### Polar-Driven, Direct-Drive Implosions on OMEGA: Observations and Simulations of Low-Mode Perturbations in the Main Fuel Layer and Hot Spot



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#### Summary

### Low adiabat, directly driven polar-drive (PD) experiments on OMEGA validate our ability to predict low-mode perturbations expected for the NIF PD point design

UR :

- Low adiabat, directly driven implosions are being performed on OMEGA emulating the configuration of the NIF beams.
- The implosions follow the 1-D simulations up to shell stagnation as evidenced by x-ray radiography and particle measurements.
- Low mode perturbations of the shell, due principally to illumination nonuniformity, are well reproduced by 2-D simulations.
- Perturbations of the hot spot depart from 2-D simulations and may have contributions from the target stalk not included in the simulations.



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### Abel inversion can be used to determine the plasma density from x-ray radiographs\*

Absorption of backlighter x rays along a path follows the relation

$$I = I_0 \exp\left[-\int \kappa(\boldsymbol{E}, \boldsymbol{r}) \, \mathrm{d}\boldsymbol{z}\right].$$

The inverse Abel transform gives the radially dependent opacity

$$\kappa(\boldsymbol{E},\boldsymbol{r}) = \frac{1}{\pi} \int_{\boldsymbol{r}}^{\infty} \frac{\mathrm{d}}{\mathrm{d}\boldsymbol{y}} \left\{ \ln\left[\frac{\boldsymbol{I}(\boldsymbol{y})}{\boldsymbol{I}_0}\right] \right\} \frac{\mathrm{d}\boldsymbol{y}}{\sqrt{\boldsymbol{y}^2 - \boldsymbol{r}^2}}.$$

If the mass absorption coefficient is approximately constant through the plasma, as is the case for bound-free absorption by inner-shell electrons, then O(r) = k(F, r)/k(F, r)

$$\rho(\mathbf{r}) = \kappa(\mathbf{E}, \mathbf{r}) / \mu_{\text{eff}}(\mathbf{E}),$$

where  $\mu_{\text{eff}}(E)$  is the mass absorption coefficient averaged over the effective energy band of the radiograph, and can be determined as follows:

$$M_{\text{shell}} = \int \rho(r) dV = 4\pi \int \frac{\kappa(E, r)}{\mu_{\text{eff}}(E)} r^2 dr.$$

<sup>\*</sup>F. J. Marshall et al., Phys. Rev. Lett. <u>102</u>, 185004 (2009).

# Abel inversion is used to compute the density profiles from framed x-ray radiographs of a polar-driven, direct-drive implosion



### The measured areal-density time history is consistent with 1-D simulations up to bang time



### The observed shell perturbations are accurately reproduced by DRACO 2-D simulations



F. J. Marshall et al., J. Phys. IV France 133, 153 (2006).

 $-\ell n (I/I_0)$ 

### Low-mode perturbations of the hot spot are determined from framed x-ray images of target self-emission



### The observed hot-spot perturbations deviate significantly from *DRACO* 2-D simulations

OMEGA shot 49331, PD implosion Framed x-ray images 2- to 4-keV emission 2.40 ns 2.46 ns 2.51 ns Stalk direction indicated with arrows The differences in measured and simulated shapes may be due to the stalk, whose effect is not included in the simulations. Viewed from  $\theta = 101^{\circ}$ 

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DRACO/Spect3D simulation

 $\overleftarrow{100 \ \mu m}$ 

#### Summary/Conclusions

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- The implosions follow the 1-D simulations up to shell stagnation as evidenced by x-ray radiography and particle measurements.
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## The density distributions are determined for each time from the radiographs and can be compared to simulation



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#### 40 of the OMEGA beams are used to emulate the NIF 48 beam indirect-drive configuration



from 21° to 59°, are used to

emulate the NIF geometry.

 Additional OMEGA beams are used for x-ray backlighting.