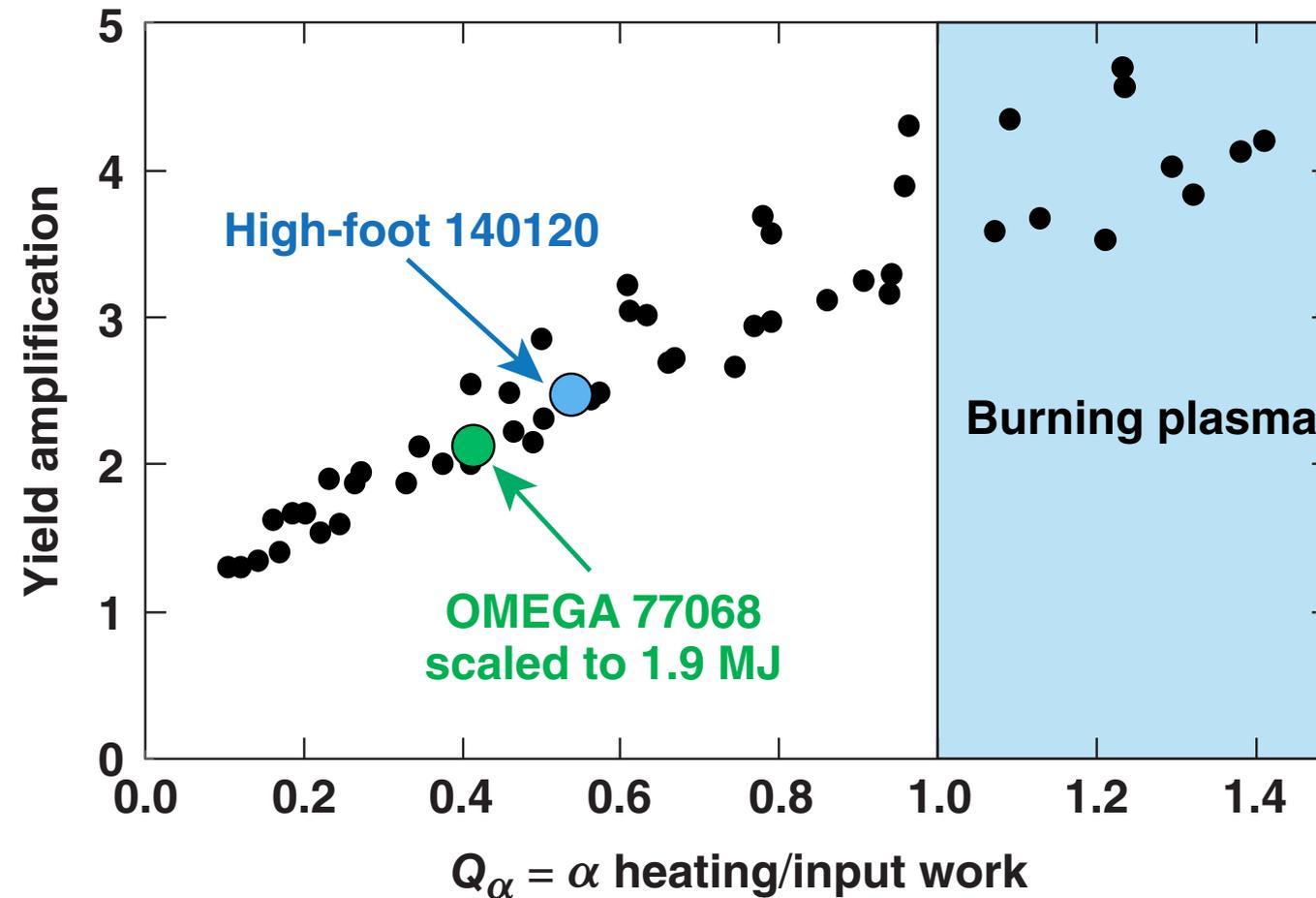


Alpha Heating and Burning Plasmas in Inertial Confinement Fusion



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Summary

The onset of the burning-plasma regime can be identified through experimental observables related to the yield amplification from α heating



- The fundamental parameter characterizing burning plasmas is $Q_{\alpha} = \alpha \text{ heating} / PdV \text{ work}$
- Current high-foot (HF)* implosions at the National Ignition Facility (NIF) have achieved $Q_{\alpha}^{\text{hs}} \approx 0.5$ to 0.6 with a yield amplification, caused by α heating, of about $2.3\times$ at 1.9-MJ laser energy
- For a high-foot-like* target, the onset of the burning-plasma regime in the hot spot ($Q_{\alpha}^{\text{hs}} = 1$) requires ~ 50 kJ of fusion energy
- Hydro-equivalent** extrapolations of direct-drive OMEGA implosions to 1.9-MJ symmetric illumination indicate performance similar to indirect drive with a yield amplification of about $2\times$ and over 100 kJ of fusion energy

*O. A. Hurricane *et al.*, *Nature* **506**, 343 (2014).

R. Nora *et al.*, *Phys. Plasmas* **21, 056316 (2014).

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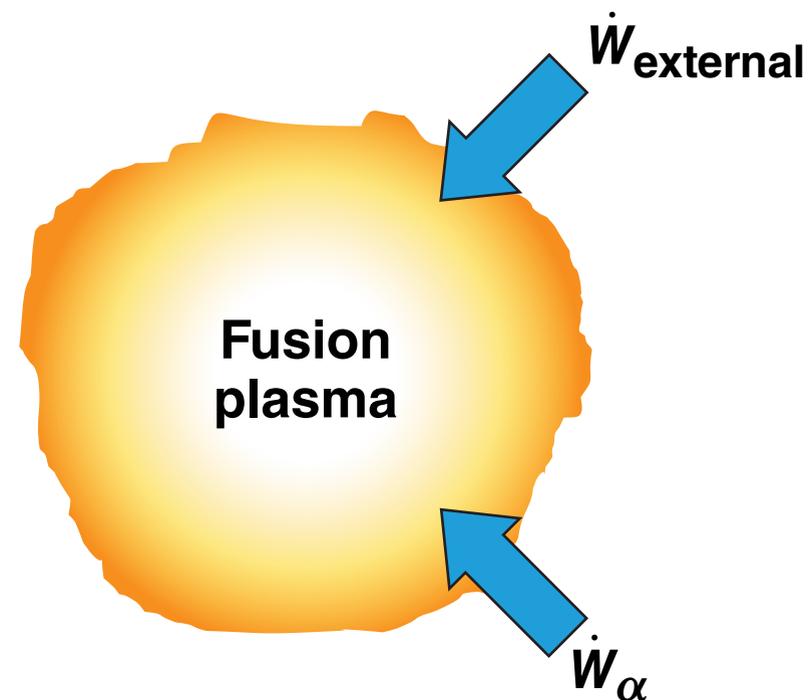
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Outline

- **Defining burning plasmas and α heating**
- **Inferring burning plasma regimes in inertial confinement fusion (ICF)**
- **α heating in indirect drive on the NIF**
- **Extrapolation of OMEGA implosions to symmetric direct drive on the NIF**

Definition of Burning Plasmas

In a burning plasma, α heating is the dominant power input to the fusion plasma



$$Q_{\alpha} \equiv \frac{\dot{W}_{\alpha}}{\dot{W}_{\text{ext}}}$$

Burning plasmas
 $Q_{\alpha} > 1$

Steady-state energy balance with power input

$$\text{External input} \longrightarrow \dot{W}_{\text{ext}} + C_{\alpha} P^2 V = \frac{3 P}{2 \tau} V \longleftarrow \text{Energy losses}$$

$\dot{W}_{\alpha} = \alpha \text{ heating for } \langle \sigma v \rangle \sim T^2$

J. P. Freidberg, *Plasma Physics and Fusion Energy* (Cambridge University Press, Cambridge, England, 2007).

The yield amplification caused by α heating depends exclusively on Q_α

Pressure with α from the power balance

$$\dot{W}_{\text{ext}} (1 + Q_\alpha) = \frac{3/2 PV}{\tau_{\text{ICF}}} \quad \tau_{\text{ICF}} \sim \frac{1}{\sqrt{P}} \quad P_\alpha \sim \left[\frac{\dot{W}_{\text{ext}}}{V} (1 + Q_\alpha) \right]^{2/3}$$

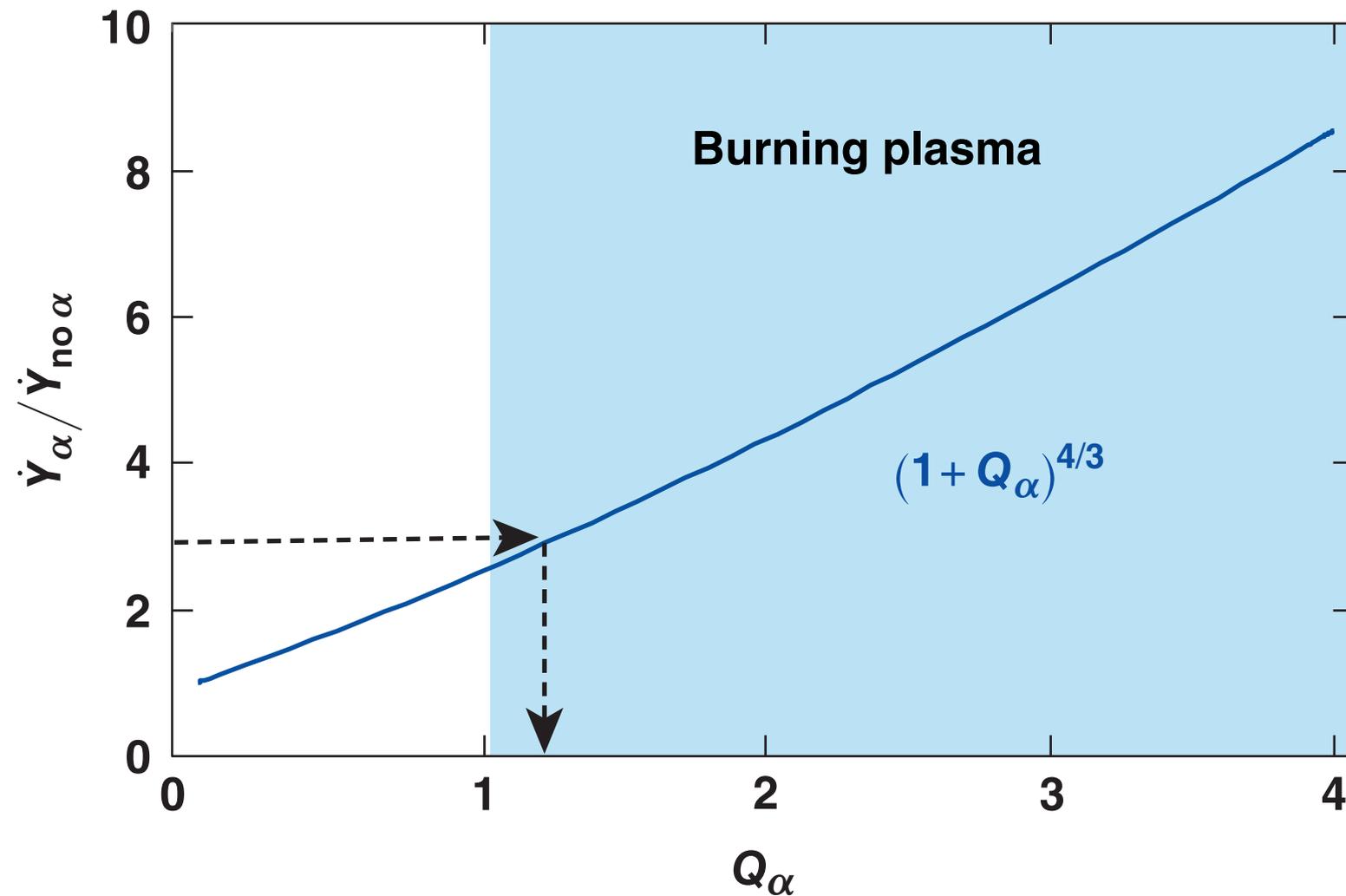
Pressure without α from the power balance

$$P_{\text{no } \alpha} \sim \left(\frac{\dot{W}_{\text{ext}}}{V} \right)^{2/3}$$

Yield amplification is a unique function of Q_α

$$\frac{\dot{Y}_\alpha}{\dot{Y}_{\text{no } \alpha}} = \left(\frac{P_\alpha}{P_{\text{no } \alpha}} \right)^2 = (1 + Q_\alpha)^{4/3}$$

If the yield amplification Y_{amp} is known, Q_{α} is determined from the $Y_{\text{amp}}-Q_{\alpha}$ plot



The definition of Q_α is modified to capture the transient character of ICF implosions

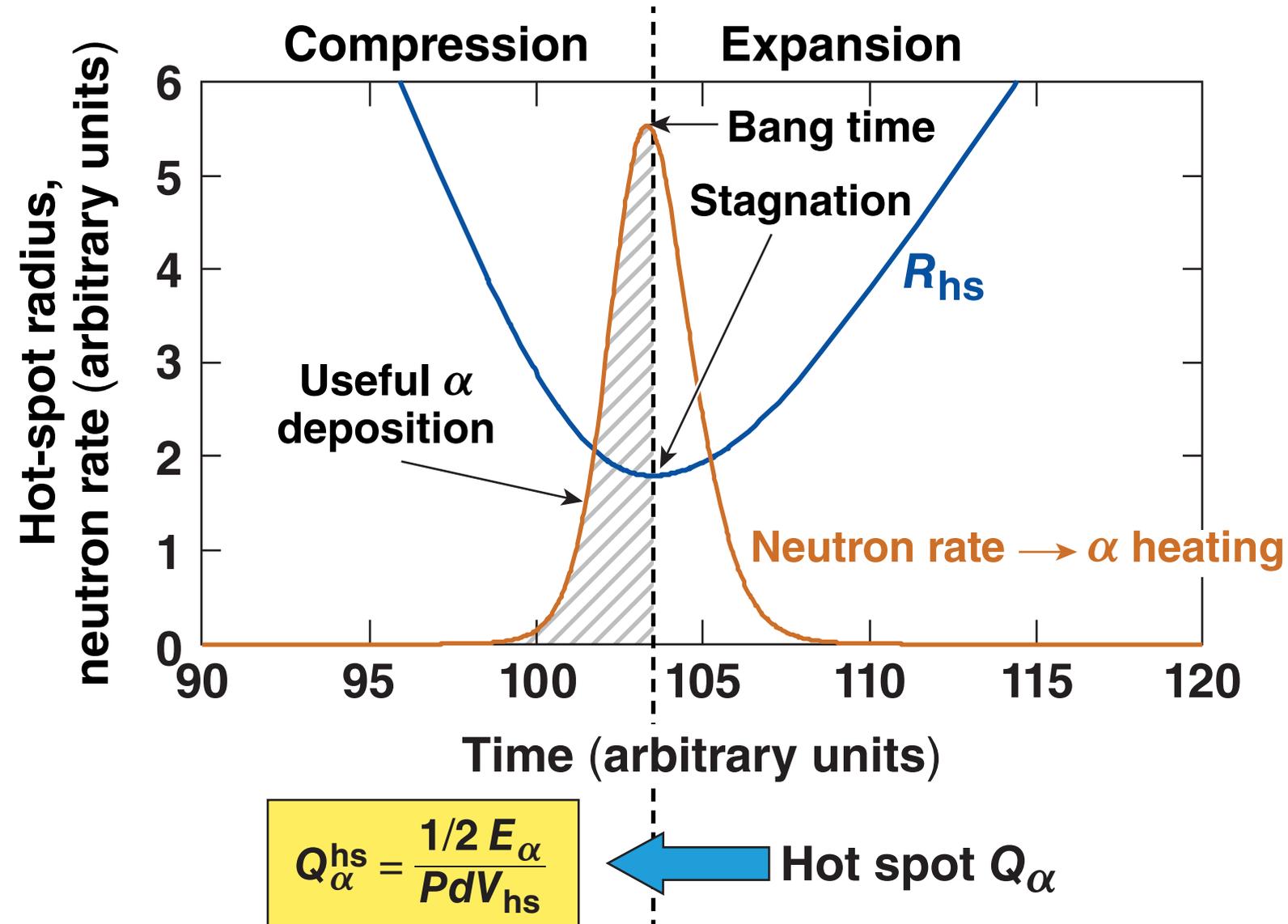
Steady state: $Q_\alpha = \frac{\dot{W}_\alpha}{\dot{W}_{\text{ext}}}$

ICF definition:

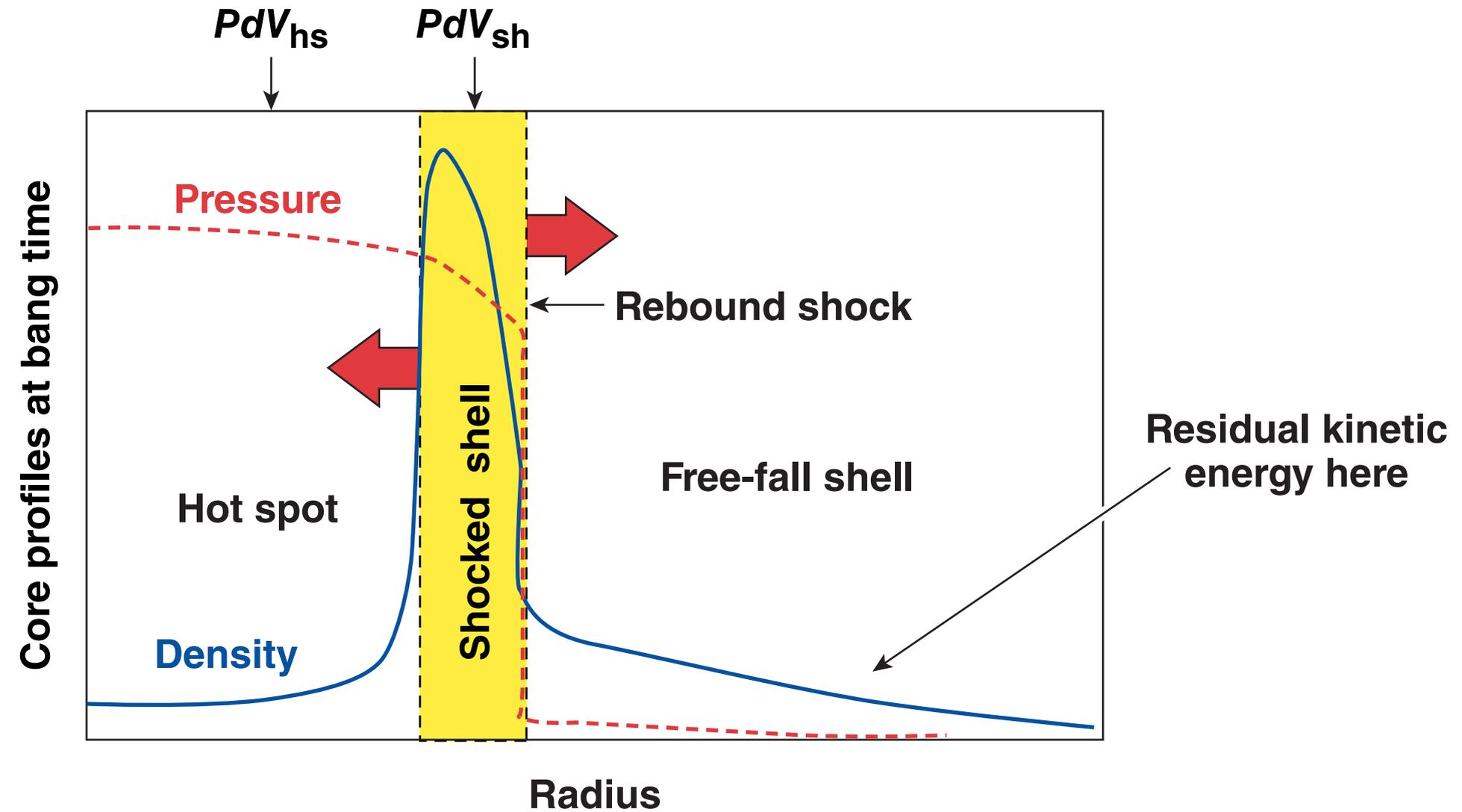
$$Q_\alpha^{\text{hs}} = \frac{\frac{E_\alpha}{2\tau_E}}{\frac{PdV_{\text{hs}}}{\tau_E}} = \frac{1}{2} \frac{E_\alpha}{PdV_{\text{hs}}}$$

- PdV_{hs} = input energy delivered to the hot spot caused by compression
- τ_E = ICF energy confinement time
- E_α = total α energy produced

Only the α energy deposition during hot-spot compression should be included when determining Q_α



In ICF the input energy is distributed between the hot spot and the shell



Two burning-plasma regimes are identified:
 α heating exceeds PdV work to the hot spot
 α heating exceeds PdV to hot spot + shell

Hot spot Q_α

$$Q_\alpha^{\text{hs}} = \frac{\frac{1}{2} E_\alpha}{|PdV_{\text{hs}}|}$$

- First burning-plasma regime

$$Q_\alpha^{\text{hs}} \geq 1$$

Total Q_α

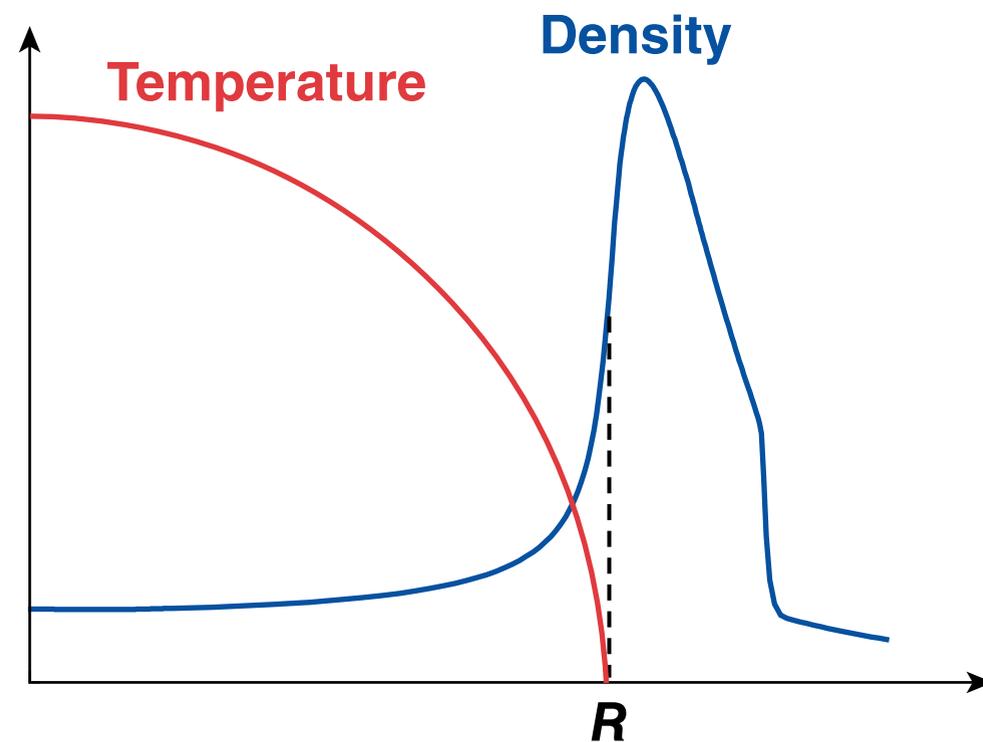
$$Q_\alpha^{\text{tot}} = \frac{\frac{1}{2} E_\alpha}{|PdV_{\text{hs}}| + |PdV_{\text{sh}}|}$$

- Second burning-plasma regime

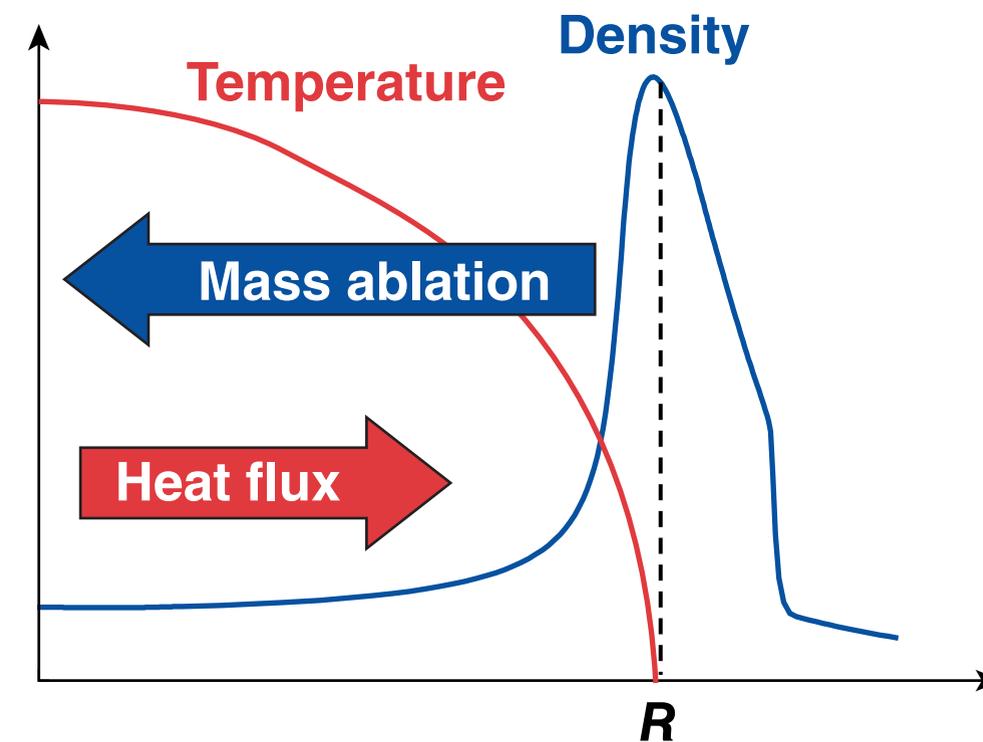
$$Q_\alpha^{\text{tot}} \geq 1$$

In simulations the PdV work is calculated by considering a fixed mass enclosed by the hot-spot volume at peak neutron production

- Hot-spot mass defined at peak neutron-production time



- Hot-spot mass traced back in time before peak neutron production



$$M_{\text{hs}}(t) = \int_0^{R(t)} 4\pi r^2 \rho dr = \text{constant} = \text{hot-spot mass at bang time}$$

The PdV work delivered to the hot spot can be exactly calculated in 1-D

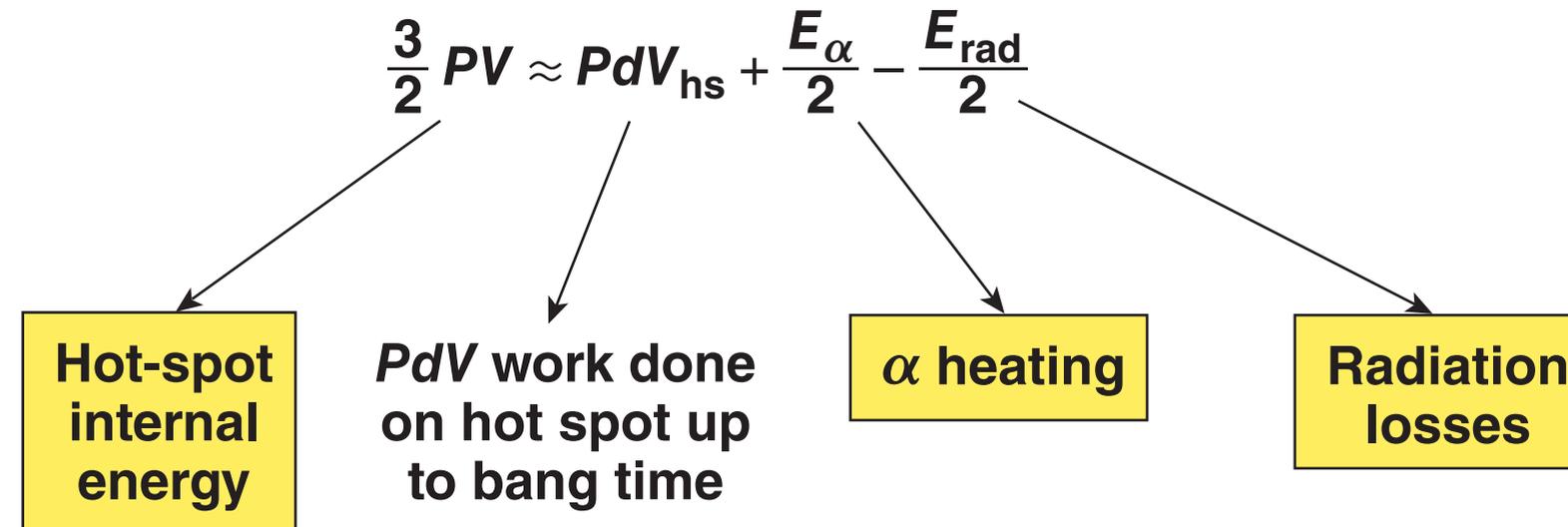
$$\text{Hot-spot } PdV \text{ work up to stagnation: } PdV_{\text{hs}} = \int_0^{t_{\text{stag}}} 4\pi PR^2 \dot{R} dt$$

Valid in the presence of ablation driven by heat conduction, α heating, and radiation for a Lagrangian mass

$$M_{\text{hs}}(t) = \int_0^{R(t)} 4\pi r^2 \rho dr = \text{constant} = \text{hot-spot mass at bang time}$$

In experiments and multidimensional simulations, the bang-time PdV work can be estimated from the hot-spot energy balance

Hot-spot energy balance at bang time



Experimental observables are used to infer hot-spot energy, α energy, and radiation losses

Hot-spot energy from burnwidth, neutron yield, T_i , and self-emission

$$Y_N \approx \frac{P^2}{16} \frac{\langle \sigma V \rangle}{T^2} V_{hs} \tau \quad \rightarrow \quad \frac{3}{2} PV \approx \sqrt{\frac{36 Y_N T^2 V_{hs}}{\langle \sigma V \rangle \tau}}$$

τ : from neutron or x-ray burnwidth
 T : from neutron time of flight
 V_{hs} : from size of x-ray self-emission image

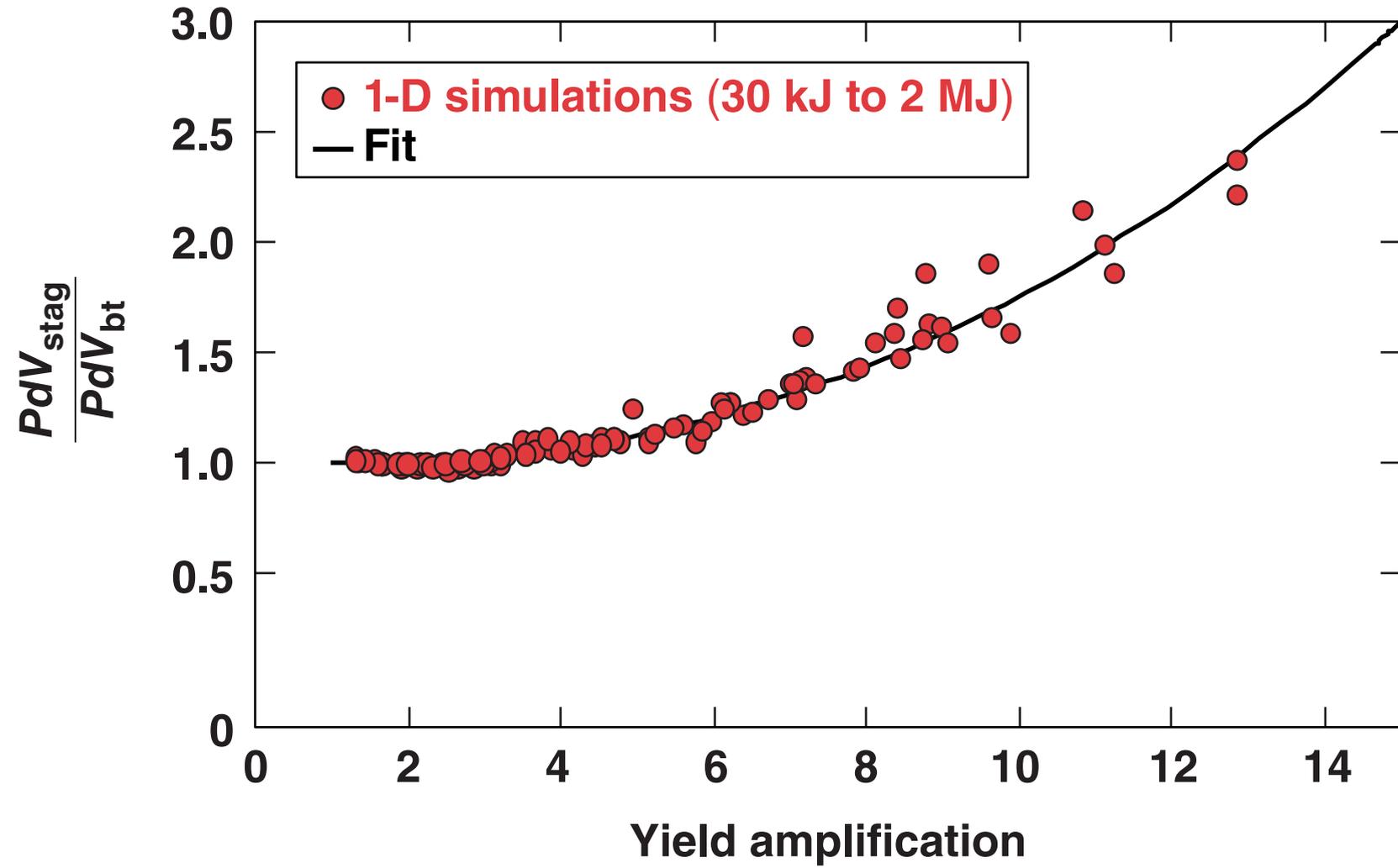
α energy from neutron yield

$$E_\alpha = (3.5 \text{ MeV}) Y_N$$

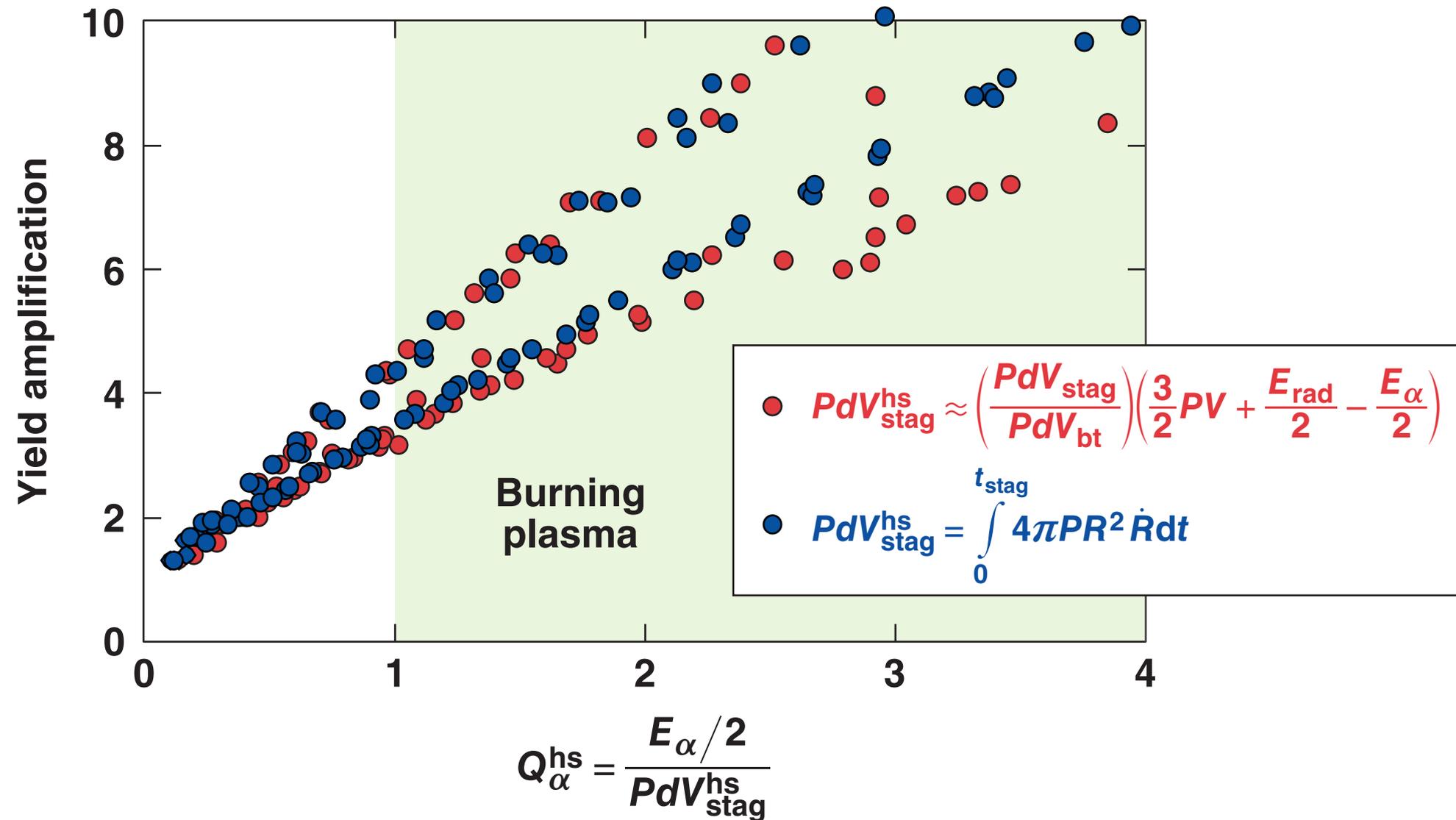
Radiation losses from α energy and temperature

$$E_{\text{rad}} \approx E_\alpha \cdot \left(\frac{3.8 \text{ keV}}{T_i} \right)^{2.8}$$

The bang-time PdV is less than the stagnation PdV because of the rapid expansion of the hot spot at large-yield amplifications



The yield amplification from simulations (30-kJ to 2-MJ energy) using both ways of calculating PdV is strongly correlated with Q_α



The physics of the hot-spot formation and shell deceleration depends on three dimensionless parameters

$$\chi_{\alpha} \equiv \frac{P\tau}{24 T^2 / \epsilon_{\alpha} \langle \sigma v \rangle} = \frac{E_{\alpha}}{E_{hs}}$$

← Lawson parameter (α heating)

$$\beta_{rad} \equiv \frac{E_{rad}}{E_{hs}}$$

← Radiation losses

$$A_{sh}^{dec} = \frac{R_{hs}(t_{dec})}{\Delta(t_{dec})}$$

← Aspect ratio at beginning of deceleration

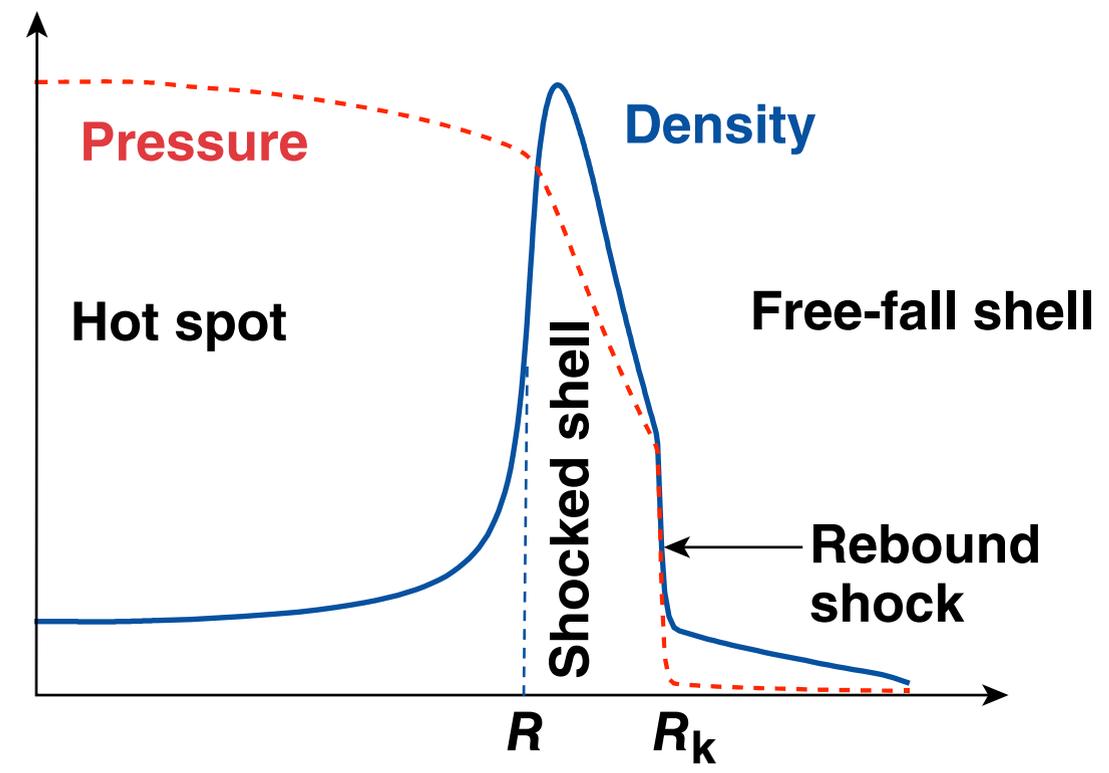
$$\text{Mach} \equiv \frac{V_i}{C_s^{hs}}$$

← Shell Mach number

P. Y. Chang *et al.*, Phys. Rev. Lett. **104**, 135002 (2010);
R. Betti *et al.*, Phys. Plasmas **17**, 058102 (2010);
J. Lindl *et al.*, Phys. Plasmas **21**, 129902(E) (2014).

The second burning-plasma regime requires calculating the total PdV to the hot spot and shell

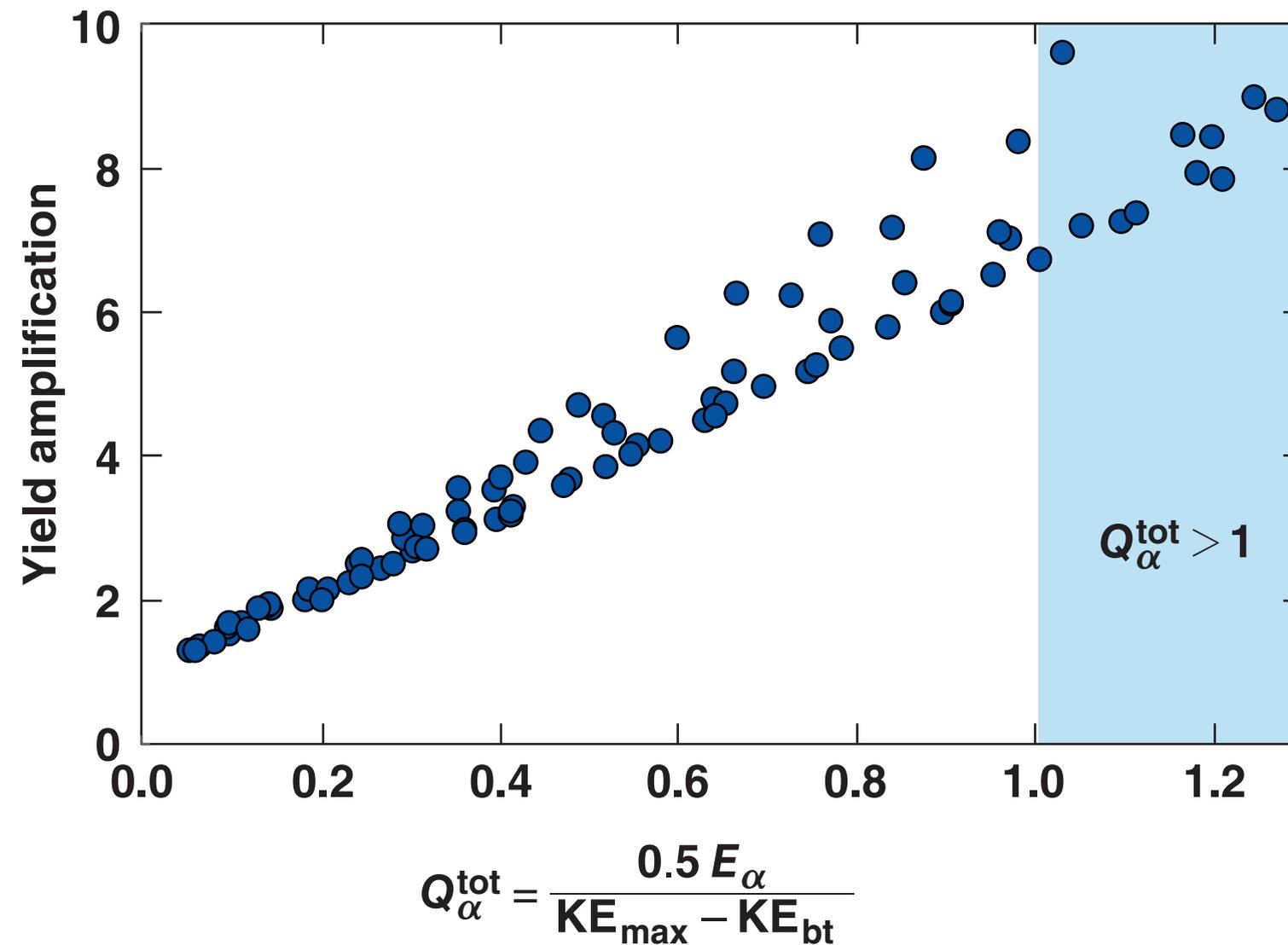
$$Q_{\alpha}^{\text{tot}} = \frac{0.5 E_{\alpha}}{PdV_{\text{hs}} + PdV_{\text{sh}}}$$



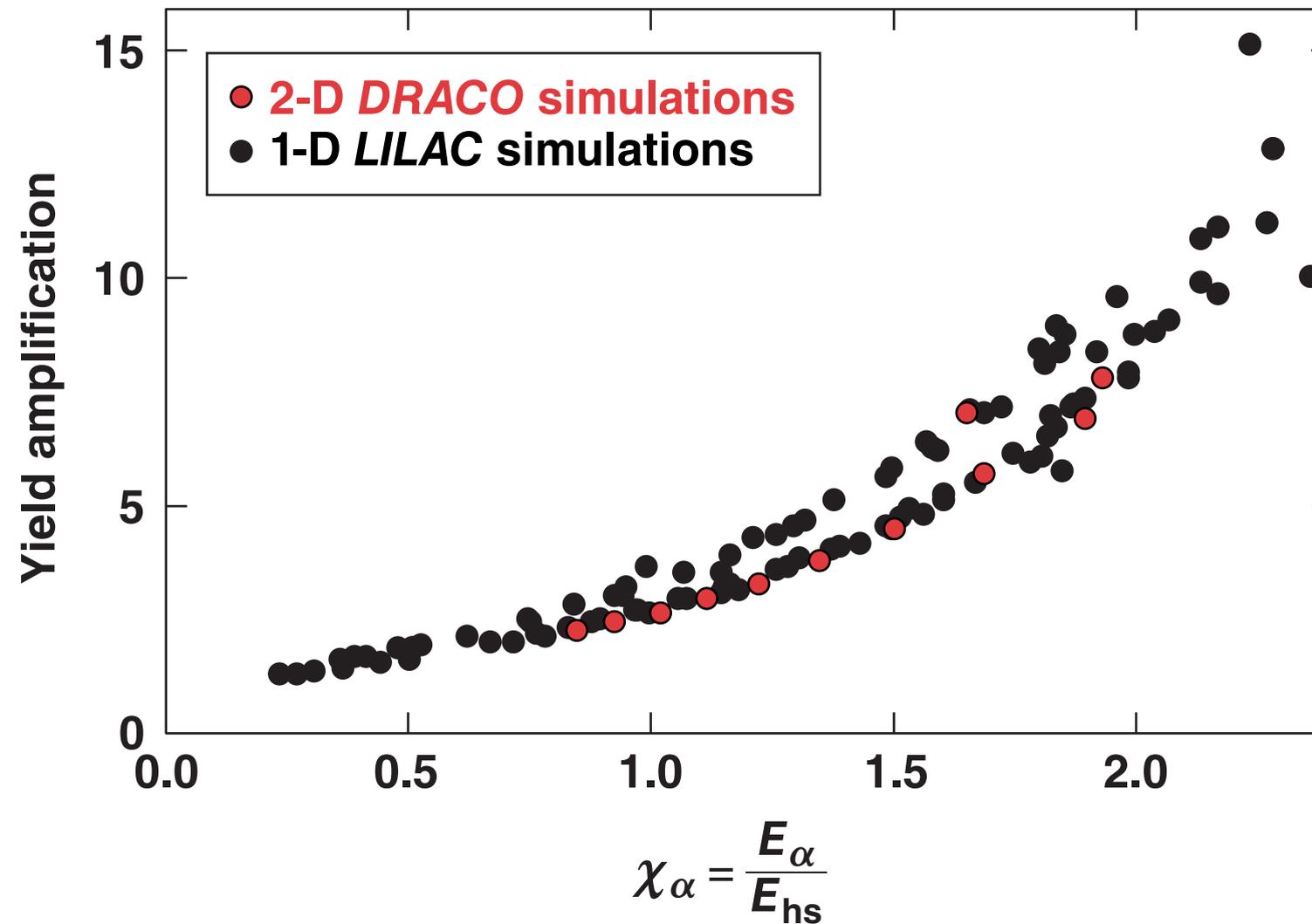
Valid
in 3-D
↓

$$\text{Total } PdV \text{ work} = PdV_{\text{hs}} + PdV_{\text{sh}} = - \int_0^{t_{\text{bang}}} 4\pi P R_k^2 \dot{R}_k dt = KE_{\text{max}} - KE_{\text{bt}}$$

The yield amplification from simulations (30-kJ to 2-MJ energy) is strongly correlated with Q_{α}^{tot}



The yield amplification is mostly a function of the measurable Lawson parameter χ_α



α Heating in Indirect- and Direct-Drive ICF

In the indirect-drive high-foot shot 140120, α heating caused a yield amplification of about 2.3×



$$\langle P \rangle \approx 170 \text{ Gbar}$$

$$\langle T \rangle \approx 5 \text{ keV}$$

$$R_{\text{hs}} \approx 35 \mu\text{m}$$

$$\tau_{\text{burn}} \approx 161 \text{ ps}$$

$$E_{\text{hs}} \approx 4.6 \text{ kJ}$$

$$E_{\alpha} \approx 5.2 \text{ kJ}$$

$$E_{\text{fusion}} \approx 26 \text{ kJ}$$

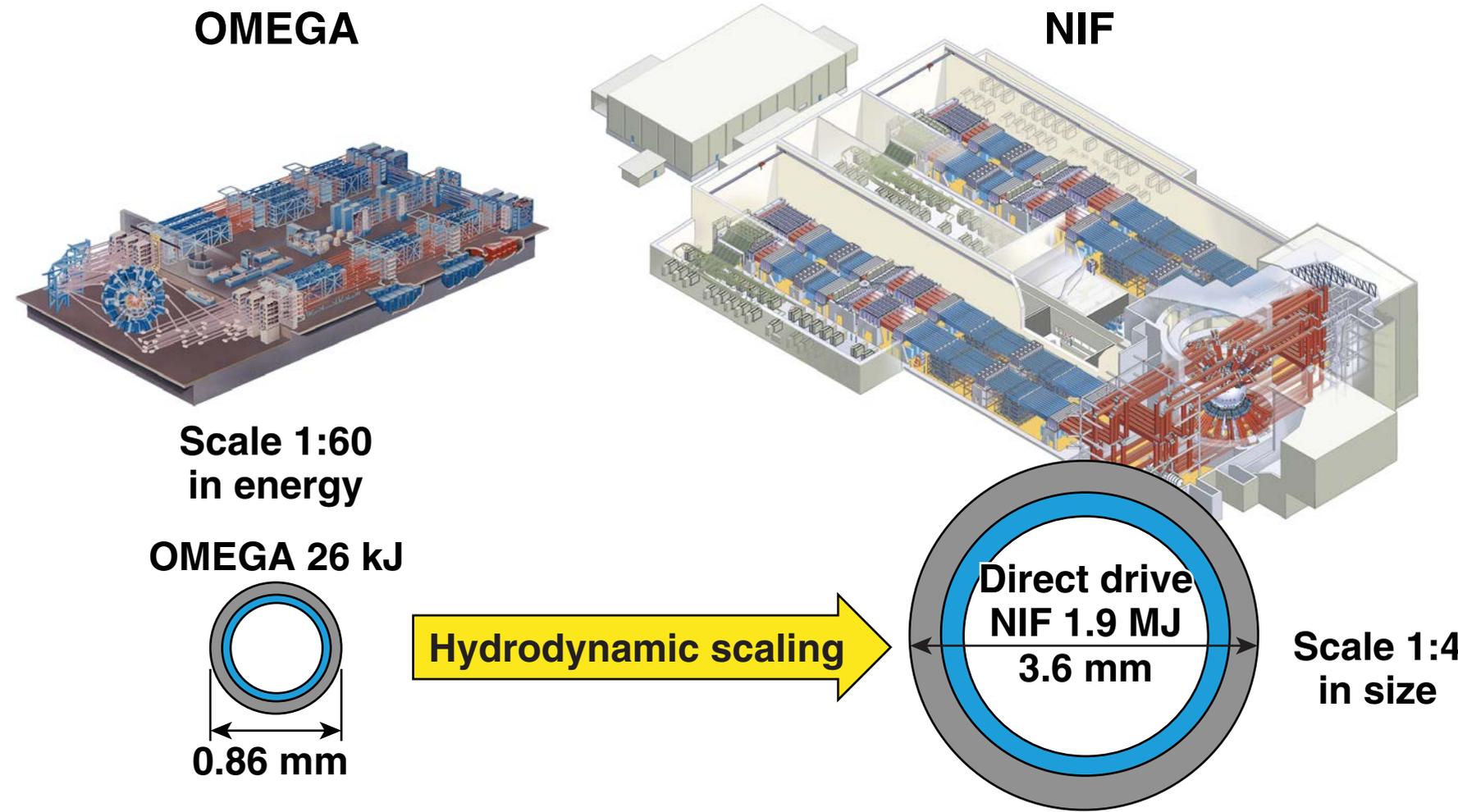
$$\chi_{\alpha} = \frac{E_{\alpha}}{E_{\text{hs}}} \approx 1.1$$

$$\frac{Y_{\alpha}}{Y_{\text{no } \alpha}} \approx 2.3$$

$$Q_{\alpha}^{\text{hs}} \approx 0.55$$

P. K. Patel, Lawrence Livermore National Laboratory, private communication (2014);
T. Döppner *et al.*, Phys. Rev. Lett. **115**, 055001 (2015).

Hydrodynamic equivalence provides a tool to scale the performance of OMEGA direct-drive implosions to NIF energies for symmetric illumination



Hydrodynamic scaling does *not* account for differences in laser-plasma interactions between OMEGA and the NIF.

The hydrodynamic scaling leads to equal pressures at stagnation

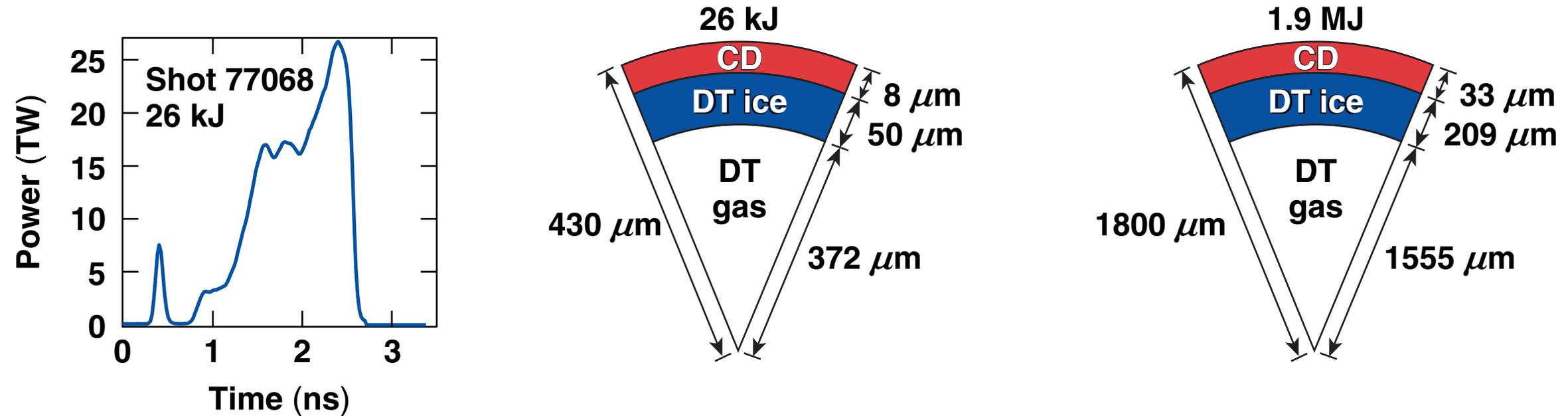
- In-flight scaling: $V_{\text{imp}} \sim \text{const}$ $\alpha \sim \text{const}$ RT * growth factors $\sim \text{const}$
 $R \sim E_L^{1/3}$ $P_L \sim E_L^{2/3}$ $\tau_{\text{pulse}} \sim E_L^{1/3}$

In the absence of α heating:

- Stagnation scaling: $P \sim \text{const}$ $T \sim R^{0.2}$ $V_{\text{hs}} \sim R^3$
 $\tau_{\text{burn}} \sim R$ $\rho R_{\text{tot}} \sim R$

From 26 kJ to 1.9 MJ $\rightarrow P$ is the same; R , τ_{burn} , ρR are up 4 \times , and T is up 30%.

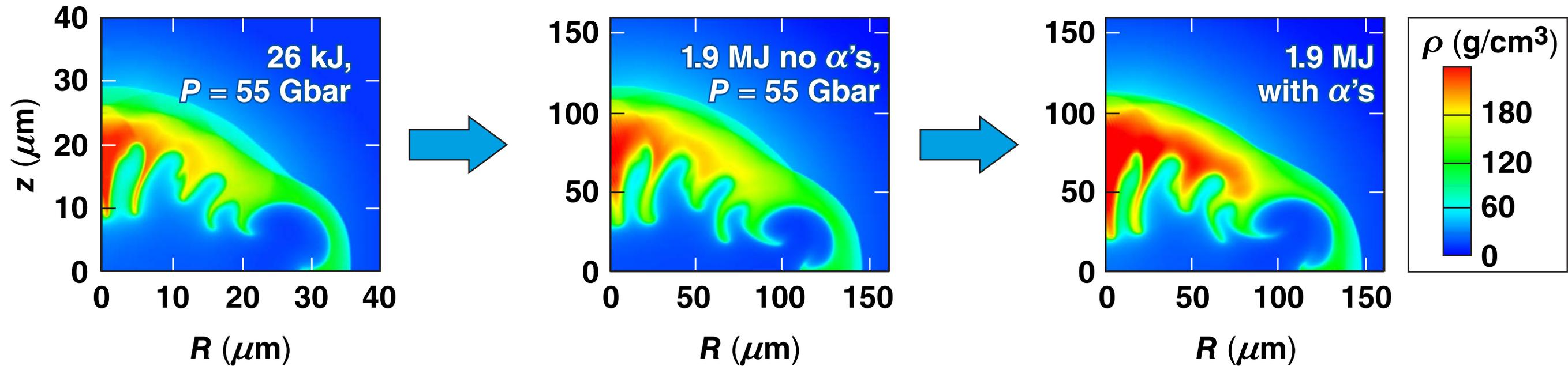
The best-performing OMEGA implosion is readily scaled up to NIF energies



Observable	Experiment	1-D simulation
Yield	5.3×10^{13}	1.7×10^{14}
P (Gbar)	56 (± 7)	97
T_i (keV)	3.6 (± 0.3)	3.82
R_{hs} (μm)	22 (± 1)	22
Burnwidth (ps)	66 (± 10)	61
ρR (g/cm ²)	0.198 (± 0.018)	0.211

A multimode ice perturbation is used to degrade the target performance to reproduce the OMEGA experiment; the same perturbation is applied to the 1.9-MJ target

Density at bang time

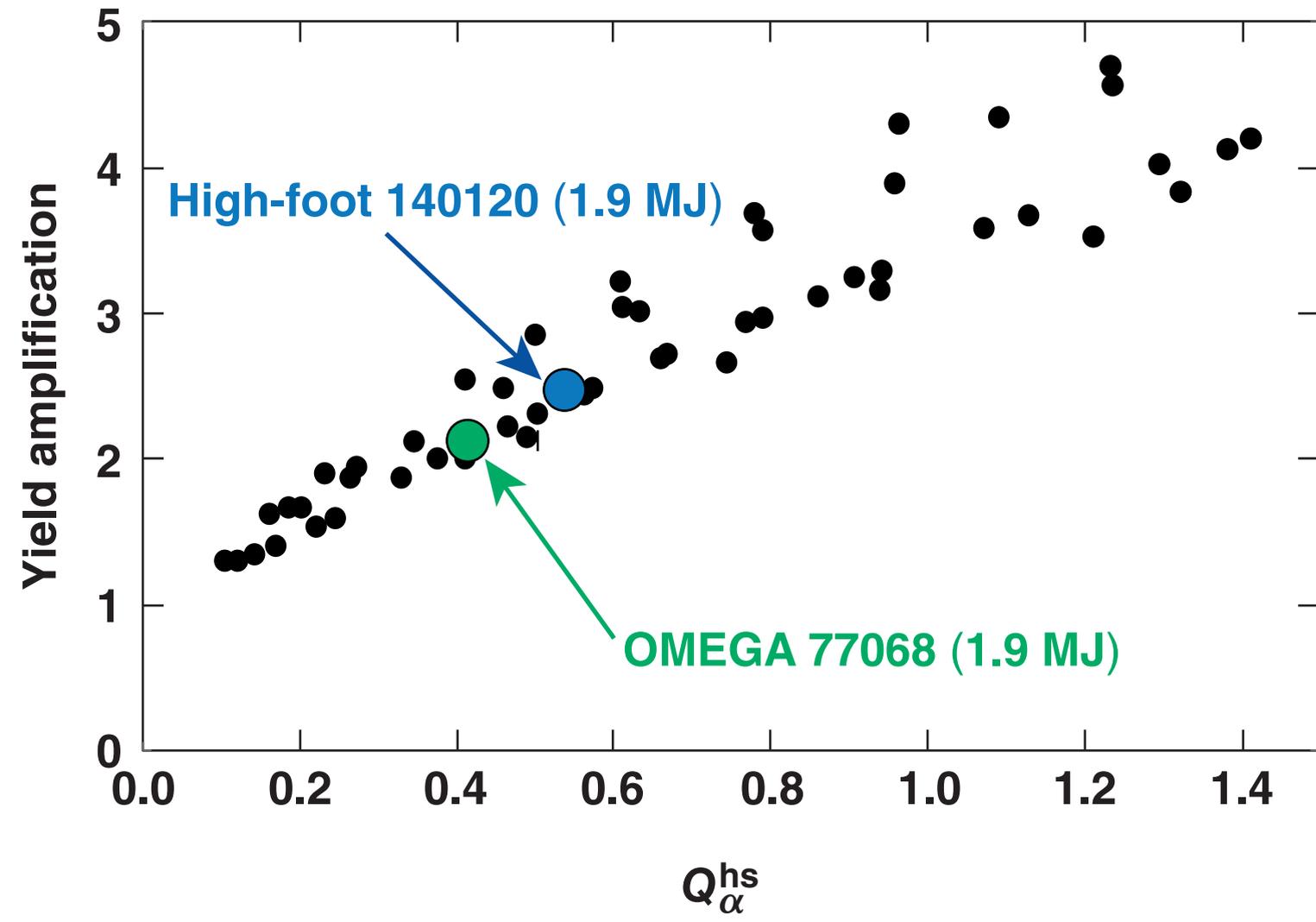


Yield with α 's = 125 kJ
 Yield no α 's = 63 kJ

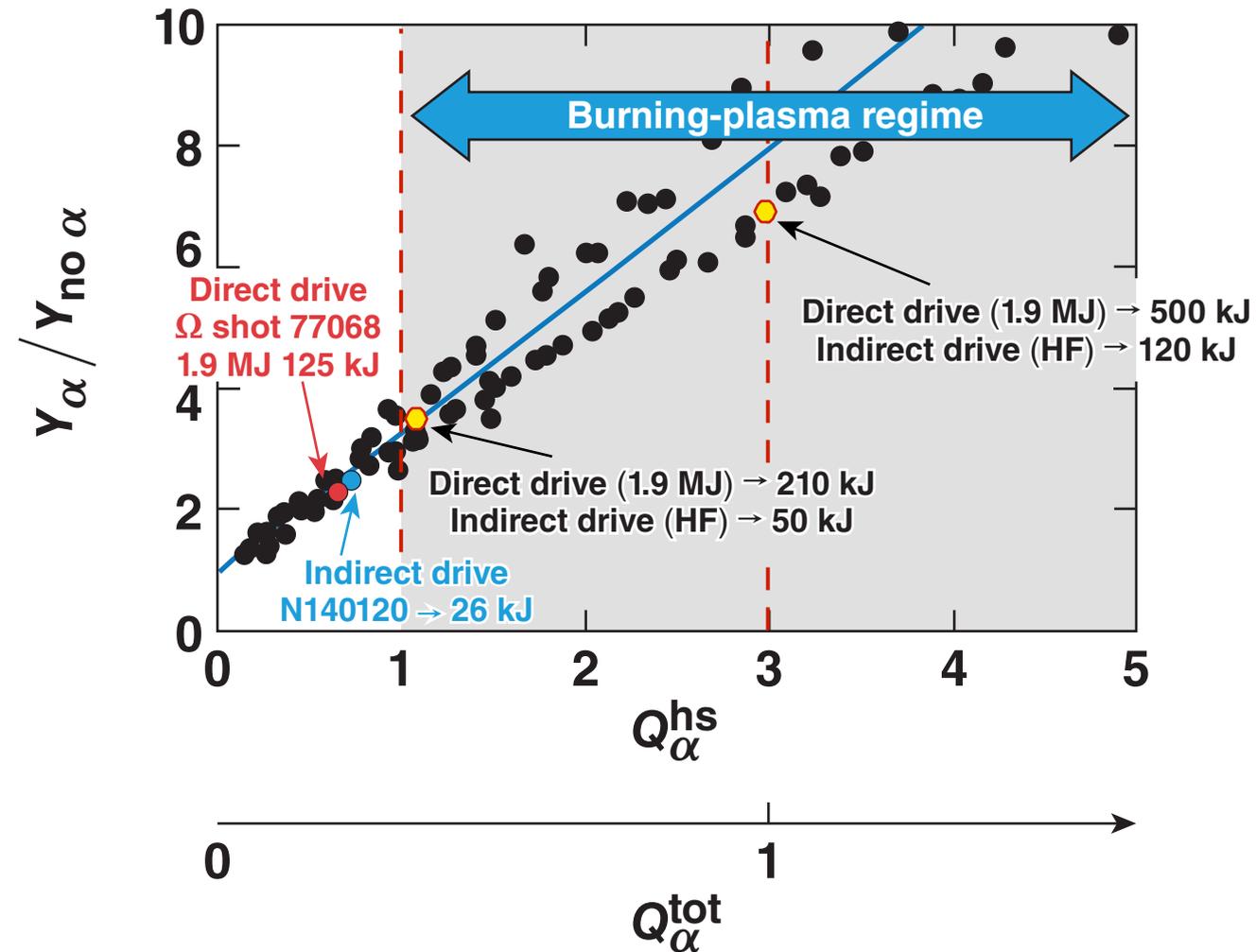
$$\frac{\gamma_{\alpha}^{1.9 \text{ MJ}}}{\gamma_{\text{no } \alpha}^{1.9 \text{ MJ}}} \approx 2$$

A. Bose *et al.*, GO5.00004, this conference;
 K. M. Woo *et al.*, GO5.00003, this conference.

OMEGA shot 77068 scaled up to 1.9 MJ achieves a yield enhancement of $\approx 2\times$ and $Q_\alpha \approx 0.4$



Access to the burning-plasma regime requires about 50 kJ of HF targets in indirect drive and about 200 kJ of fusion energy for direct drive



Both direct and indirect drive must double the yield amplification to access the burning-plasma regime.

The onset of the burning-plasma regime can be identified through experimental observables related to the yield amplification from α heating

- The fundamental parameter characterizing burning plasmas is $Q_{\alpha} = \alpha \text{ heating} / PdV \text{ work}$
- Current high-foot (HF)* implosions at the National Ignition Facility (NIF) have achieved $Q_{\alpha}^{\text{hs}} \approx 0.5$ to 0.6 with a yield amplification, caused by α heating, of about $2.3\times$ at 1.9-MJ laser energy
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R. Nora *et al.*, *Phys. Plasmas* **21, 056316 (2014).