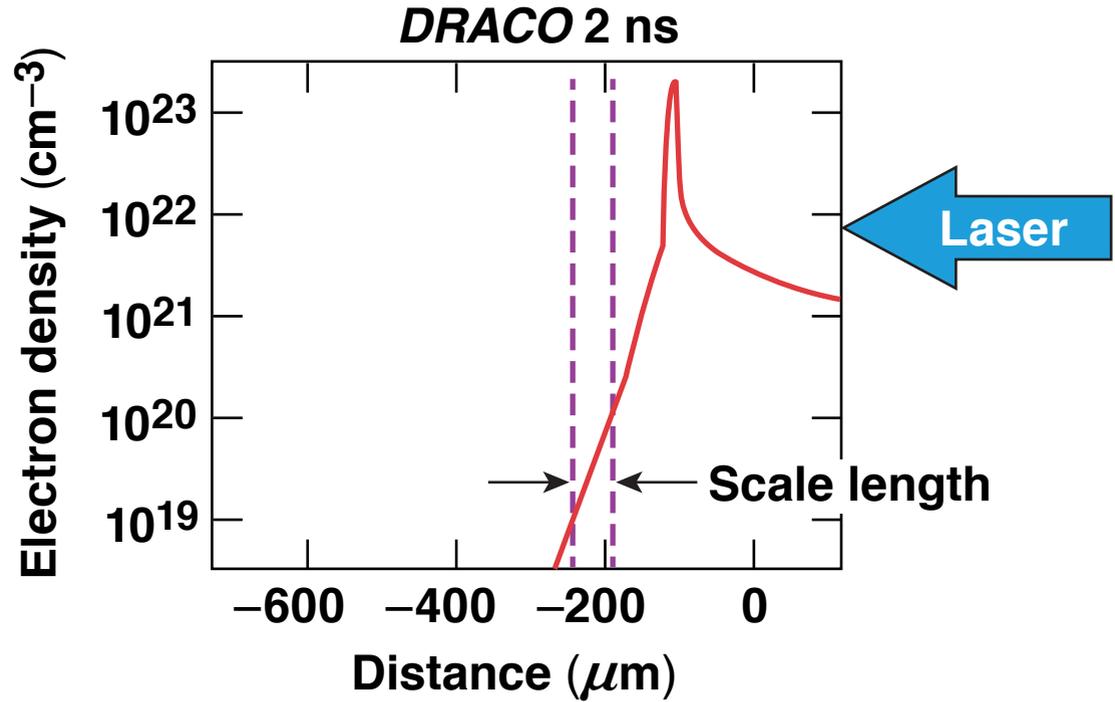
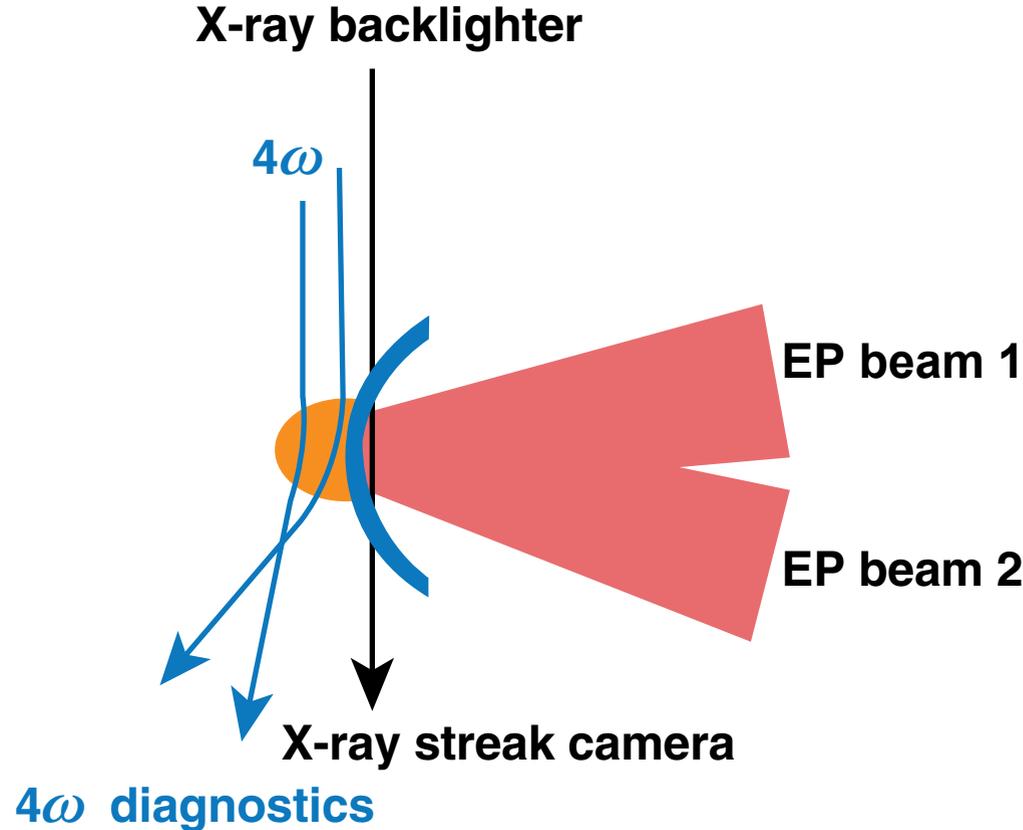


# Numerical Simulations of Shock-Release OMEGA EP Experiments



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# The measured electron density scale length in the released shock is significantly longer than that predicted by hydro simulations

- Material release from the inner shell of an implosion determines the mass of the hot spot and the onset of deceleration
- $4\omega$  probe diagnostics were used on OMEGA EP to measure the electron density profile in the released shock\*
- *DRACO*-simulated shell trajectories depend on the thermal transport models and agree with those measured using side-on x-ray radiography\*\*

# Collaborators

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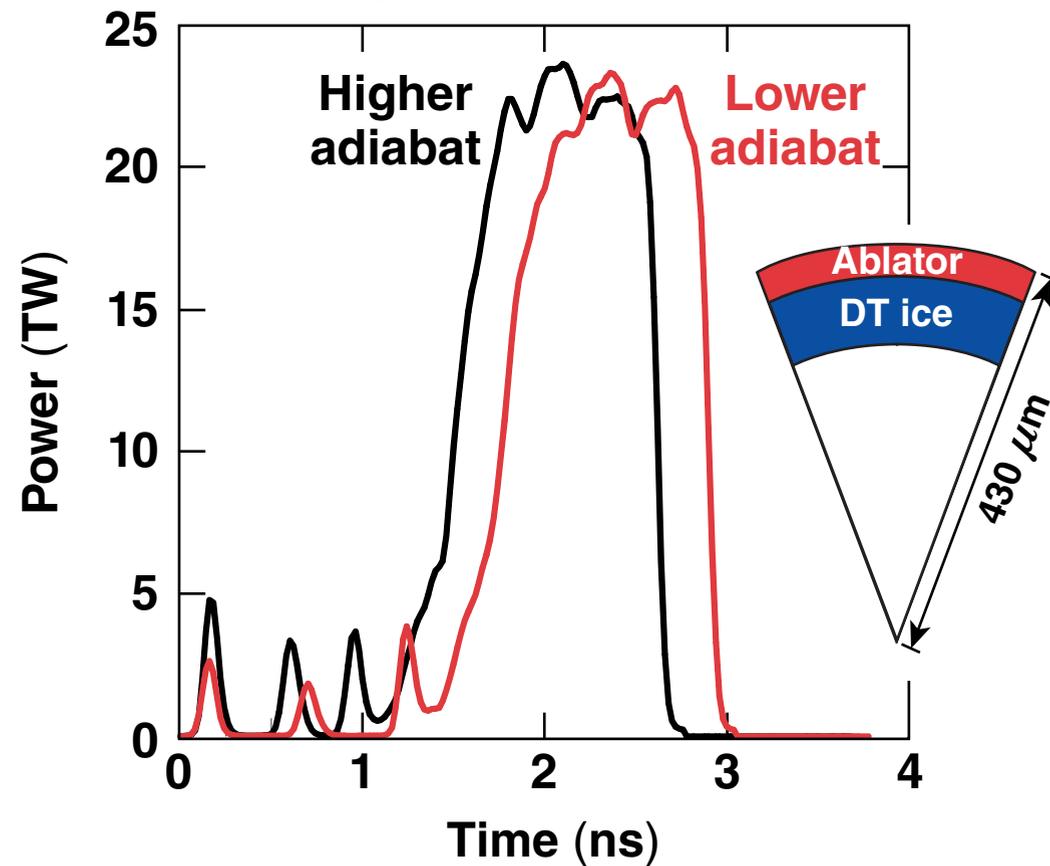


**D. Haberberger, J. Carroll-Nellenback, D. Cao, D. H. Froula,  
V. N. Goncharov, S. X. Hu, I. V. Igumenshchev, J. P. Knauer, J. A. Marozas,  
A. V. Maximov, P. B. Radha, S. P. Regan, and T. C. Sangster**

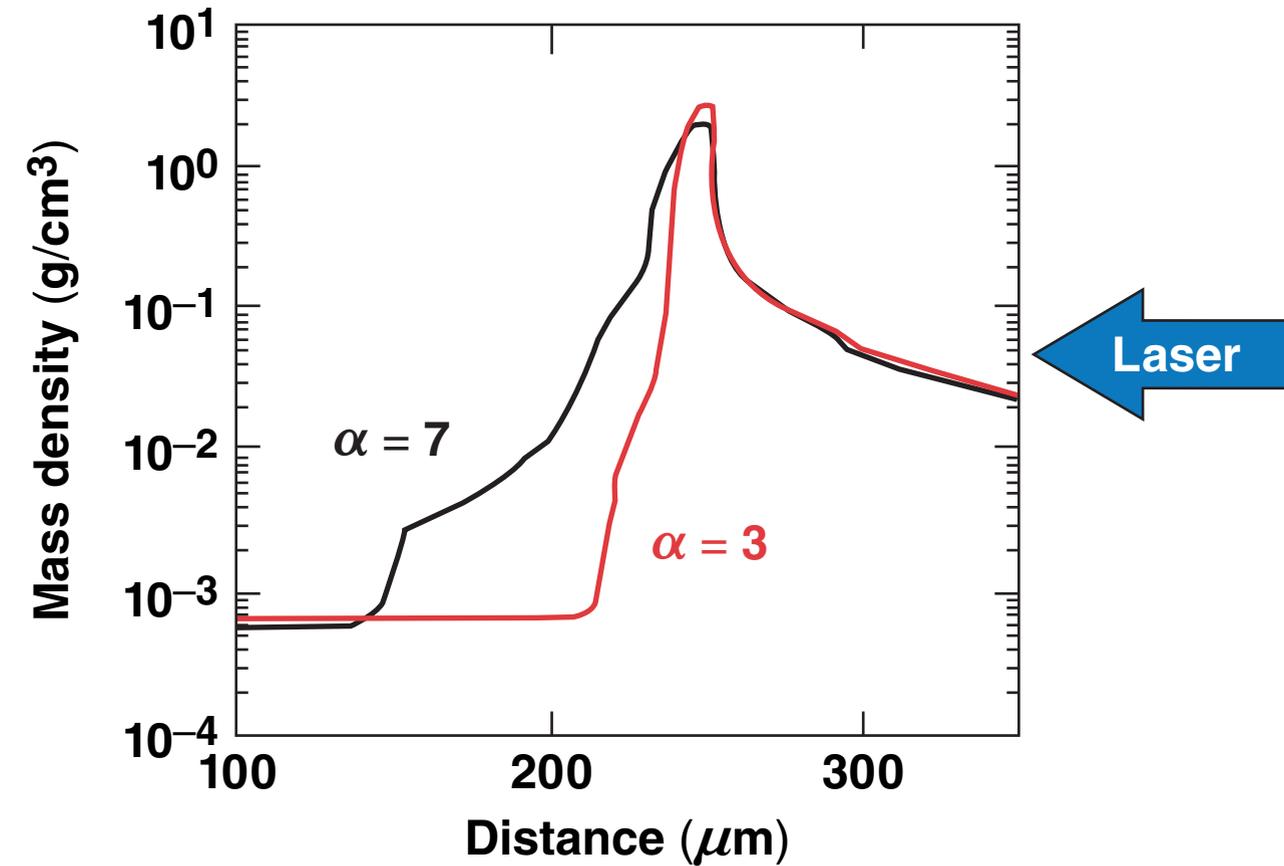
**University of Rochester  
Laboratory for Laser Energetics**

# The density profile in the rarefaction wave affects the hot-spot formation and depends on the shell adiabat and the shock-release physics

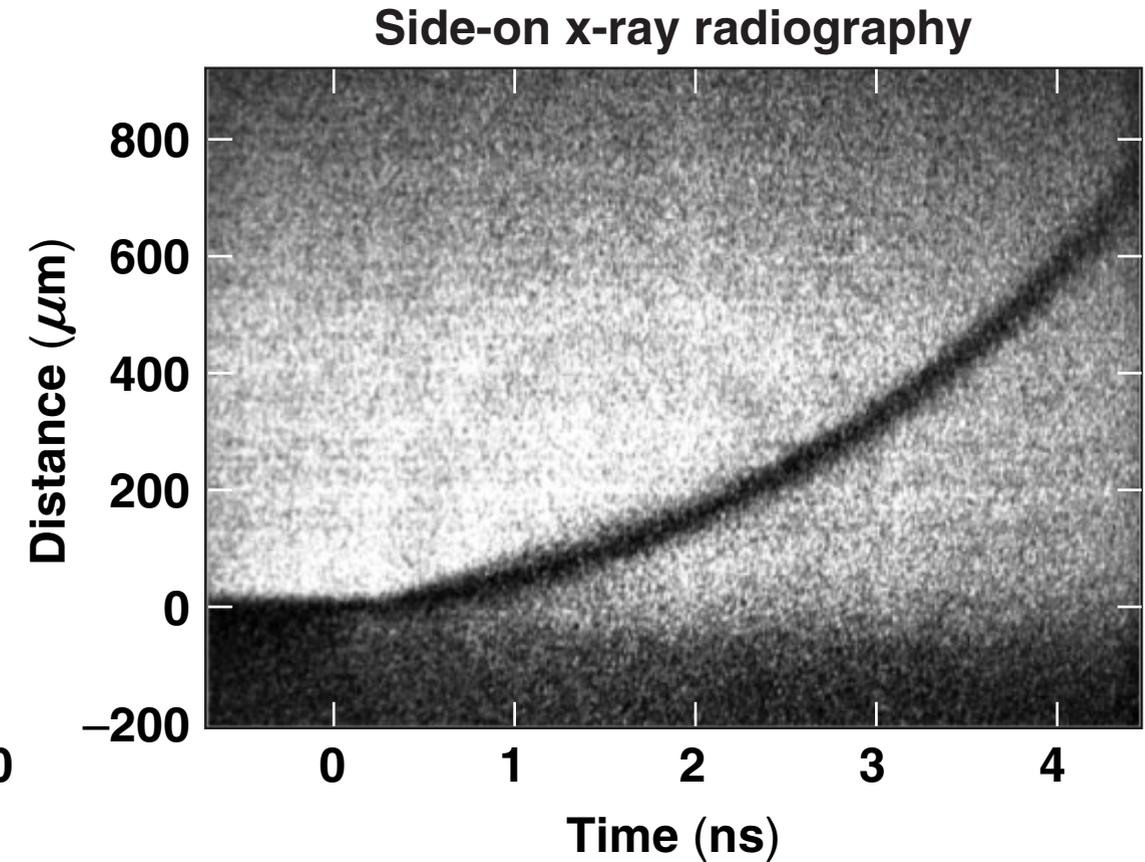
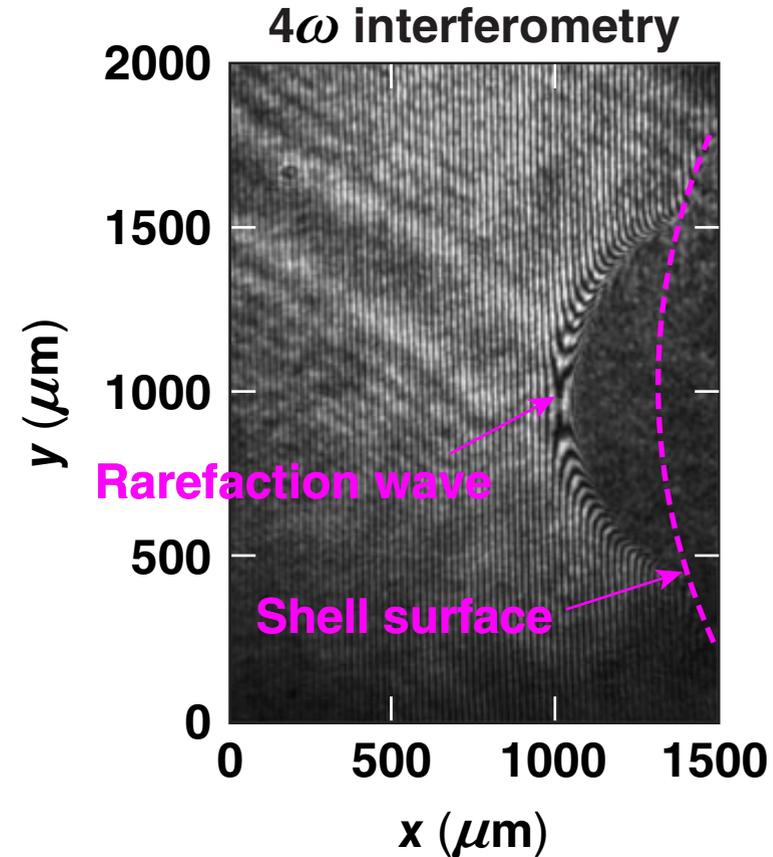
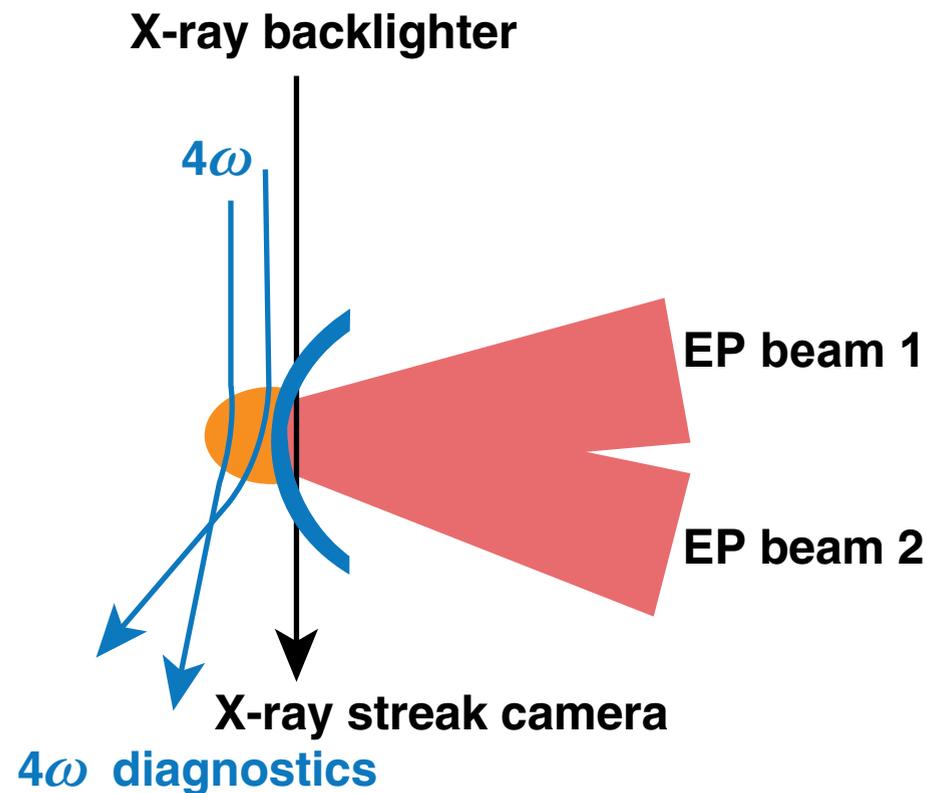
Adiabat and IFAR are controlled by pulse shaping



LILAC simulation



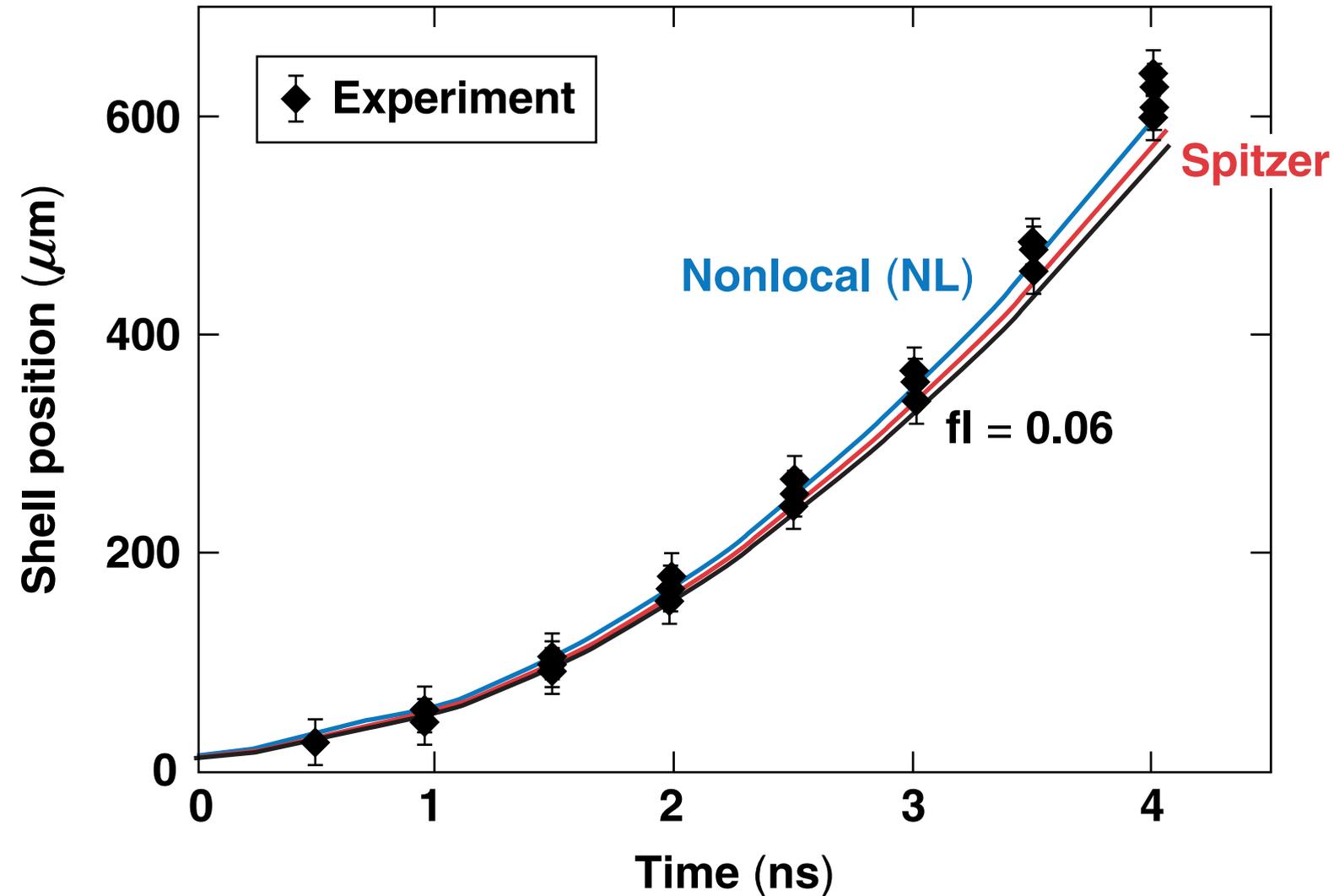
# The shock-release experiment uses a $4\omega$ probe to measure the low-density plasma profile and side-on x-ray radiography to measure the shell trajectory



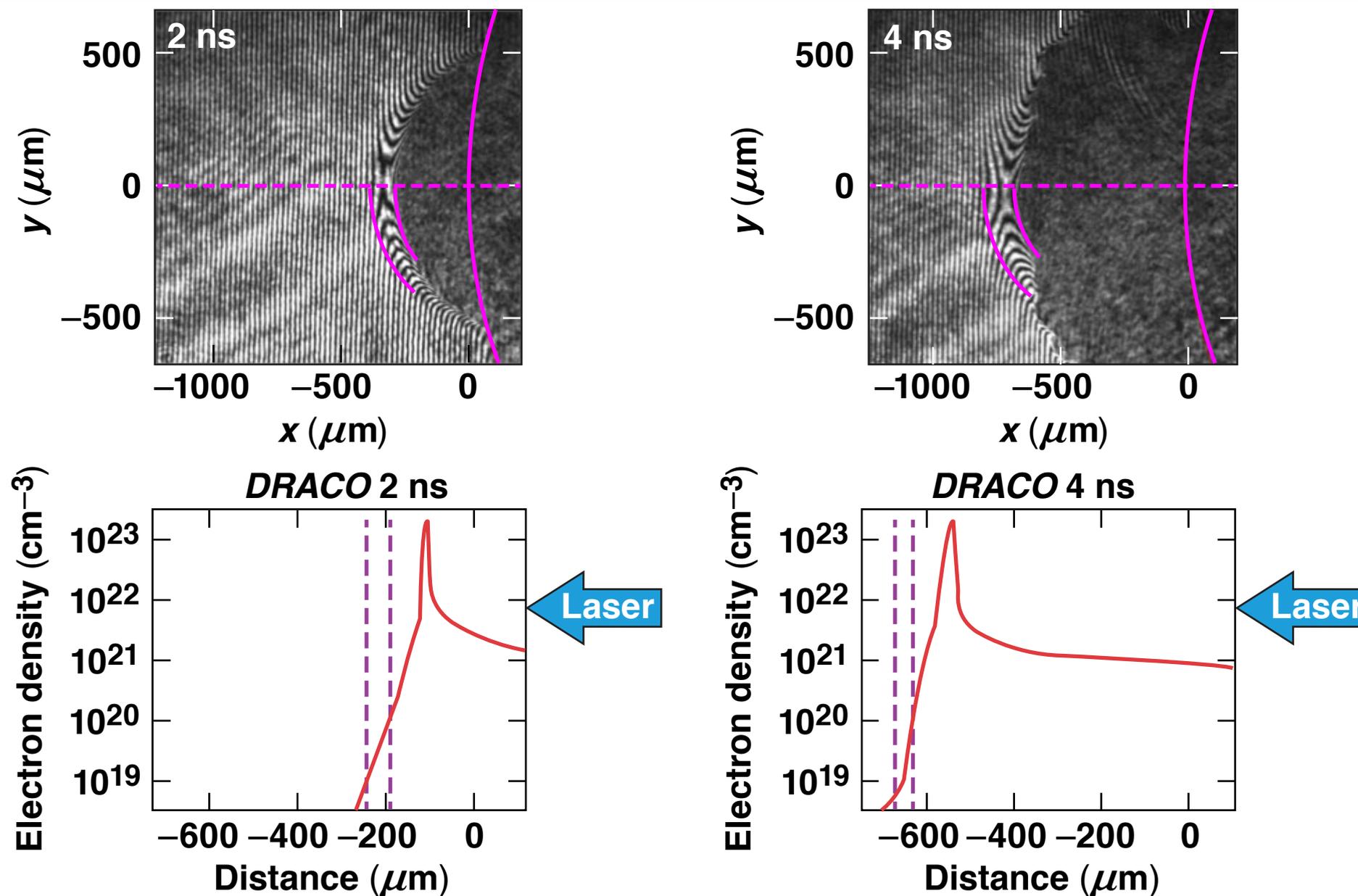
- $37 \mu\text{m}$  CH
- 4.1-mm-diam spherical cap
- 5-ns square pulse (two beams)
- $4 \times 10^{14} \text{ W/cm}^2$

TC14640

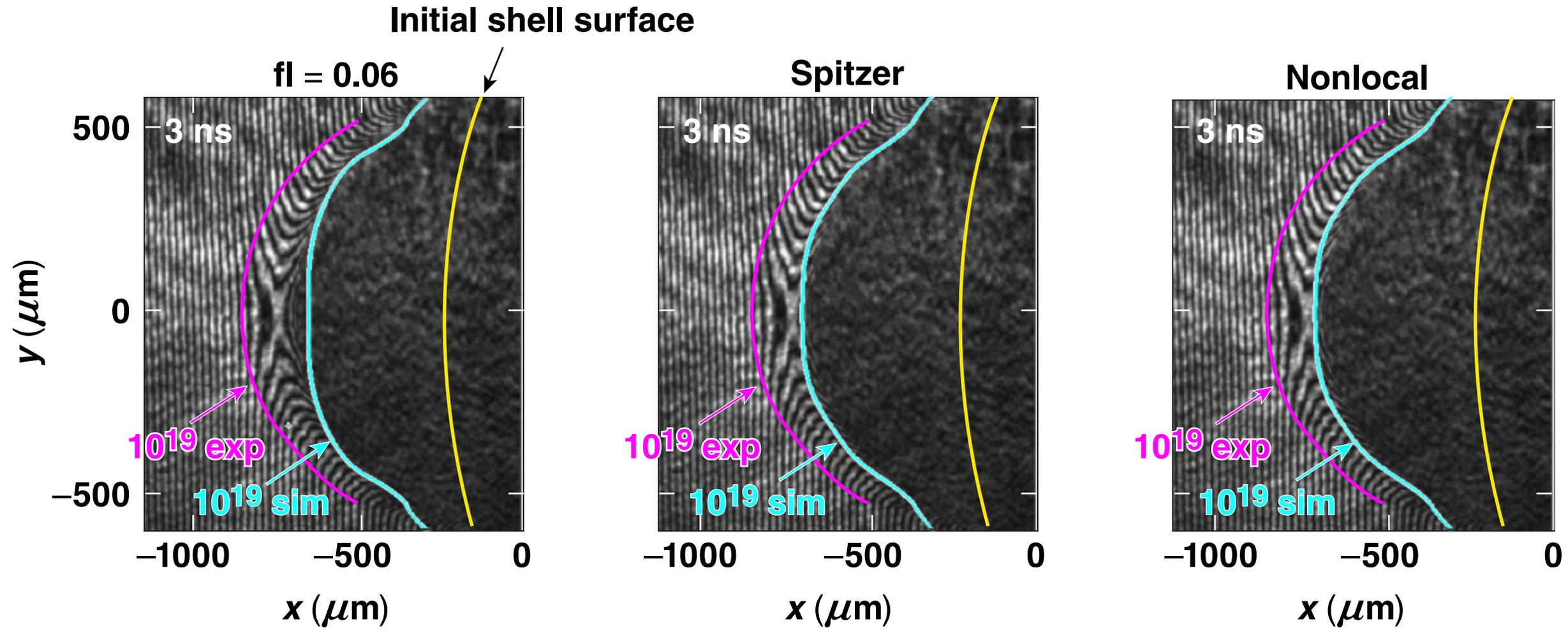
# The shell trajectory from a *DRACO* simulation using the nonlocal transport model shows good agreement with the measured shell trajectory



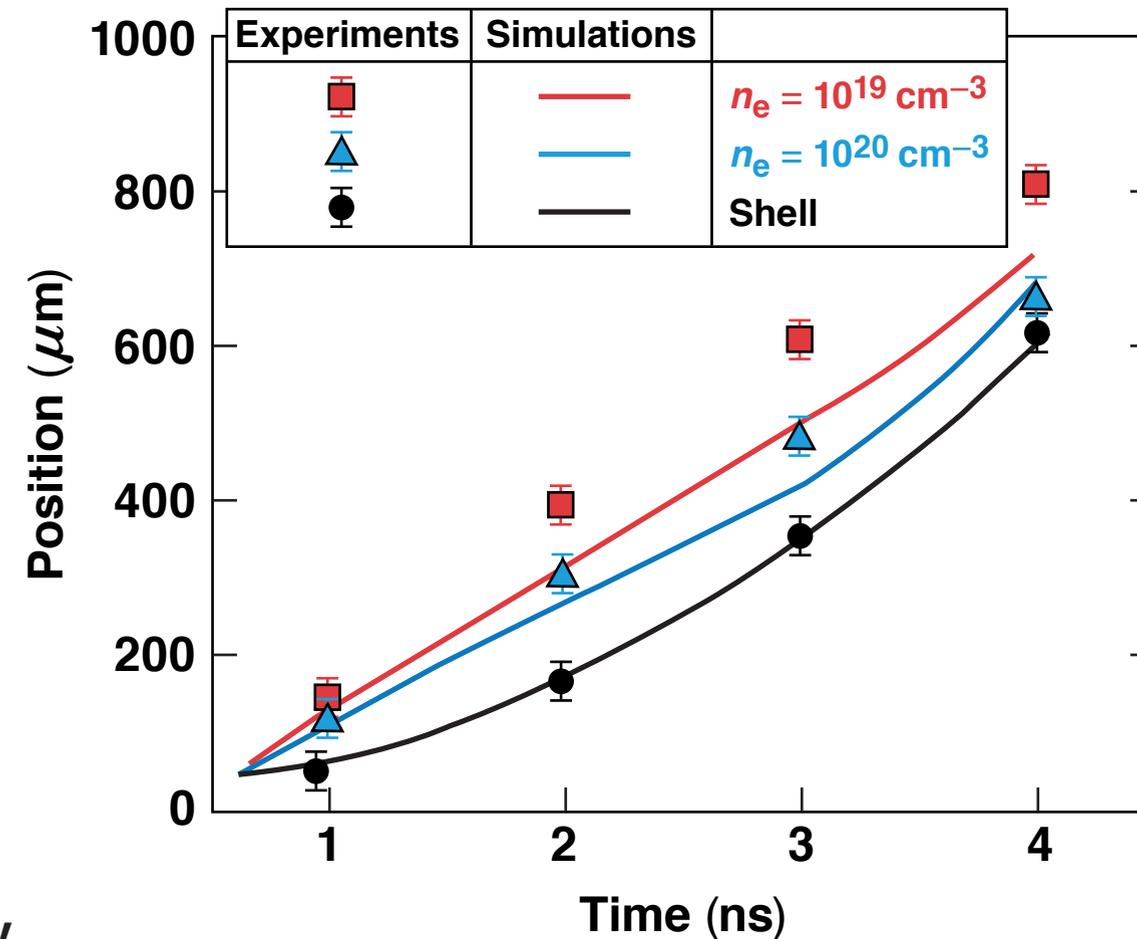
# A time series of $4\omega$ data delivers detailed temporal and spatial information about the plasma density in the rarefaction wave



# DRACO-simulated 2-D shapes of the rarefaction wave depend on the thermal transport model



# Experimentally measured scale lengths of the electron density profiles in the rarefaction wave are significantly longer than those predicted by *DRACO*



## Possible reasons for discrepancy

- Shell adiabat and sound speed are not accurately modeled
- EOS, plasma Z, and index of refraction are not accurate for the plasma conditions of the released shock
- Lack of important physics in the simulations such as ion viscosity and species separation

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