









Extended Magnetohydrodynamics in the FLASH Code



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- FLASH is a multi-physics magnetohydrodynamics (MHD) code that includes adaptive mesh refinement (AMR), radiation, laser ray-tracing, extended MHD effects, and more. As part of the extended MHD capabilities, Hall and Biermann terms are already in the release version of FLASH.
- Recent code development under the auspices of ARPA-E and DOE NNSA has greatly increased FLASH's ability to model a broader range of magnetized plasma regimes.
 - Anisotropic thermal conduction important for any problem in which heat transport is affected by magnetic fields (typically when Hall parameter ≥ 1).
 - Anisotropic magnetic resistivity the effects of anisotropy are important for strongly-magnetized plasmas, and the implicitly-solved implementation is crucial for studying problems with high resistivity (e.g., Z-pinches and cold plasmas).
 - Thermoelectric effects includes Seebeck, Righi-Leduc, and Nernst terms. We know these effects can play an important role in magneto-inertial fusion concepts.



The FLASH code



- □ *FLASH* is a publicly available, high performance computing (HPC), adaptive mesh refinement (AMR), finite-volume, hydro and MHD code with extended physics capabilities. Supported primarily by the U.S. DOE NNSA.
- FLASH is professionally managed software in continuous development for 20 years: coding standards; version control; daily automated regression testing; extensive documentation; user support; integration of extensive code contributions from external users.

> 3,500 users world wide

>1,200 papers published with FLASH







Ideal MHD uses a relatively simple Ohm's Law in which magnetic field is "frozen-in" and must follow motion of plasma.

 $\mathbf{E} = -\mathbf{u} \times \mathbf{B}$

Extended MHD uses a much more complex generalized Ohm's Law, which takes on this form in FLASH with logical switches to turn on/off each term:

$$\begin{split} \mathbf{E} &= -\mathbf{u} \times \mathbf{B} + \frac{\mathbf{J}}{n_e e} \times \mathbf{B} - \frac{\nabla P_e}{n_e e} \\ &+ \frac{\eta_{\parallel}}{\epsilon_0 \omega_{pe}^2 \tau_{ei}} \mathbf{b} (\mathbf{b} \cdot \mathbf{J}) + \frac{\eta_{\perp}}{\epsilon_0 \omega_{pe}^2 \tau_{ei}} \mathbf{b} \times (\mathbf{J} \times \mathbf{b}) - \frac{\eta_{\wedge}}{\epsilon_0 \omega_{pe}^2 \tau_{ei}} (\mathbf{b} \times \mathbf{J}) \\ &- \beta_{\parallel} \mathbf{b} (\mathbf{b} \cdot \nabla T_e) - \beta_{\perp} \mathbf{b} \times (\nabla T_e) \times \mathbf{b}) - \beta_{\wedge} (\mathbf{b} \times \nabla T_e) \end{split}$$







$$\mathbf{E} = -\mathbf{u} \times \mathbf{B} + \frac{\mathbf{J}}{n_e e} \times \mathbf{B}$$

- □ Hall term is implemented via an explicit flux-based method, and the velocity used to determine the computational time step must now include the current.
 - Dispersive Whistler waves and Hall drift waves are introduced (wave speeds can also affect time step on scales smaller than ion inertial length).
- Plasma opening switch shows how including the Hall term leads to drastically different B-field behavior.









$$\mathbf{E} = -\mathbf{u} \times \mathbf{B} - \frac{\nabla P_e}{n_e e}$$

- □ *FLASH* has two implementations of the Biermann term:
 - An explicit flux-based method similar to the method for the Hall term.
 - A source term method in which the Biermann term is not included as a magnetic flux (appropriate for 1D and 2D cylindrical problems).
- □ Thermal-magnetic waves are introduced (wave speed proportional to sound speed).
- Laser-ablated CH foil shows Biermann-generated fields (credit: Yingchao Lu).







□ Isotropic case is rather simple: $\frac{\partial e_{int}}{\partial t} + \nabla \cdot (-\kappa \nabla T) = 0.$

Anisotropic case can be written in terms of three conductivity coefficients:

$$\frac{\partial e_{int}}{\partial t} + \nabla \cdot (\kappa_{\parallel} \nabla T (\kappa_{\parallel} \kappa_{\perp}) \mathbf{b} (\mathbf{b} \cdot \nabla T) \kappa_{\wedge} (\mathbf{b} \times \nabla T)) = 0$$

 In FLASH, this is solved implicitly with an iterative matrix inversion method using LLNL's HYPRE library (Falgout+ 2004, 2005).

 $\kappa_{\perp} < \kappa_{\parallel}$ implies thermal diffusion is hindered perpendicular to magnetic fields.

Poster from 2019 APS DPP meeting showcases analytic test, verifying the code's implementation.





- In the experiment, magnetized plasma jets collide to form a target region and additional, heavier jets form a liner that compresses the target.
- In these 2D cylindrical simulations, two target-forming jets collide, and we observe different temperature profiles when using isotropic vs anisotropic thermal conduction.
- Azimuthal B-field peaks at accretion shocks that surround target, so when modeling with anisotropic thermal conduction, this creates a hot shell.
- Testing how thermal conduction behaves in different field configurations during target compression phase.
 - See Chuang Ren's talk Wed. for more details about our PLX collaboration.







Anisotropic magnetic resistivity



$$\mathbf{E} = -\mathbf{u} \times \mathbf{B} + \frac{\eta_{\parallel}}{\epsilon_0 \omega_{pe}^2 \tau_{ei}} \mathbf{b} (\mathbf{b} \cdot \mathbf{J}) + \frac{\eta_{\perp}}{\epsilon_0 \omega_{pe}^2 \tau_{ei}} \mathbf{b} \times (\mathbf{J} \times \mathbf{b}) - \frac{\eta_{\wedge}}{\epsilon_0 \omega_{pe}^2 \tau_{ei}} (\mathbf{b} \times \mathbf{J})$$

- Like thermal conduction, this is solved implicitly with HYPRE, but it is a much more complicated system of equations with coupled B-field components.
- Using the Poynting flux and conservation of energy, Ohmic heating can be included in a way that is consistent with the implicit magnetic diffusion.
- $\hfill\square$ η_{Λ} term is added to the explicit Hall flux.

Implicit solve can save orders of magnitude in computational time for problems with high resistivity.

Poster from 2021 NIF/JLF UGM describes the procedure for solving this system implicitly.











- High-Z liner is imploded with an applied axial current and compresses a DT fuel.
- □ *FLASH*'s implicit resistivity solver implementation has been verified by comparing 1D cylindrical models with results from the *MACH2* code.

30

25

20

15

10

5

120

100

140

Shell trajectory from a simplified model without radiation.

inner radius outer radius

curren

10000

1000

100

10

0

20

40

60

80

Time (ns)

Radius (µm)





R-Axis (cm)



Ruskov+ (Phys. Plasmas 2020)

Ohmic heating at liner/vacuum interface decreases resistivity → prevents B-field from diffusing into liner.







- $\mathbf{E} = -\mathbf{u} \times \mathbf{B} \beta_{\parallel} \mathbf{b} (\mathbf{b} \cdot \nabla T_e) \beta_{\perp} \mathbf{b} \times (\nabla T_e) \times \mathbf{b}) \beta_{\wedge} (\mathbf{b} \times \nabla T_e)$
- Terms can be rewritten into β_{\parallel} , (β_{\parallel} β_{\perp}), and β_{\wedge} terms. Nomenclature varies in literature, but in *FLASH*, these terms are called:
 - $\circ \beta_{\parallel}$ = Seebeck
 - $(\beta_{\parallel} \beta_{\perp})$ = Righi-Leduc. Replaced with β_{new} for Z ≥ 1 (Davies+, Phys. Plasmas 2021)
 - $\circ \beta_{\Lambda} = Nernst$
- Additionally, all terms appear in the electron heat flux: $\boldsymbol{q} = -\frac{T_e}{\rho} \boldsymbol{\beta} \cdot \boldsymbol{j}$
- All terms are implemented with an explicit flux-based approach.

Currently in the process of testing the Nernst term with a KHI problem (credit: Yingchao Lu).



We know Nernst plays an important role in some MIF concepts like mini-MagLIF (Hansen+, Phys.

Plasmas 2020).







We have implemented some of the most up-to-date transport coefficients from the literature:

- Thermal conductivity coefficients (κ_{\parallel} , κ_{\perp} , κ_{\wedge}) of Ji & Held (Phys. Plasmas 2013).
- Resistivity coefficients (η_{\parallel} , η_{\perp} , η_{\wedge}) and thermoelectric coefficients (β_{\parallel} , β_{\perp} , β_{\wedge}) of Davies+ (Phys. Plasmas 2021).

Home > Physics of Plasmas > Volume 28, Issue 1 > 10.1063/5.0023445 Full . Published Online: 25 January 2021 Accepted: December 2020 Transport coefficients for magnetic-field evolution in inviscid

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