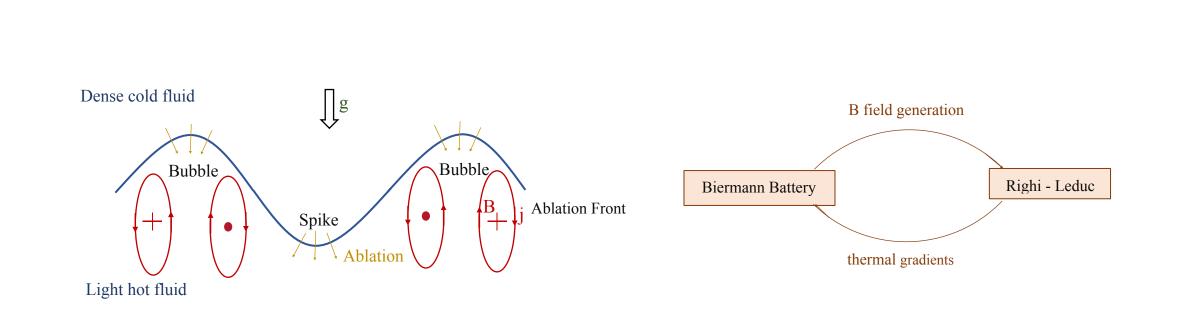
#### The Effect of Self-Generated Magnetic Fields on the Ablative Rayleigh-Taylor Instability Dynamics



F. Garcia-Rubio University of Rochester Laboratory for Laser Energetics 61st Annual Meeting of the APS Division of Plasma Physics Fort Lauderdale, Florida October 23, 2019



## The B-field is self-generated during the Rayleigh-Taylor instability, but it is weakly coupled and its effect is small in the linear regime for moderate Rem

- The self-generated magnetic field modifies the hydrodynamics by bending the heat flux lines via the Righi-Leduc term
- The Righi-Leduc term becomes important for perturbation wavelengths comparable to the distance between the ablation front and the critical surface, but computations show that the hydro B field coupling is relatively weak
- The Darrieus-Landau instability dominates over the magneto-thermal instability in the subdense region

Self-generated B field stabilizes the RTI up to a 20% in growth rate



## **Collaborators**



#### **Riccardo Betti**

University of Rochester Laboratory for Laser Energetics Department of Mechanical Engineering

Hussein Aluie University of Rochester Department of Mechanical Engineering

**Javier Sanz Recio** 

Universidad Politécnica de Madrid School of Aerospace Engineering



## Ablation fronts in Inertial Confinement Fusion are Rayleigh-Taylor unstable



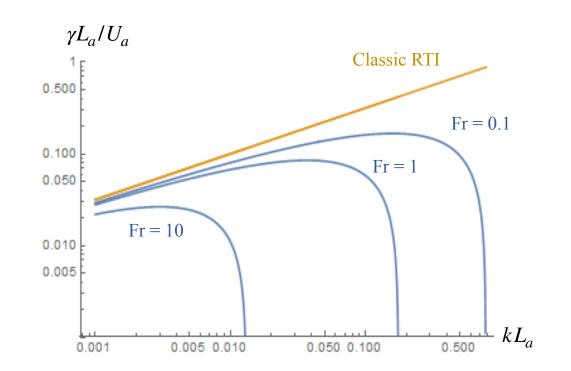


$$\gamma = 0.9\sqrt{kg} - 3kU_a$$

• Froude Number

$$Fr = \frac{Convection}{Gravity} = \frac{U_a^2}{gL_a}$$

Ablation stabilizes the RTI. The Froude number becomes the governing parameter



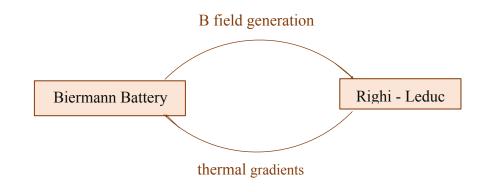


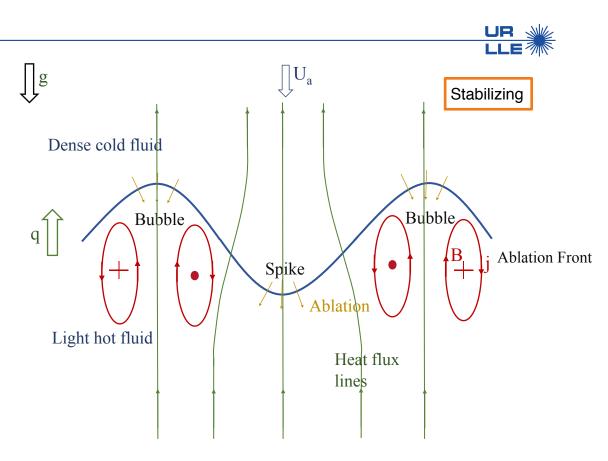
#### During the Rayleigh-Taylor instability, the B-Field is self-generated due to the Biermann-Battery term

• Biermann battery term:

$$\frac{\partial \overrightarrow{B}}{\partial t} \sim \frac{\nabla p \times \nabla n}{n^2} \propto \frac{\partial \overrightarrow{\omega}}{\partial t}$$

- The magnetic field modifies the hydrodynamics
  - Linear term: **Righi-Leduc**  $\overrightarrow{q}_{\mathsf{RL}} \sim -\frac{T^4}{n} \overrightarrow{B} \times \nabla T$
- Magnetothermal Instability





The Righi-Leduc effect bends the heat flux lines

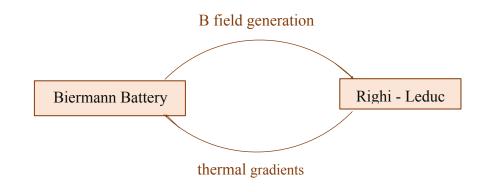


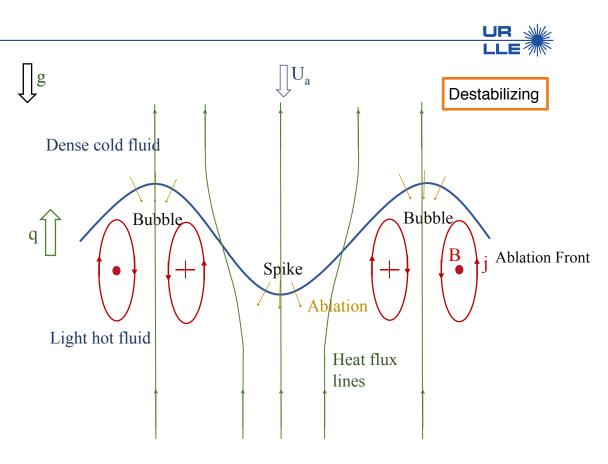
#### During the Rayleigh-Taylor instability, the B-Field is self-generated due to the Biermann-Battery term

• Biermann battery term:

$$\frac{\partial \overrightarrow{B}}{\partial t} \sim \frac{\nabla p \times \nabla n}{n^2} \propto \frac{\partial \overrightarrow{\omega}}{\partial t}$$

- The magnetic field modifies the hydrodynamics
  - Linear term: **Righi-Leduc**  $\vec{q}_{RL} \sim -\frac{T^4}{n} \vec{B} \times \nabla T$
- Magnetothermal Instability

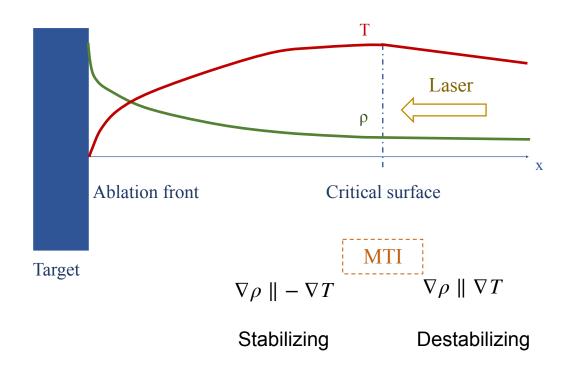




The Righi-Leduc effect bends the heat flux lines



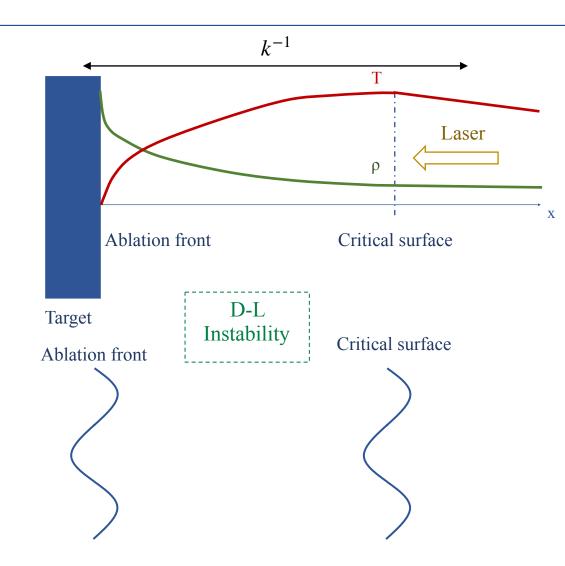
### The magneto-thermal instability takes place together with hydrodynamic instabilities



- Ablative Rayleigh-Taylor Instability (ART)
  - Sanz, Phys. Rev. Let. 73, 20 (1994)
  - Nishiguchi, Jpn. J. Appl. Phys. Vol. 41 (2002)
- Darrieus-Lanadu Instability (DL)
  - Sanz, Masse and Clavin, Phys. Plasmas 13, 102702 (2006)
- Magneto-Thermal Instability (MTI)
  - Tidman and Shanny, Phys. Fluids 17, 1207 (1974)
  - Haines Can. Journal of Physics 64(8) (1986)



### The magneto-thermal instability takes place together with hydrodynamic instabilities



- Ablative Rayleigh-Taylor Instability (ART)
  - Sanz, Phys. Rev. Let. 73, 20 (1994)
  - Nishiguchi, Jpn. J. Appl. Phys. Vol. 41 (2002)
- Darrieus-Lanadu Instability (DL)
  - Sanz, Masse and Clavin, Phys. Plasmas 13, 102702 (2006)
- Magneto-Thermal Instability (MTI)
  - Tidman and Shanny, Phys. Fluids 17, 1207 (1974)
  - Haines Can. Journal of Physics 64(8) (1986)



#### We use a magnetohydrodynamic model with Braginskii's expressions for transport terms

Induction equation

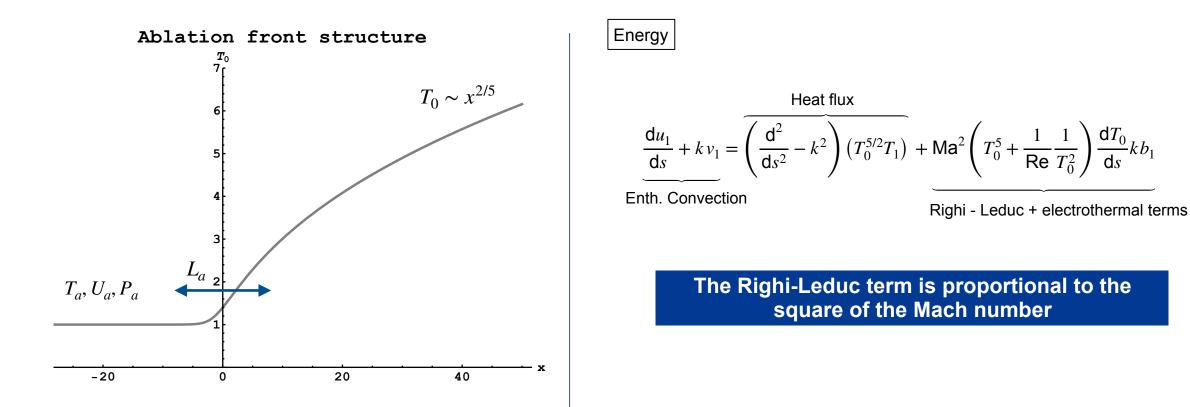
 $\frac{\partial \vec{B}}{\partial t} + \underbrace{\frac{c}{c} \nabla n \times \nabla p}_{en^2} - \underbrace{\frac{B-Convection}{\nabla \times \left(\vec{v} \times \vec{B}\right)}}_{\nabla \times \left(\vec{v} \times \vec{B}\right)} + \underbrace{\frac{B-Diffusion}{\frac{c^2 \alpha_0 m_e}{4\pi e^2} \nabla \times \left(\frac{1}{n\tau} \nabla \times \vec{B}\right)}_{ent} =$ 

$$= \underbrace{\frac{c}{4\pi e} \left(1 + \frac{\alpha_0^{''}}{\delta_0}\right) \nabla \times \left[\frac{1}{n} \overrightarrow{B} \times \left(\nabla \times \overrightarrow{B}\right)\right]}_{\text{Hall}} + \underbrace{\frac{c\beta_0^{''}}{\delta_0 m_e} \nabla \times \left(\tau \overrightarrow{B} \times \nabla T\right)}_{\text{Nernst}}.$$



#### A linear analysis of the equations has been performed to derive the stability spectrum

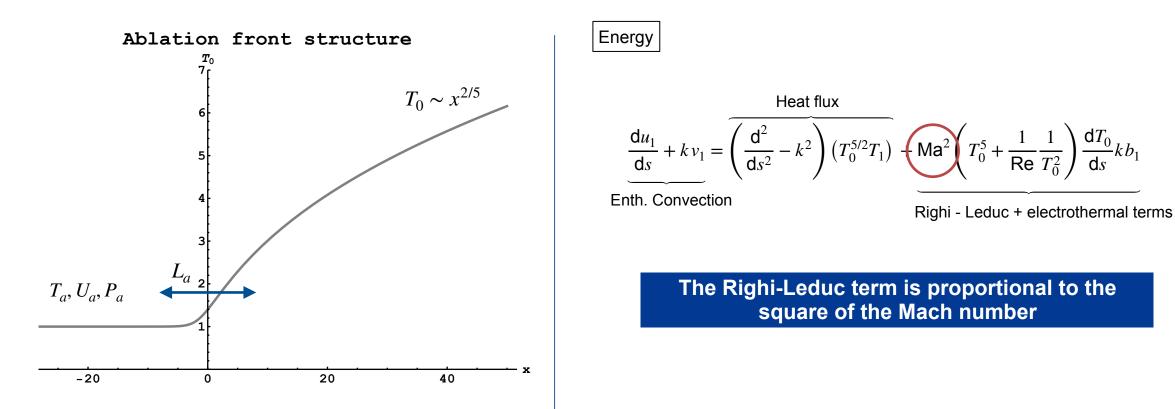
Ansatz:  $q = q_0(x) + q_1(x)\exp(\gamma t + iky)$ 



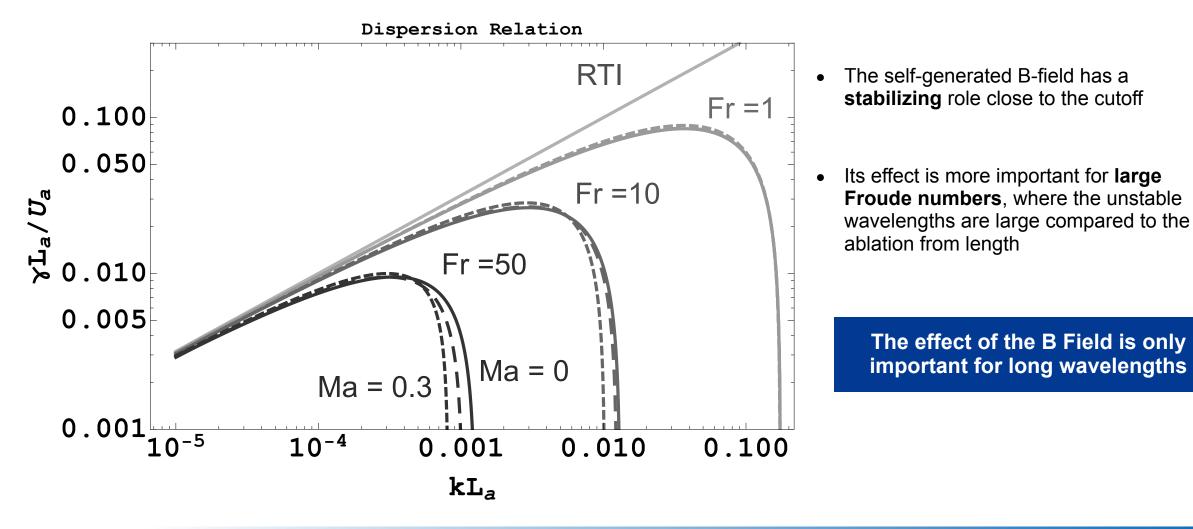


#### A linear analysis of the equations has been performed to derive the stability spectrum

Ansatz:  $q = q_0(x) + q_1(x)\exp(\gamma t + iky)$ 

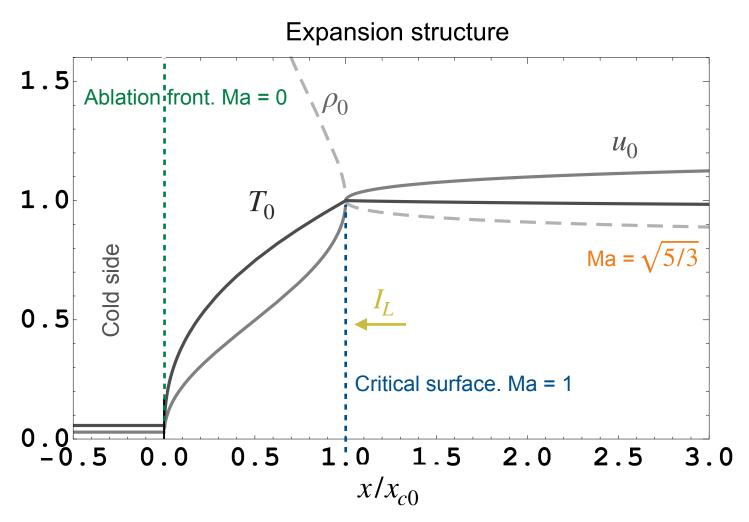


## **Dispersion relation for isobaric ablation fronts with self-generated B-fields**





#### A consistent model needs to consider non-isobaric effects and perturb the critical surface



Ablation front

 $x_a = \xi_a \exp(\gamma t + iky)$ 

Critical surface

 $x_c = x_{c0} + \xi_c \exp(\gamma t + iky)$ 

- Boundary conditions:
  - Perturbed Mach number = 1
  - Unbounded thermal mode  $\sim \exp(\lambda_T s)$
  - Unbounded magnetic mode  $\sim \exp(\lambda_B s)$



#### A consistent model needs to consider non-isobaric effects and perturb the critical surface

**Expansion structure** Perturbed mass and momentum fluxes 1.5  $\rho_0 u_1 + \rho_1 u_0 = fk \quad Ma$   $2u_1 + \rho_1 u_0^2 + p_1 = qk^{3/5}$ Mass Ablation front. Ma = 0  $\rho_0$ Momentum  $u_0$ 1.0 **Dispersion relation**  $T_0$  $Ma = \sqrt{5/3}$ Cold side  $\gamma^{2} + k(1+f)\gamma - k^{2}f - \frac{k}{Fr}(1-qFrk^{3/5}) = 0$ . 0.5 . 1 . .  $Fr \gg 1$ Critical surface. Ma = 1  $\gamma = \sqrt{\frac{k}{\mathrm{Fr}} \left(1 - q \mathrm{Fr} k^{3/5}\right)}$ 0.0<sup>E</sup> \_0 3.0 0.5 2.5 2.0 5 1:0 1.5 0.0 $x/x_{c0}$ 

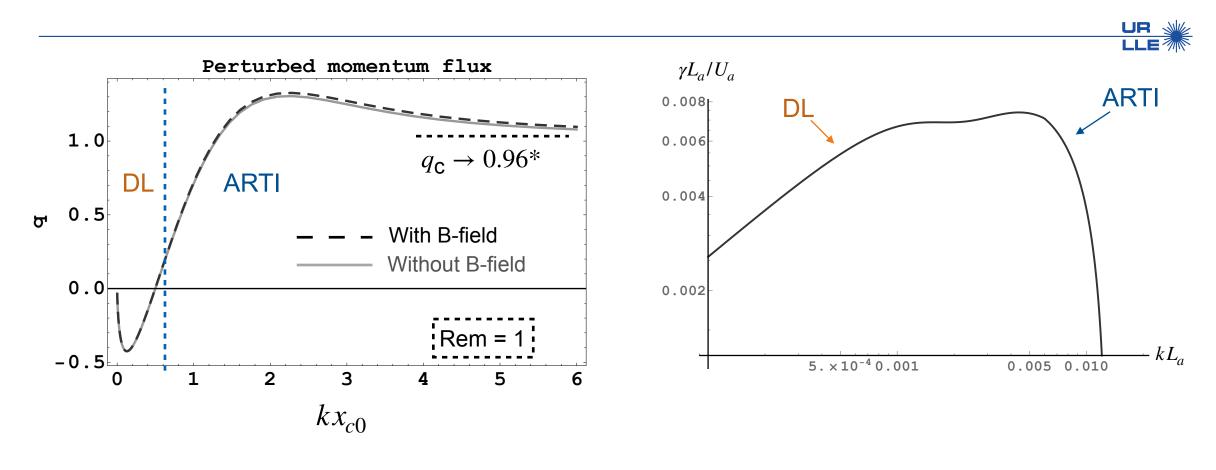


#### A consistent model needs to consider non-isobaric effects and perturb the critical surface

**Expansion structure** Perturbed mass and momentum fluxes 1.5  $\rho_0 u_1 + \rho_1 u_0 = fk \quad Ma$   $2u_1 + \rho_1 u_0^2 + p_1 = qk^{3/5}$ Mass Ablation front. Ma = 0  $\rho_0$ Momentum  $u_0$ 1.0 **Dispersion relation**  $T_0$  $Ma = \sqrt{5/3}$ Cold side  $\gamma^{2} + k(1+f)\gamma - k^{2}f - \frac{k}{Fr}(1-qFrk^{3/5}) = 0$ х. 0.5 . .  $Fr \gg 1$ Critical surface. Ma = 1  $\gamma = \sqrt{\frac{k}{\mathrm{Fr}} \left(1 - q \mathrm{Fr} k^{3/5}\right)}$ 0.0<sup>E</sup> \_0 3.0 0.5 2.5 2.0 5 1:0 1.5 0.0 $x/x_{c0}$ 



### The effect of the B field is small for Rem ~ O(1)

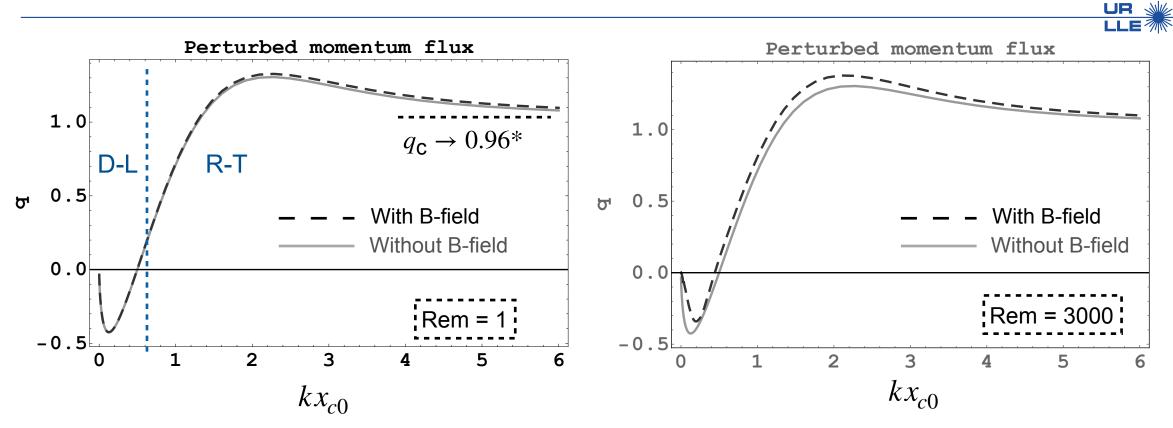


The self-generated B-field stabilizes  $\sim$  5% for low Rem

\* Sanz, Phys. Review Letters 73, 20 (1994), Sanz Phys. Review E 53, 4 (1996)



# Increasing Rem enhances the stabilizing effect of the B field, but it remains small for moderate Rem



The self-generated B-field stabilizes up to  $\sim$  20% for moderate Rem

\* Sanz, Phys. Review Letters 73, 20 (1994)



## The B-field is self-generated during the Rayleigh-Taylor instability, but it is weakly coupled and its effect is small in the linear regime for moderate Rem

- The self-generated magnetic field modifies the hydrodynamics by bending the heat flux lines via the Righi-Leduc term
- The Righi-Leduc term becomes important for perturbation wavelengths comparable to the distance between the ablation front and the critical surface, but computations show that the hydro B field coupling is relatively weak
- The Darrieus-Landau instability dominates over the magneto-thermal instability in the subdense region

Self-generated B field stabilizes the RTI up to a 20% in growth rate

