Three-Dimensional Modeling of Laser–Plasma Confinement in a Strong Magnetic field



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Extended MHD is needed to simulate plasma disc formation and magnetic confinement observed in laser-wire interaction experiments on Zebra

• The plasma disc is pinched and magnetically confined by the azimuthal magnetic field from the wire

- These simulations are in general agreement with the experiments at the Zebra facility using $\lambda = 1.06$ - μ m light and 1 MA of current
- A simulation using an extended magnetohydrodynamic (MHD) model captures the azimuthal plasma expansion and the evolution of plasma perturbations seen in experiments





Collaborators



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Studies of laser-generated magnetized plasmas have an important impact on magnetized inertial confinement fusion (ICF) and laboratory astrophysics

- Multiple magneto-inertial fusion concepts involve the interaction of magnetic fields and laser plasmas
- The coupling of laser-ablated plasmas in strong magnetic fields is also similar to magnetized astrophysical phenomena with the plasma β parameter changing from $\beta < 1$ to $\beta \gg 1$

 $\beta = \frac{nkT}{B^2/2\mu}$

TC15107

NASA/SDO/Goddard Scientific Visualization Studio D. J. Strozzi, *et al.*, No. LLNL-CONF-672979. Lawrence Livermore National Lab. (LLNL), Livermore, CA (United States), 2015 M. R. Gomez, *et al.*, Physical review letters <u>113.15</u>, 155003 (2014). J. R. Davies, *et al.*, Physics of Plasmas <u>26.6</u>, 062701 (2017).

MIFEDS: magneto-inertial fusion electrical discharge system MagLIF: magnetized liner inertial fusion NIF: National Ignition Facility LLNL: Lawrence Livermore National Laboratory

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The Zebra pulsed-power generator at UNR can be coupled with the Leopard laser for magnetized laser-plasma studies





- Zebra generator
 - load current 1 MA with LCM 1.7 MA*
 - current rise time 80 ns
 - storage energy 150 kJ
 - impedance 1.9 Ω



- Leopard laser
 - short pulse: 15 J, 0.35 ps
 - $I \sim 10^{19} \text{ W/cm}^2$, K = 106
 - long pulse: 35 J, 0.8 ns
 - $I \sim 6 \times 10^{15} \text{ W/cm}^2$

*V. V. Ivanov *et al.*, Plasma Phys. Control. Fusion <u>59</u>, 085008 (2017). LCM: liquid-crystal modulator UNR: University of Nevada, Reno



Disc-type plasma structures have been observed in recent experiments in megagauss magnetic fields



- The laser was focused to a spot of 30 μm with an intensity of ~3 \times 10 15 W/cm² for ~ 1 ns
- UV shadowgraphs show no disc plasma without the current
- laser probing and x-ray spectroscopy
 - -measured electron density $n_{\rm e} \sim 10^{19} \, {\rm cm}^{-3}$
 - -density in the rings of 7 \times 10 $^{18}\,cm^{-3}$
 - -electron temperature $T_e = (200 \text{ to } 400) \text{ eV}$





Cylindrical (r-z) simulations with azimuthal symmetry show axial collimation in the external magnetic field not seen without the magnetic field





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The magnetic-field evolution over time shows a pinching effect that leads to disc formation



- The laser pulse ablates the plasma, pushing the external field and generating Biermann battery (BB) magnetic fields $B_{BB} \sim \nabla n_e \times \nabla T_e$
- BB fields introduce asymmetry to the disc
- The magnetic field pinches down on the plasma
- As the plasma expands, the temperature drops, and magnetic fields diffuse into the disc



The extended MHD model leads to faster disc structure formation compared to the resistive MHD model



$$\vec{E} = -\frac{1}{c} \vec{U} \times \vec{B} + \vec{\eta} \times \vec{J} + \frac{\vec{J} \times \vec{B}}{cen_{e}} - \frac{\nabla \vec{P}_{e}}{en_{e}} - \frac{k}{e} \vec{\beta} \times \nabla T_{e}$$

 $\vec{E} = -\frac{1}{c} \vec{U} \times \vec{B} + \vec{\eta} \times \vec{J}$



Plasma structures are similar in simulations and experiments and measured parameters are in general agreement





6 ns after laser pulse



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Backup





The plasma β parameter transitions from larger than unity to less than unity around the disc, and the Hall parameter changes from small to large



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The field curvature of the disc and pressure gradient suggest the possibility of interchange instability



 $\kappa \times \nabla p > 0$ allowing for possible interchange instability

