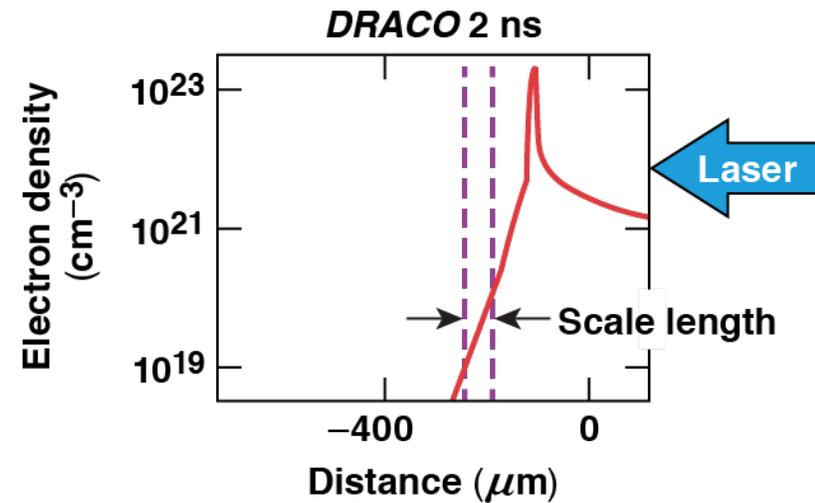
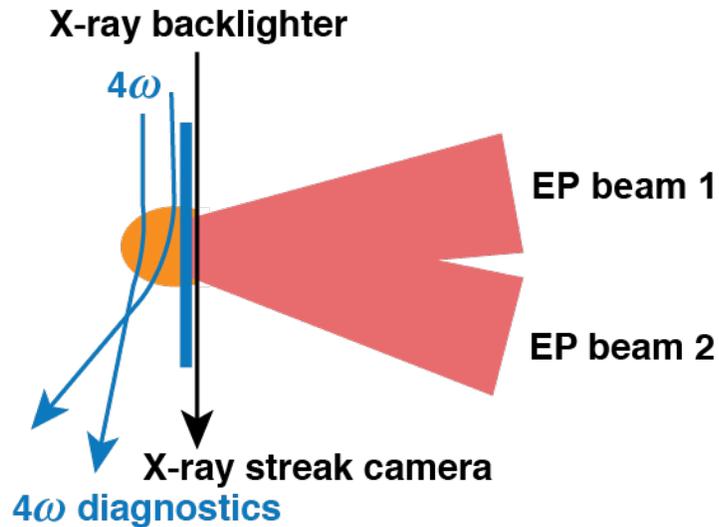


# Shock-Release Experiments on OMEGA EP



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the American Physical Society  
Division of Plasma Physics  
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# Radiation transport strongly affects the expansion of the rarefaction wave formed by the shock released from the laser-driven CH foil

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- The buried Au layer reduced the extent and scale length of the density profile in the rarefaction wave, consistent with radiation preheat hypothesis
- The Al layer on the laser side of the target did not affect the release, which ruled out the laser shinethrough as the early expansion mechanism

\* D. Haberberger *et al.*, Phys. Rev. Lett. **123**, 235001 (2019).  
VISAR: velocity interferometer system for any reflector

# Collaborators

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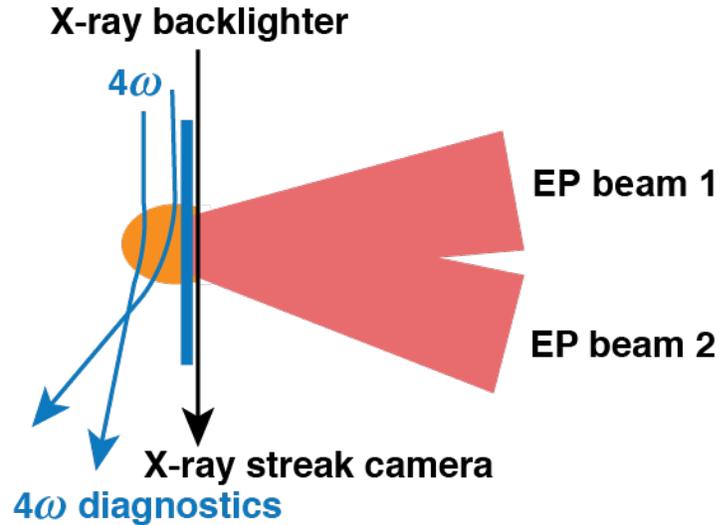


**D. Haberberger, J. P. Knauer, S. X. Hu, S. T. Ivancic, J. Carroll-Nellenback,  
D. Cao, I. V. Igumenshchev, V. V. Karasiev, P. B. Radha, A. V. Maximov, S. P. Regan,  
T. C. Sangster, R. Boni, P. M. Nilson, V. N. Goncharov, and D. H. Froula**

**University of Rochester  
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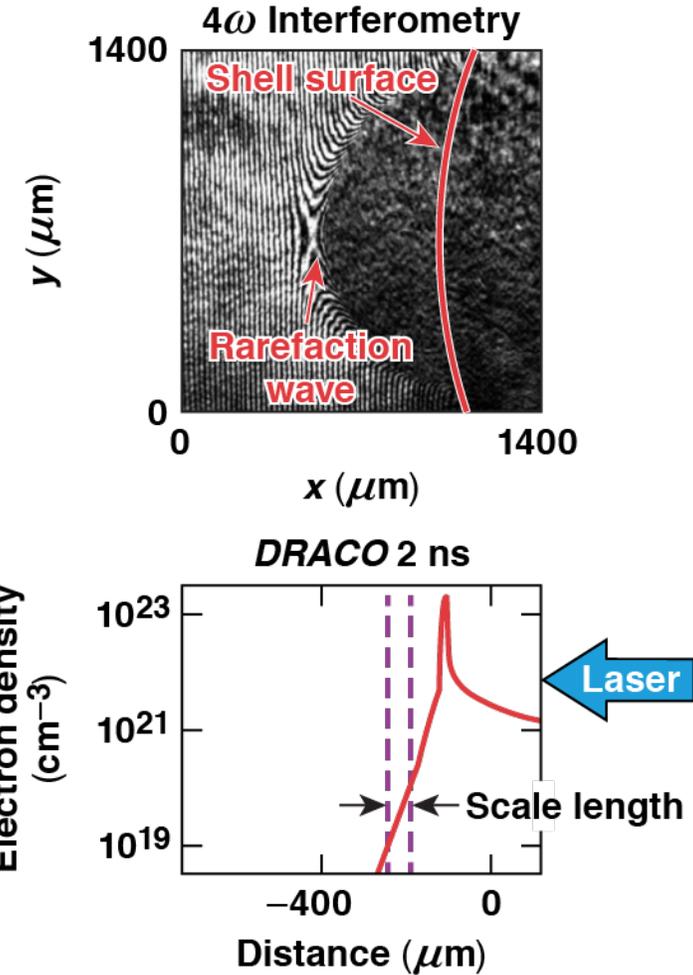
**M. D. Rosen and V. A. Smalyuk  
Lawrence Livermore National Laboratory**

# The shock-release experiments used a $4\omega$ probe to measure the low-density plasma profile and side-on x-ray radiography to measure the shell trajectory



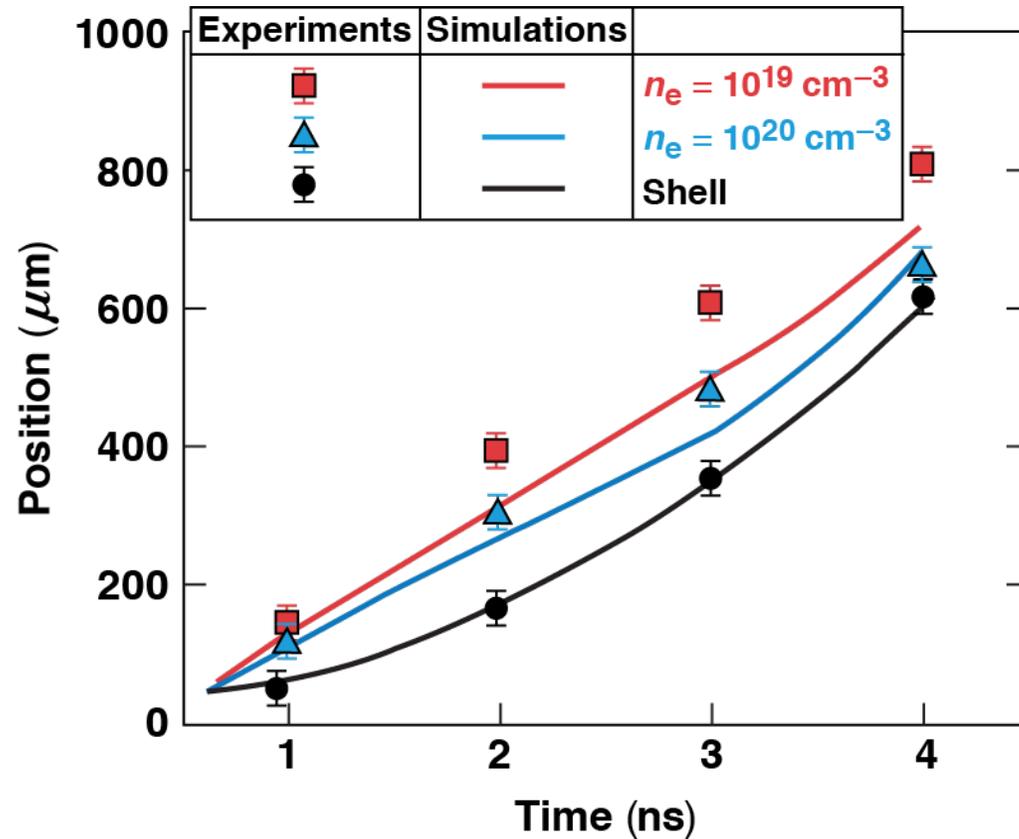
TC14640b

- 37  $\mu\text{m}$  CH
- 4.1-mm-diam spherical cap or flat
- 5-ns square pulse
- $3 \times 10^{14}$  W/cm<sup>2</sup>



TC14642b

# The experimentally measured extent and scale length of the electron density profiles in the rarefaction wave are significantly longer than those predicted by standard radiation-hydrodynamics simulations

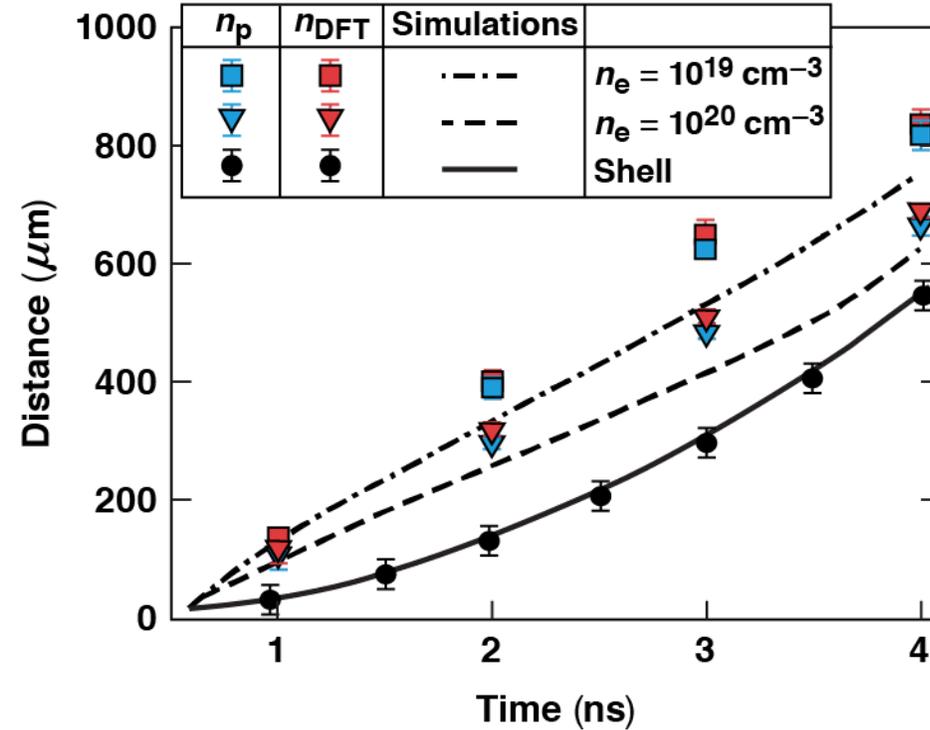
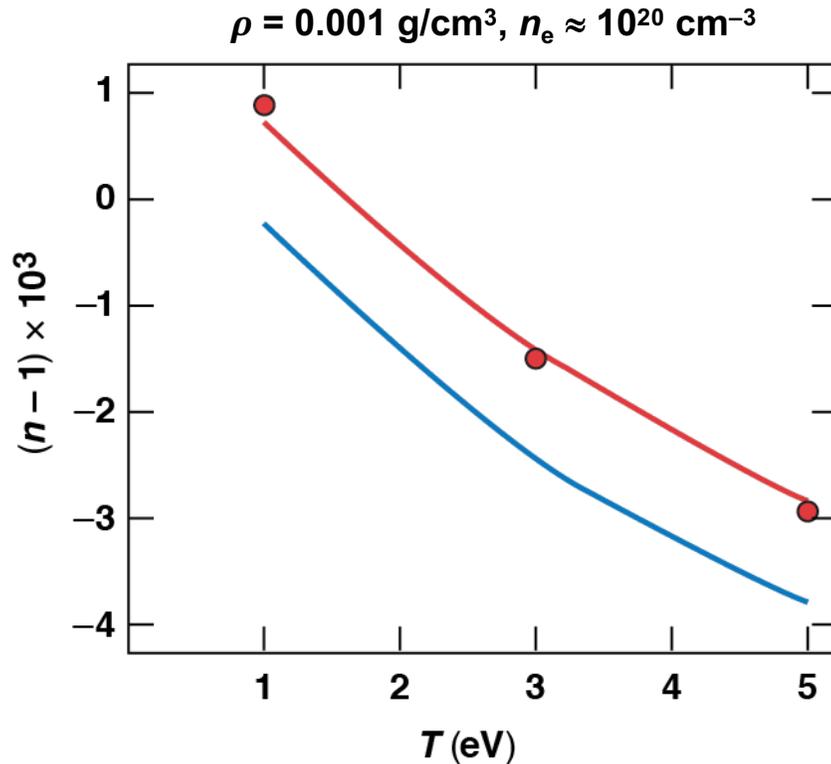


## The reasons for the difference

- The way we infer the electron density using a textbook plasma index of refraction is wrong
- Some physics is missing from or incorrect in the simulations

TC14644b

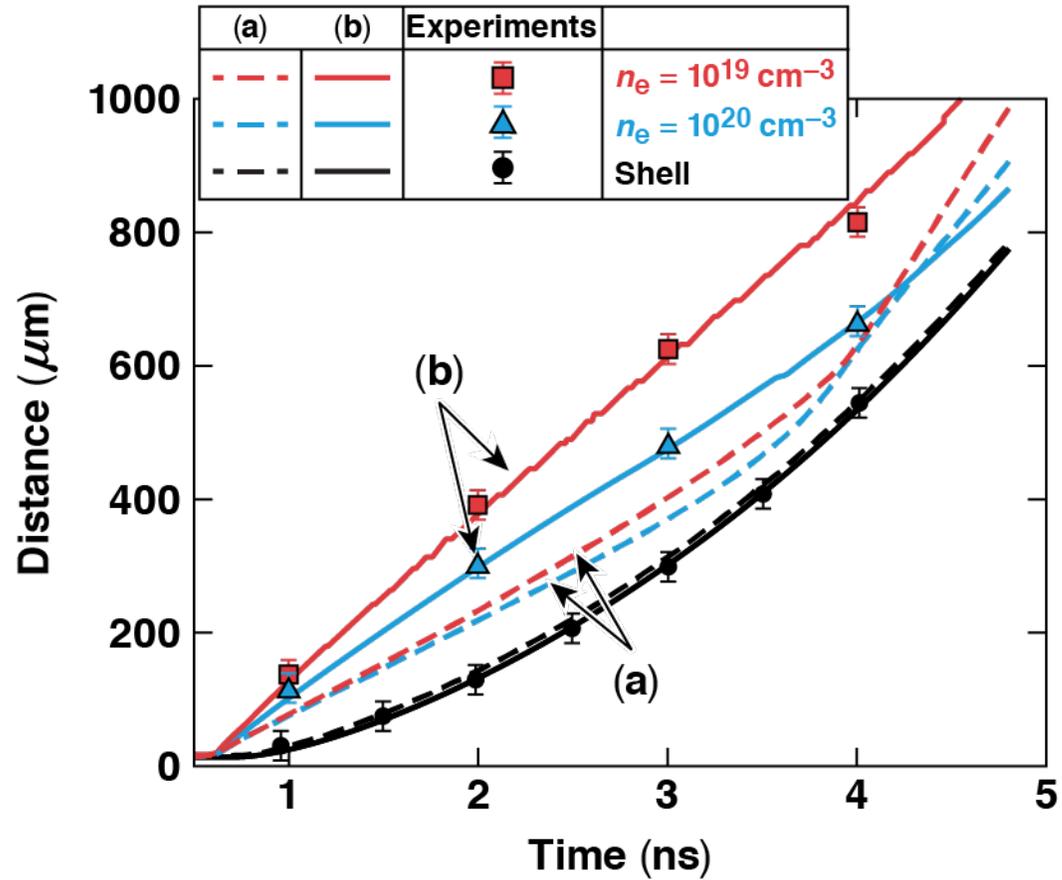
# Using a more-accurate index of refraction from *ab-initio* calculations\* does not significantly change the electron density profiles obtained from $4\omega$ probe images



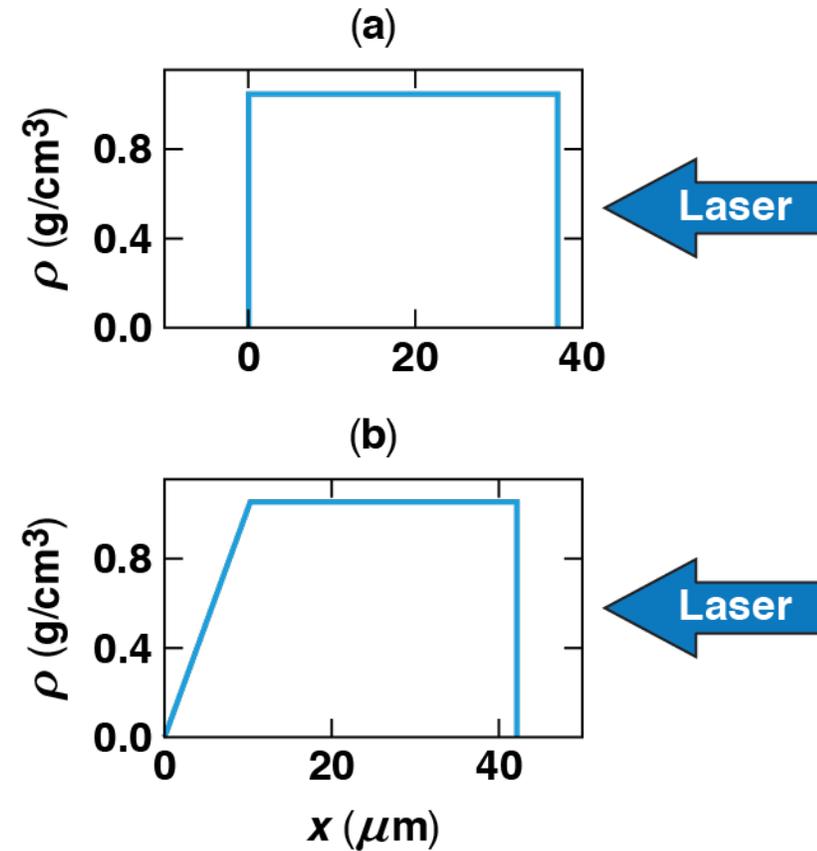
—  $n_{\text{DFT}} = \sqrt{1 + 4\pi(1.7 \text{ \AA}^3 \times ni - 4.9 \text{ \AA}^3 \times n_e)}$  – DFT *ab-initio* calculations  
—  $n_p = \sqrt{1 - 4\pi 4.9 \text{ \AA}^3 \times n_e}$  – plasma dielectric constant  $n_p = \sqrt{1 - n_e/n_c}$

\* A. Shvydky, A. V. Maximov, V. V. Karasiev, D. Haberberger, and V. N. Goncharov, "Ionization State and Dielectric Permittivity in Cold Rarefied CH Plasmas of Inertial Confinement Fusion," to be submitted to Physical Review E.

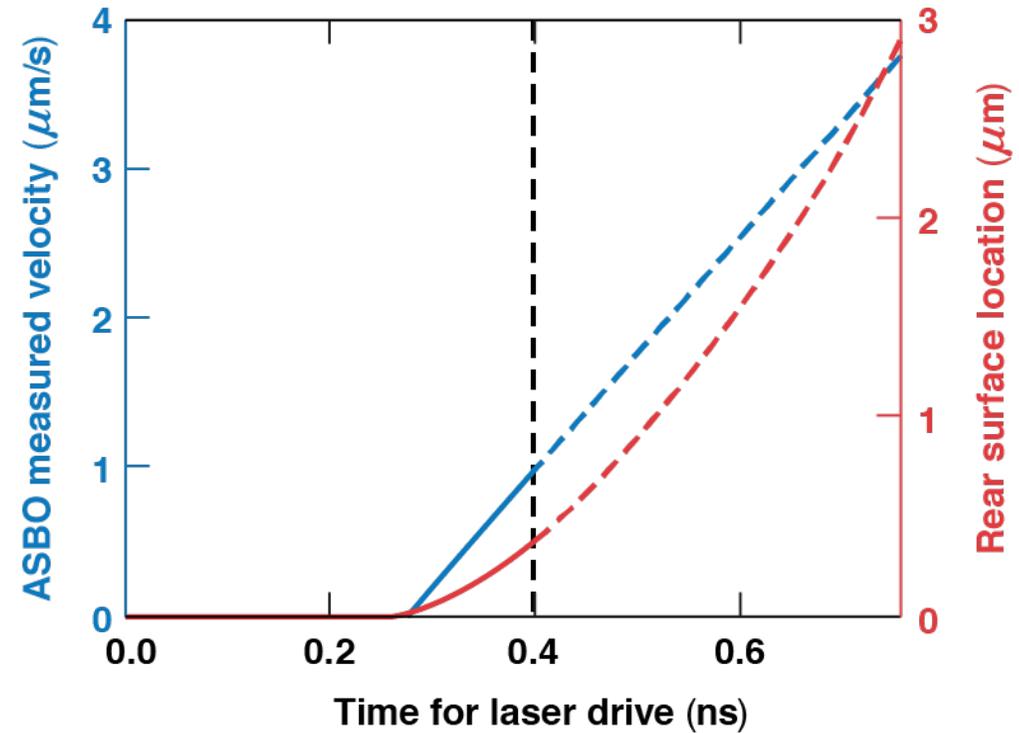
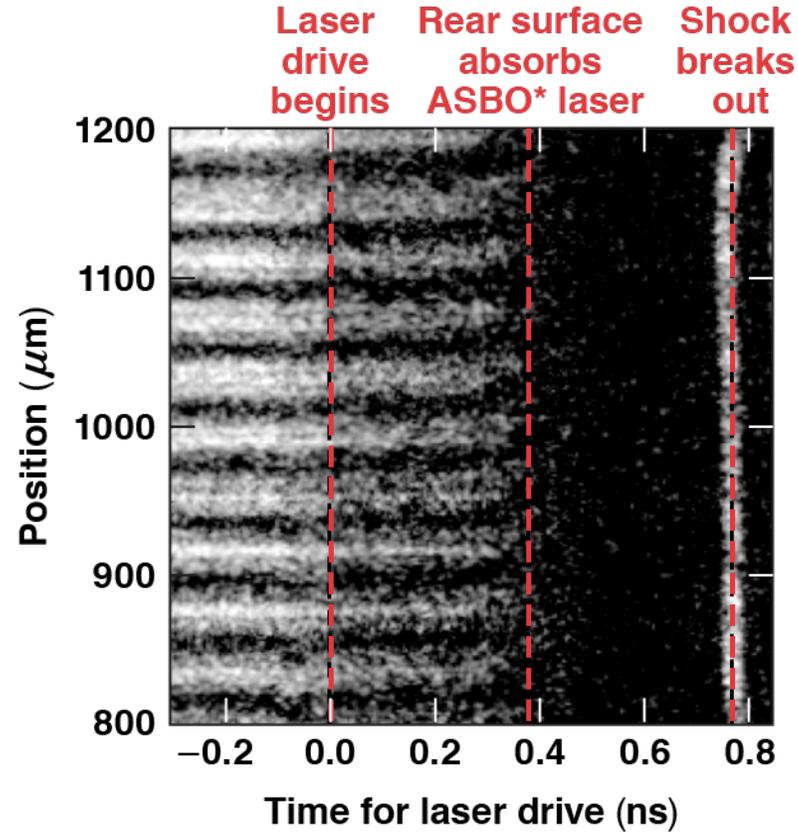
# The density profile on the back side of the shell before the shock breakout strongly affects the rarefaction wave expansion



TC15202a



# VISAR shows movement of the back side of the CH foil before the shock breakout

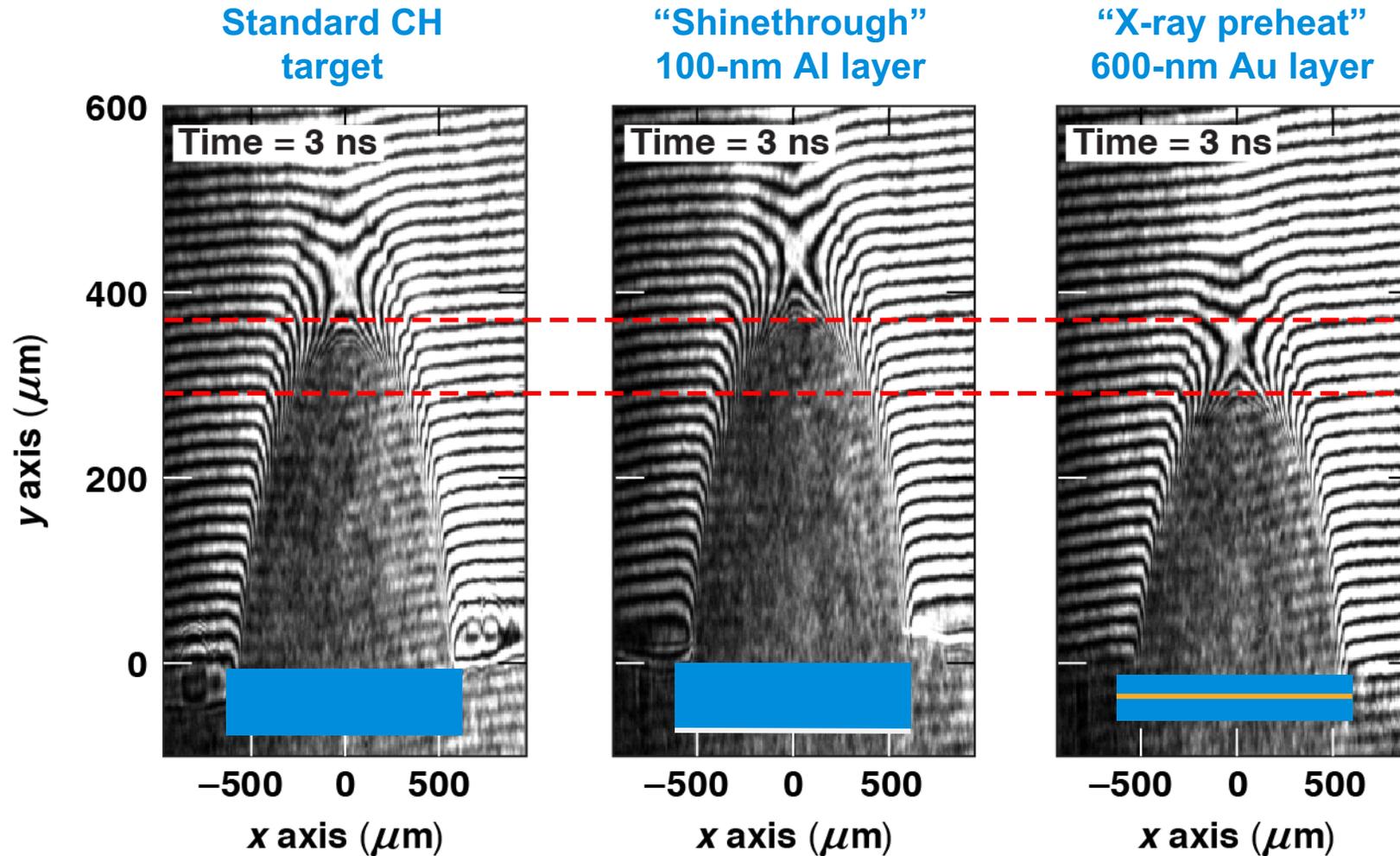


TC15601

Coronal x-ray preheat or laser shinethrough?

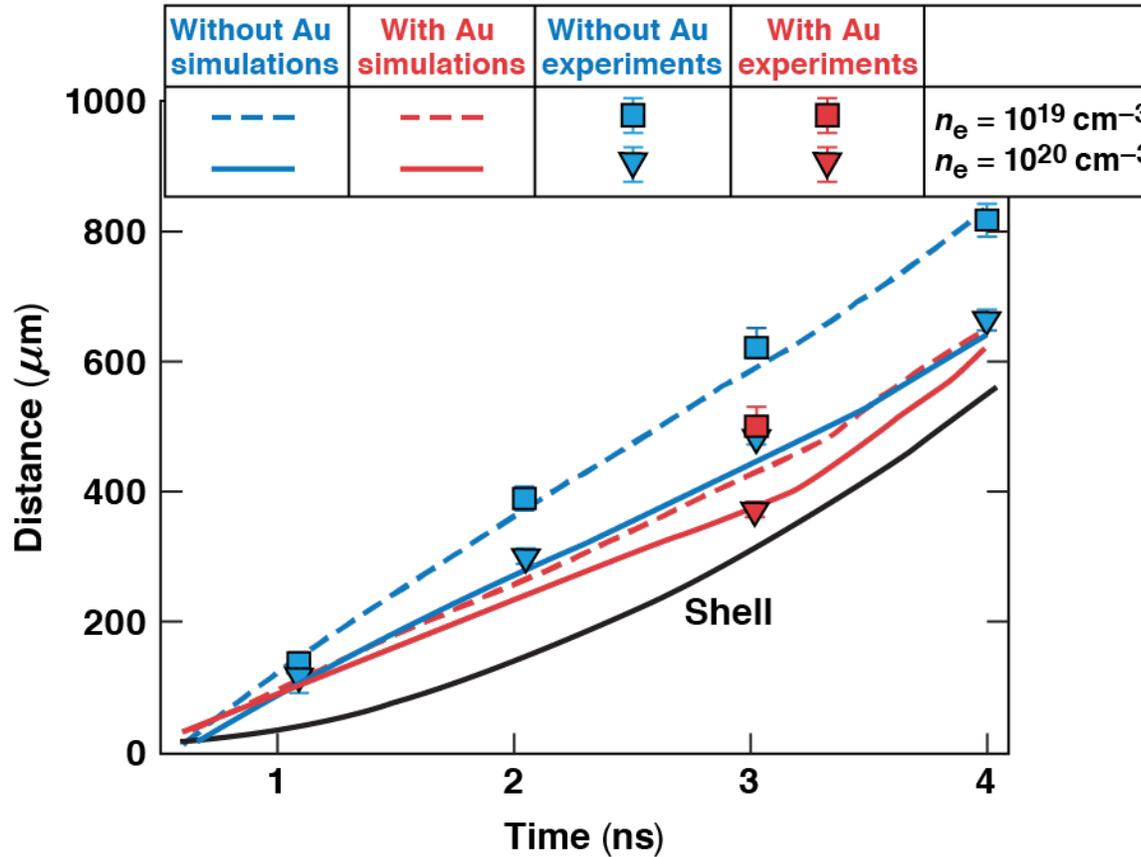
\* ASBO: active shock breakout

# The latest experiments were aimed at testing x-ray preheat as well as shinethrough as the culprit for pre-expansion of the rear side of the target

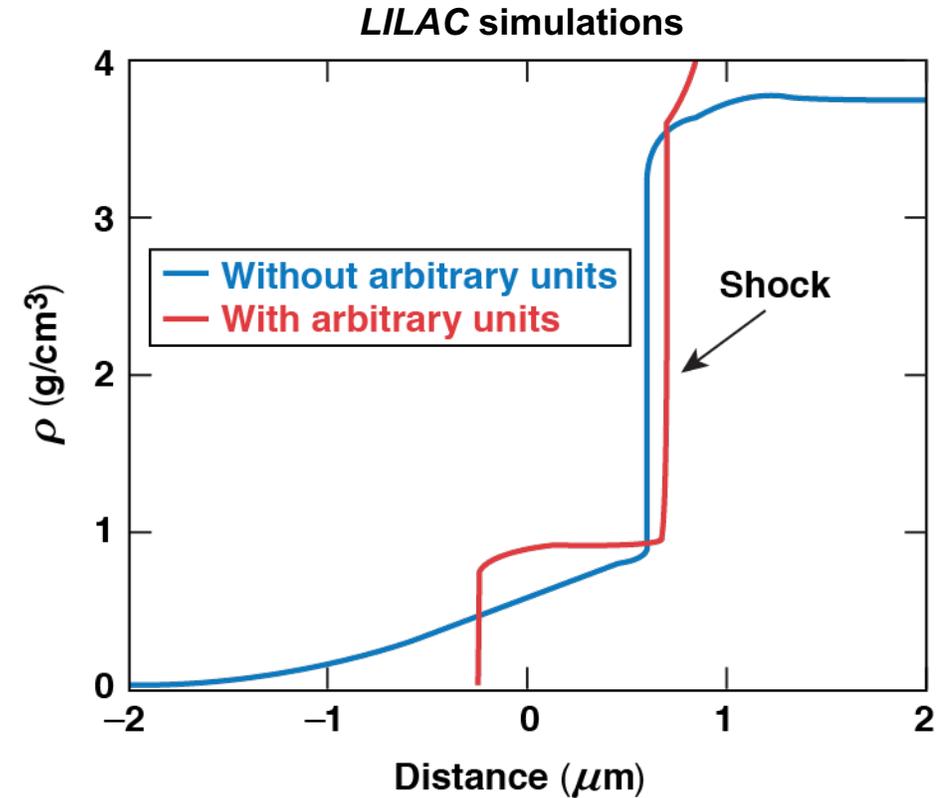


TC15602

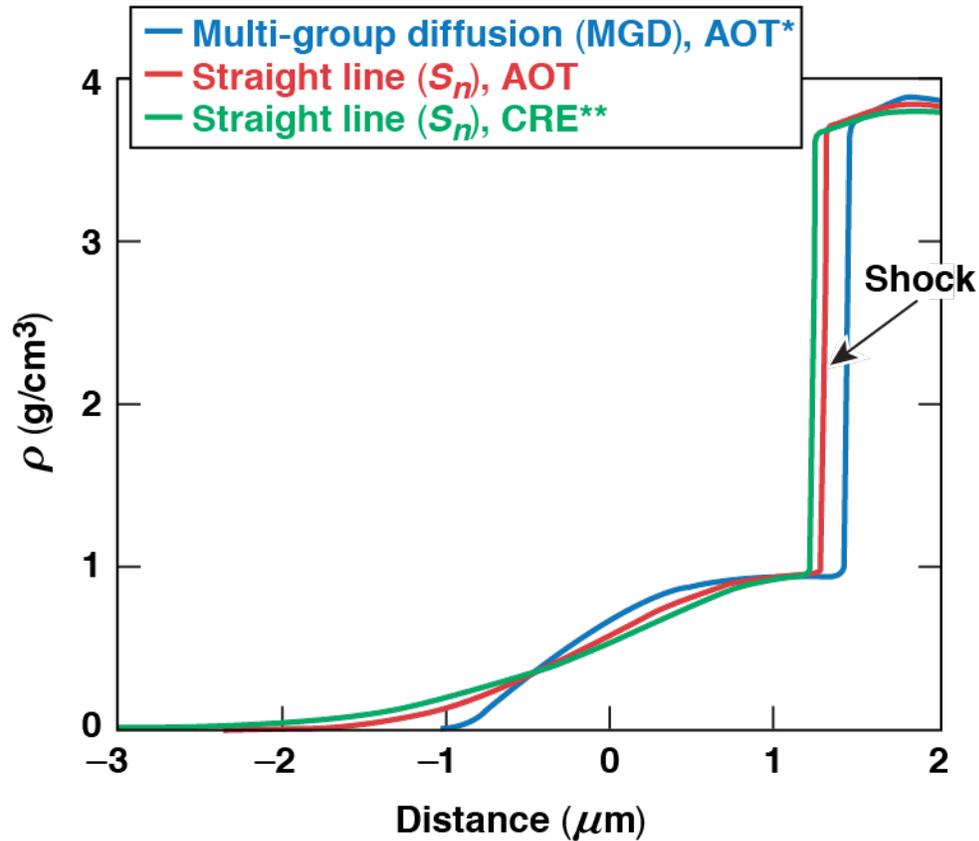
# The buried Au layer significantly affects the position of the rarefaction wave



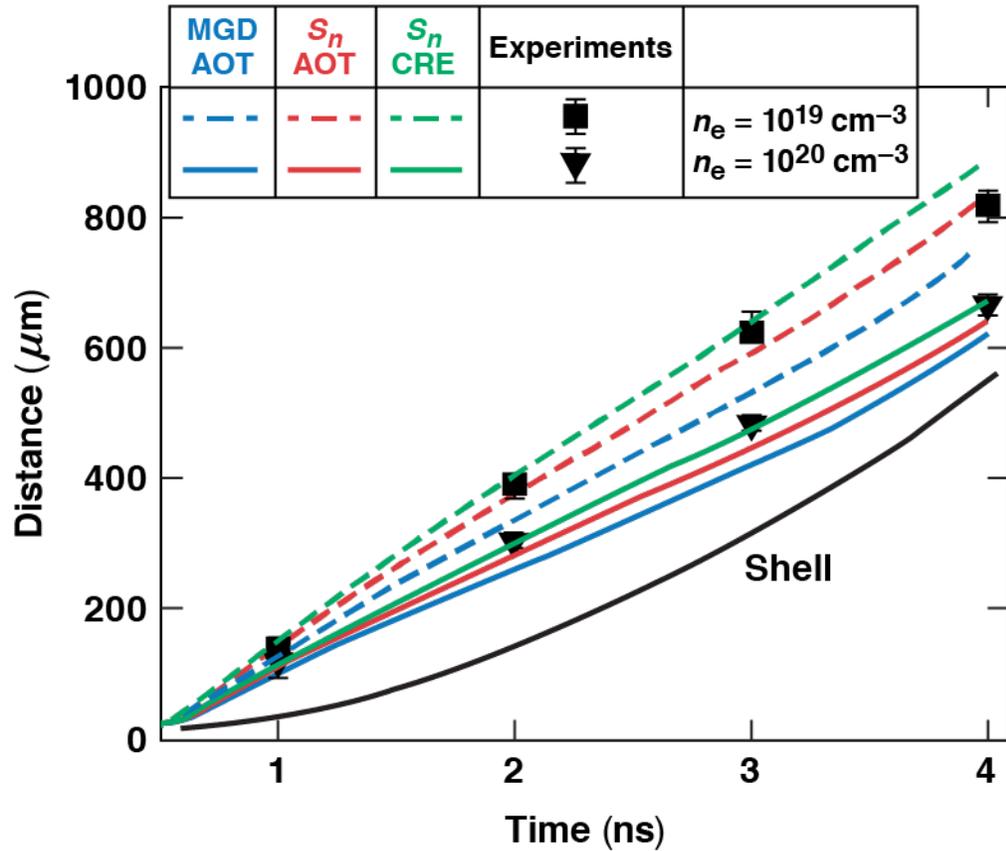
TC14644d



# The early expansion and therefore the release strongly depend on the radiation transport models used in the simulations



TC14644e



W. F. Huebner *et al.*, Los Alamos National Laboratory, Los Alamos, NM, Report LA-6760-M (1977);  
 \* AOT: astrophysical opacity tables  
 R. Epstein *et al.*, Bull. Am. Phys. Soc. **43**, 1666 (1998);  
 \*\* CRE: collisional radiative equilibrium

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