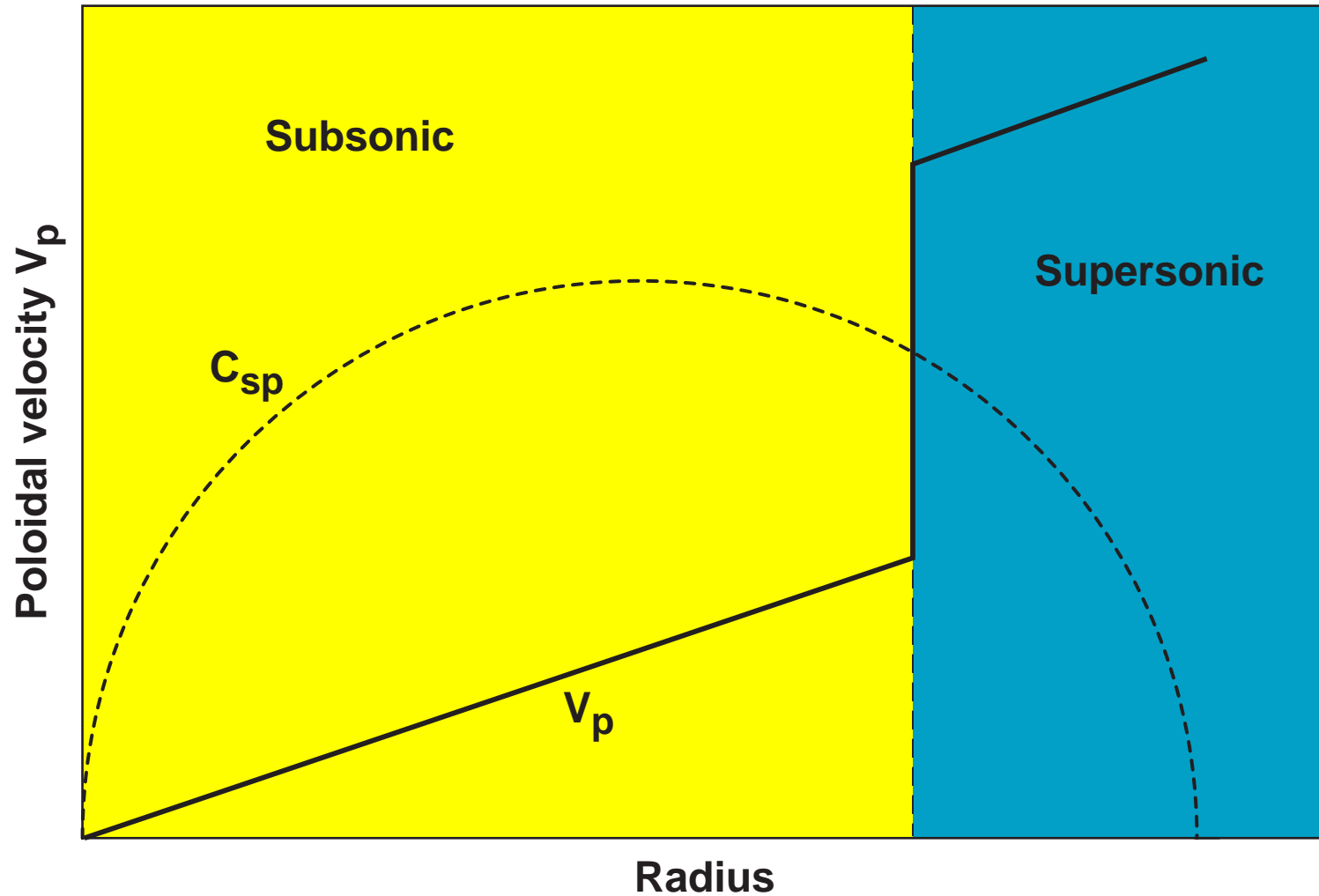
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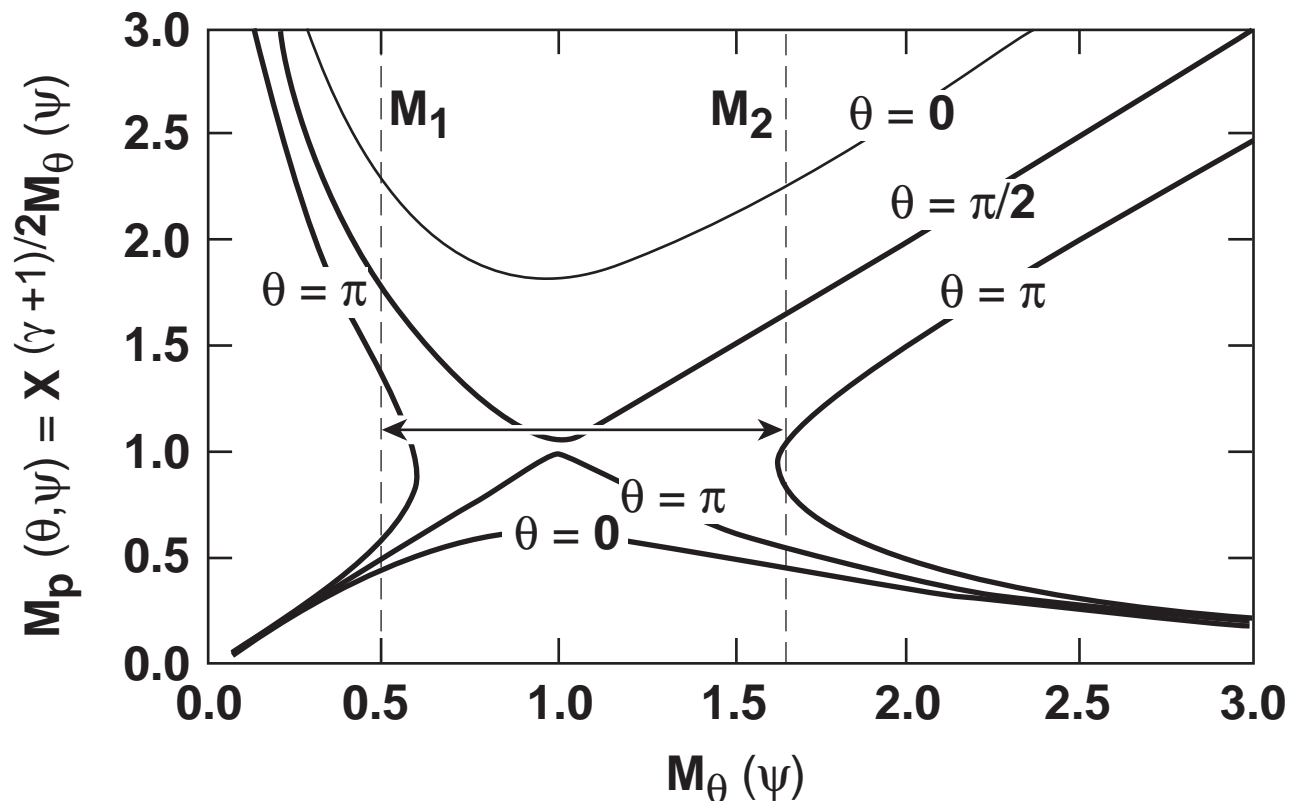
Summary

Transonic equilibria are radially discontinuous



The solution of the Bernoulli equation yields a radially discontinuous poloidal Mach number M_p at $\theta = \pi$

- Poloidal Mach number M_p versus free function $M_\theta(\psi)$.
- No solutions at $\theta = \pi$ for $M_1 < M_\theta < M_2$.
- The free function M_θ must be $< M_1$ (or $> M_2$).
- The poloidal Mach number must be radially discontinuous.



MHD equilibrium equations with arbitrary flow

$$\nabla \cdot (\rho \mathbf{v}^{\mathbf{r}}) = 0$$

$$\rho \mathbf{v}^{\mathbf{r}} \cdot \nabla \mathbf{v}^{\mathbf{r}} = -\nabla P + \mathbf{J}^{\mathbf{r}} \times \mathbf{B}^{\mathbf{r}} + \nabla v_{\theta} \nabla \mathbf{v}^{\mathbf{r}}$$

$$\nabla \cdot (\rho \mathbf{S} \mathbf{v}^{\mathbf{r}}) = v_{\theta} (\mathbf{e}_{\theta} \cdot \nabla \mathbf{v}_{\theta})^2 / T$$

$$\mathbf{p} = (1 + Z) \rho T / m_i$$

$$\mathbf{S} = c_v \log(\mathbf{p} / \rho^{\gamma})$$

$$\mu_0 \mathbf{J}^{\mathbf{r}} = \nabla \times \mathbf{B}^{\mathbf{r}}$$

$$\nabla \cdot \mathbf{B}^{\mathbf{r}} = 0$$

$$\nabla \times (\mathbf{v}^{\mathbf{r}} \times \mathbf{B}^{\mathbf{r}}) = 0$$

Look for continuous solutions in $\theta \rightarrow$ neglect viscosity;
the MHD equilibrium depends on five free functions

$$\nabla \cdot \mathbf{B} = 0 \rightarrow \mathbf{B}_p = \frac{1}{R} \nabla \Psi \times \mathbf{e}_\phi$$

$$\nabla \times (\mathbf{v} \times \mathbf{B}) = 0 \rightarrow \mathbf{v} = v_{||} \mathbf{b} - \Omega(\Psi) R \mathbf{e}_\phi$$

$$\nabla \cdot \rho \mathbf{v} = 0 \rightarrow \rho = D(\psi) / (v_{||} / B)$$

$$\nabla \cdot (\rho \mathbf{Sv}) = 0 \rightarrow \nabla \cdot \mathbf{p}^{1/\gamma} \mathbf{v} = 0 \rightarrow \mathbf{p} = Q(\psi) / (v_{||} / B)$$

$$\phi\text{-momentum} \rightarrow \mathbf{B}_\phi R = F(\Psi) + D(\Psi) v_\phi R$$

$$\mathbf{v}\text{-momentum:} \rightarrow \rho \left[\frac{v^2}{2} + \Omega(\Psi) R v_\phi \right] + \frac{\gamma}{\gamma - 1} \mathbf{p} = \frac{N(\Psi)}{(v_{||} / B)}$$

\rightarrow $D(\psi)$, $Q(\psi)$, $F(\psi)$, $\Omega(\psi)$, $N(\Psi)$ are free functions
of Ψ determined by the transport equations.

$\nabla\psi$ -component of momentum = Grad-Shafranov



$$\Delta^* \psi + \mathbf{e}_\psi \cdot \nabla \left(\frac{\mathbf{B}_\phi^2 R^2}{2} \right) + R^2 \mathbf{e}_\psi \cdot \nabla p +$$
$$- \left(\rho v_\phi^2 / \mathbf{B}_p^2 \right) \mathbf{B}_z = \mathcal{O} \left(\rho v_p^2 / \mathbf{B}_p^2 \right) \sim \epsilon^2$$

Ordering:

$$\epsilon = a/R_0 < 1$$

$$\beta \sim \epsilon^2, \mathbf{B}_p / \mathbf{B}_\phi \sim \epsilon,$$

$$\mathbf{v}_p \sim \mathbf{C}_s,$$

$$\mathbf{v}_\phi \sim (\epsilon \mathbf{C}_s \text{ or } \mathbf{C}_s)$$

$$\mathbf{C}_{sp} = \frac{\mathbf{B}_p}{\mathbf{B}} \mathbf{C}_s = \text{poloidal sound speed}$$

$$\mathbf{M}_p = \frac{\mathbf{v}_p}{\mathbf{C}_{sp}} \sim 1 = \text{poloidal Mach number}$$

The Bernoulli equation must be solved to find the poloidal Mach number M_p

Definitions:

$$\rho = \mathbf{D}(\psi)/\mathbf{X}$$

$$\mathbf{p} = \mathbf{P}(\psi)/\mathbf{X}^\gamma$$

$$\Delta^* \psi + \mathbf{e}_\psi \cdot \nabla \left(\mathbf{R}^2 \mathbf{p} + \mathbf{B}_\phi^2 \mathbf{R}^2 / 2 \right) = \mathbf{O}(\epsilon)$$

$$\mathbf{C}_s(\psi) = \sqrt{\gamma \mathbf{P}(\psi) / \mathbf{D}(\psi)}$$

$$\mathbf{v}_p = \mathbf{X} \mathbf{C}_s(\psi) \mathbf{M}_\theta(\psi) \mathbf{B}_p / \mathbf{B}$$

$$\mathbf{v}_\phi \cong \mathbf{C}_s(\psi) \left[\mathbf{M}_\phi(\psi) + (\mathbf{X} - 1) \mathbf{M}_\vartheta(\psi) \right]$$

$$\mathbf{B}_\phi \mathbf{R} = \mathbf{F}(\psi) \left[1 + \mathbf{O}(\epsilon^2 \mathbf{X}) \right]$$

Bernoulli equation

$$\mathbf{R}^{-2} \mathbf{M}_\vartheta^2(\psi) \mathbf{X}^{\gamma+1} - \left\{ \frac{2}{\gamma-1} + \mathbf{M}_\vartheta^2(\psi) + \left[\mathbf{M}_\vartheta(\psi) - \mathbf{M}_\phi(\psi) \right] (\mathbf{R}^2 - 1) \right\} \mathbf{X}^{\gamma-1} + \frac{2}{\gamma-1} = 0$$

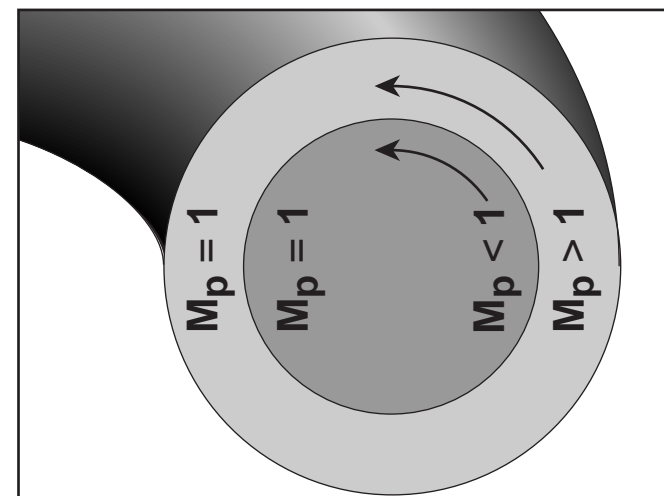
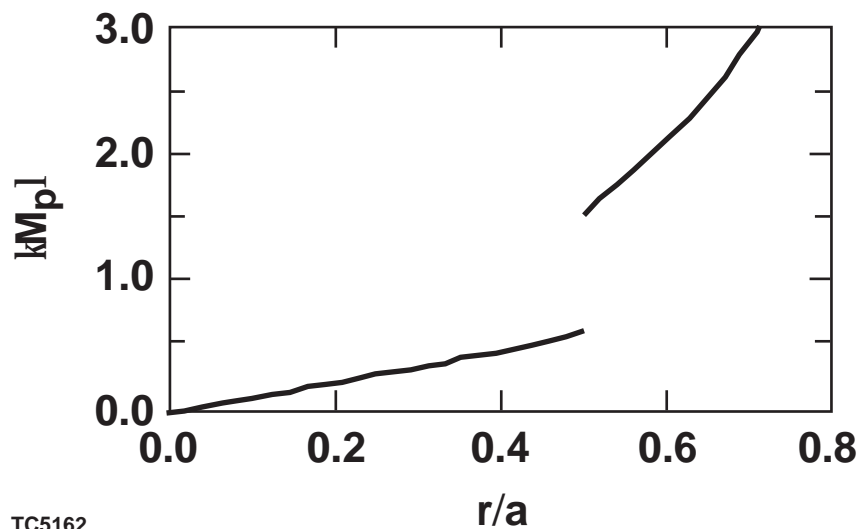
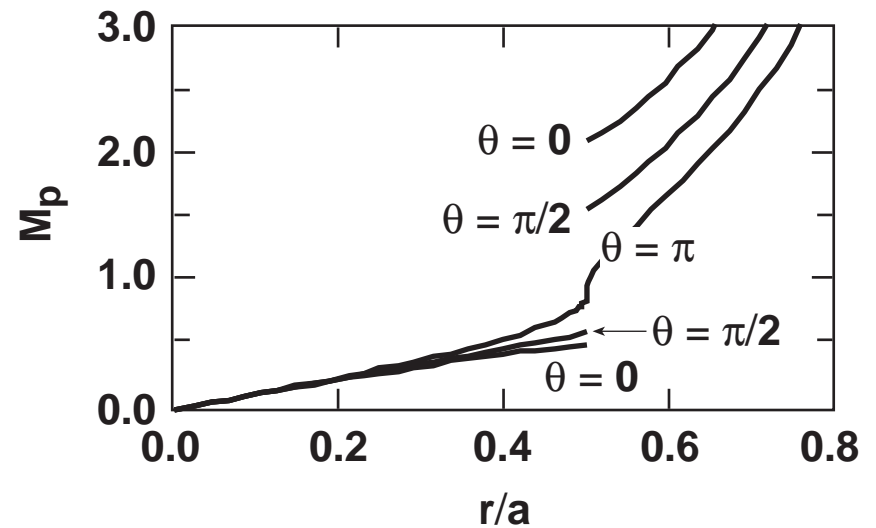
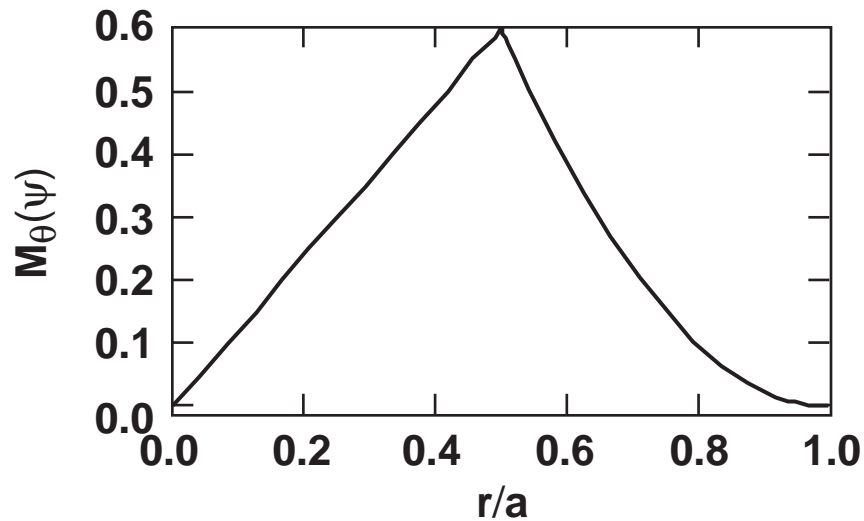
$$\mathbf{R} = 1 + \epsilon \cos \vartheta$$

$$\mathbf{X} = \mathbf{X}(\psi, \vartheta)$$

$$\mathbf{M}_p = \mathbf{X}^{(\gamma+1)/2} \mathbf{M}_\theta(\psi)$$

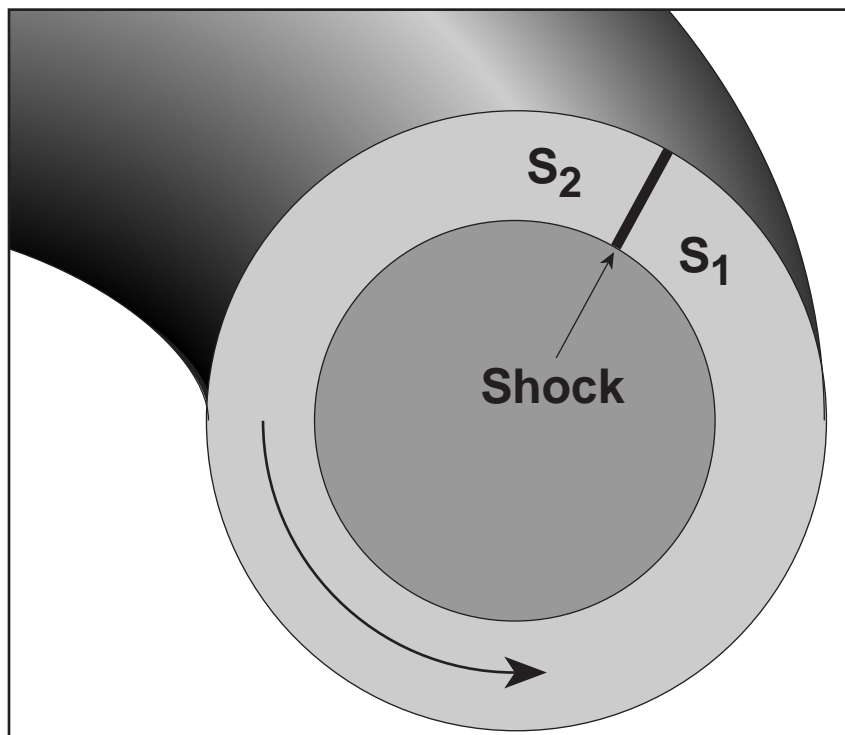
$\mathbf{P}(\psi)$, $\mathbf{D}(\psi)$, $\mathbf{F}(\psi)$, $\mathbf{M}_\phi(\psi)$, and $\mathbf{M}_\theta(\psi)$ are new free functions.

Shockless transonic equilibria have discontinuous Mach number profiles



Shocked solutions do not represent “real” equilibria

- Sources of irreversibility cannot be stationary in a periodic system.



S = entropy

$$S_2 > S_1$$

$$= \bullet (\rho S \vec{v}) = v_q (\mathbf{e}_q \bullet = v_q)^2 / T$$

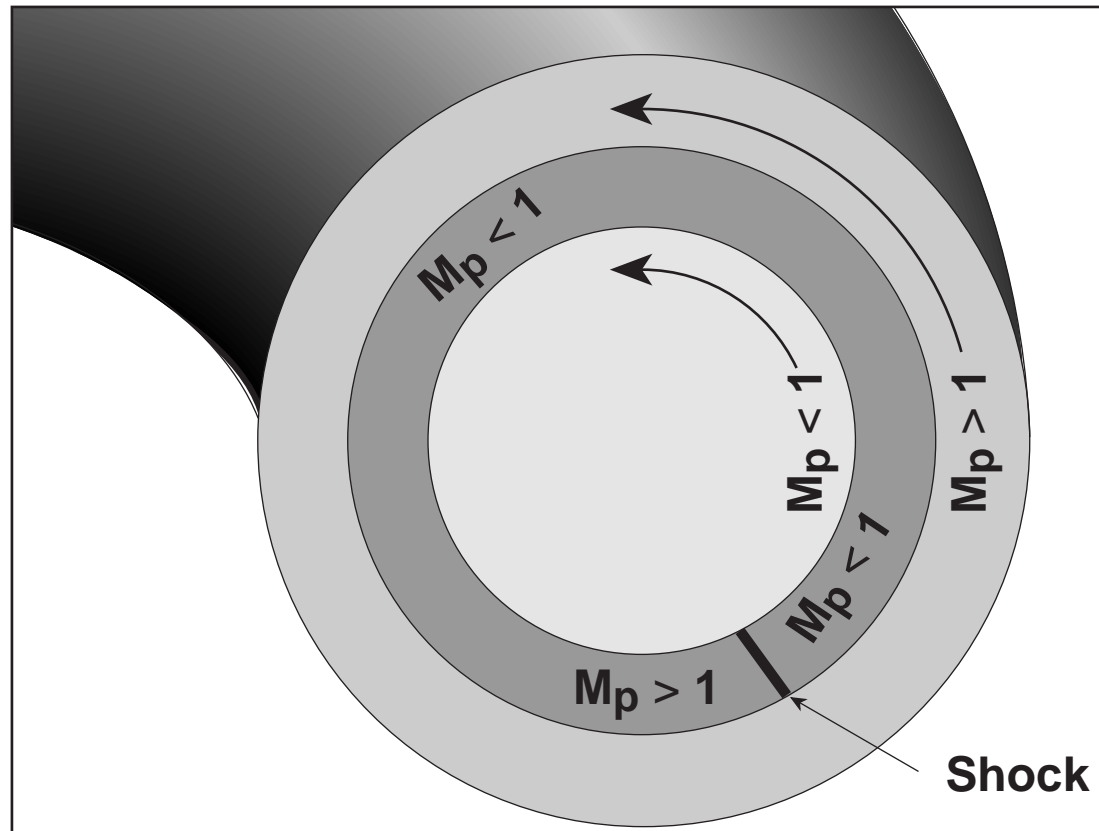
$$\Delta S \sim \epsilon^{3/2} S$$

- If the shock were stationary, the entropy would constantly increase.

- The shock must travel: $\theta_s = \theta_s(t)$

- The shock velocity is slow: $\dot{\theta}_s(t) \sim \epsilon^{1/2} \frac{v_\theta}{r}$

Can equilibria with shocks exist?



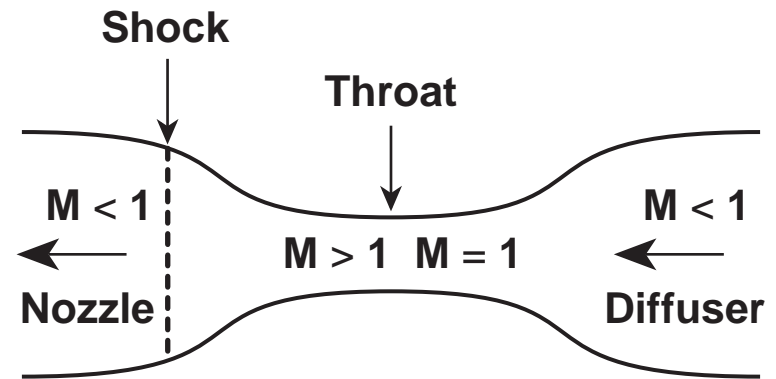
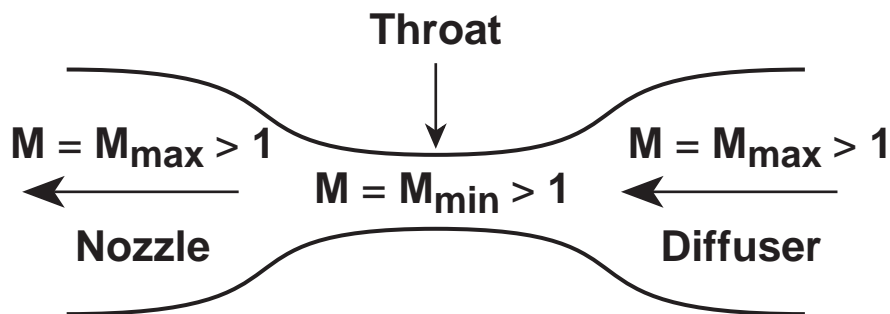
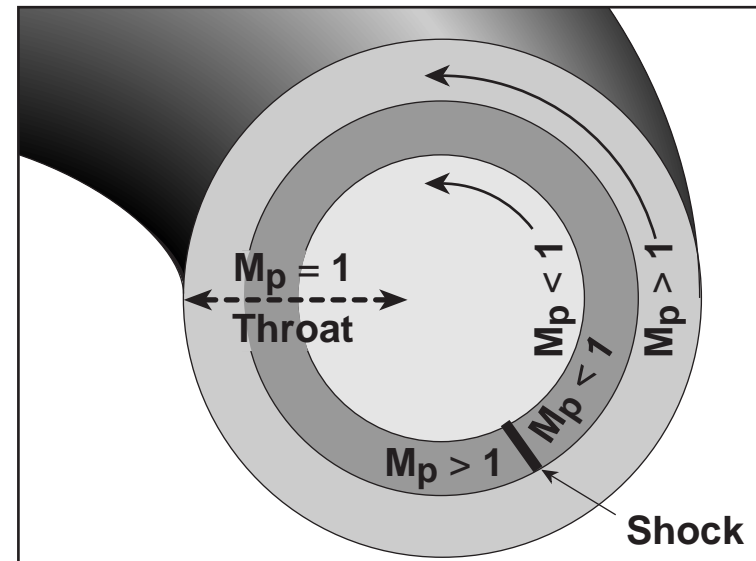
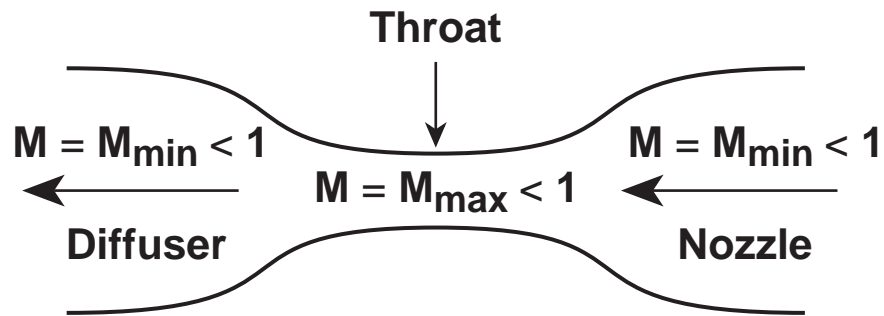
Hazeltine, Lee, Rosenbluth (1971); Shaing, Hazeltine, and Sanuki (1992).

Shock dynamics

- **Shocks cannot persist at a steady state.**
- **If a shock forms, it weakens while moving poloidally toward the inner midplane.**
- **The shock vanishes after reaching the inner midplane.**
- **Traveling shocks leave behind the radially discontinuous equilibria previously described.**

The tokamak magnetic flux surfaces act as a nozzle/diffuser

- The flow is either “well” supersonic or “well” subsonic.



Summary

- In the presence of transonic poloidal flow, the isentropic MHD equations yield two classes of equilibria:
 - shockless equilibria with radially discontinuous profiles and
 - equilibria with weak shocks.
- Because of the finite irreversibility, the isentropic MHD model does not correctly describe shocks (weak or strong) in a periodic system.
- The energy-conserving (including the viscous heating) MHD model shows that shocked equilibria evolve into shockless radially discontinuous equilibria.

The shock trajectory $\theta_s(t)$ can be determined by the solvability conditions of the time-dependent MHD equations

Example:

$$\nabla \cdot (\rho \mathbf{v}_p) = -\partial_t \rho \quad \partial_t \sim \epsilon^{1/2} \frac{v_p}{r} \quad \rho = \rho_0 + \epsilon^{1/2} \rho_{1/2} + \epsilon \rho_1 \dots$$

$$\mathbf{v}_p = \mathbf{v}_{p0} + \epsilon^{1/2} \mathbf{v}_{p(1/2)} + \epsilon \mathbf{v}_{p1} \dots$$

$$\left\{ \begin{array}{l} \epsilon^0 \text{ - order} \\ \nabla \cdot (\rho_0 \mathbf{v}_{p0}) = 0 \rightarrow \mathbf{B}_p \cdot \nabla (\rho_0 \mathbf{v}_{p0} / \mathbf{B}_p) = 0 \rightarrow \rho_0 \mathbf{v}_{p0} / \mathbf{B}_p = \mathbf{D}_0[\psi, \theta_s(t)] \end{array} \right.$$

$$\left\{ \begin{array}{l} \epsilon^{1/2} \text{ - order} \\ \nabla \cdot [\rho_{1/2} \mathbf{v}_{p0} + \rho_0 \mathbf{v}_{p(1/2)}] = -\partial_t \rho_0 \\ \mathbf{B}_p \cdot \nabla [(\rho_{1/2} \mathbf{v}_{p0} + \rho_0 \mathbf{v}_{p(1/2)}) / \mathbf{B}_p] = -\partial_t \rho_0 \end{array} \right. \longrightarrow \int_{-\pi}^{+\pi} \frac{r \partial_t \rho_0}{\mathbf{B}_p \cdot \nabla \theta} d\theta = 0$$

$$\left\{ \begin{array}{l} \epsilon^1 \text{ - order} \\ \dots \end{array} \right.$$

Shock poloidal trajectory

$$\frac{d\theta_s}{dt} = \frac{\pi}{3} [2(\gamma + 1) \epsilon]^{1/2} \frac{C_{sp}}{r} \frac{\cos^3(\theta_s/2)}{\sin\theta_s}$$

- If $\theta_s(0) > 0$, then $\dot{\theta}_s(t) > 0$ and the shock travels counterclockwise.
- If $\theta_s(0) < 0$, then $\dot{\theta}_s(t) < 0$ and the shock travels clockwise.
- The shock stops when it reaches the inward midplane $\theta_s = \pm\pi$.
- The shock disappears at the inward midplane:

$$\Delta M_p|_{\pm\pi} = 2\sqrt{2(\gamma + 1) \epsilon} \cos\left(\frac{\theta_s}{2}\right)\Big|_{\pm\pi} = 0$$

Shocks in the upper/lower mid-plane travel leaving behind a super/subsonic flow

