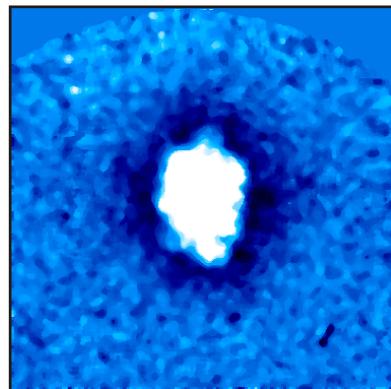


Low-Adiabat, Polar-Drive-Implosion Experiments on OMEGA

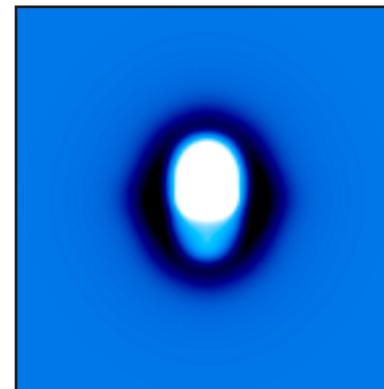


OMEGA shot 52136, Saturn target, 15-atm-D₂-filled,
24-μm CH shell, LA1501 pulse shape



Experiment

200 μm



DRACO/Spect3D

Ti backlit images
~4.7 keV
View angle $\theta_v = 63^\circ$

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University of Rochester
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50th Annual Meeting of the
American Physical Society
Division of Plasma Physics
Dallas, TX
17–21 November 2008

Summary

Low-adiabat, polar-drive experiments on OMEGA are validating the NIF designs



- Shaped pulses are used to keep the main fuel layer on a low adiabat (~3).
- Beam repointing is used to optimize the implosion symmetry on both standard and Saturn-type targets.
- Measurements of implosion core size and shape determined from framed x-ray radiographs compare favorably to 2-D *DRACO* simulations.
- Inferred areal densities agree with values determined from D³He proton spectra.

Recent publications:

F. J. Marshall *et al.*, J. Phys. IV France 133, 153 (2006).
J. A. Marozas *et al.*, Phys. Plasmas 13, 056311 (2006).

Related Talk:

A. Shvydky (TO5.00007).

Collaborators



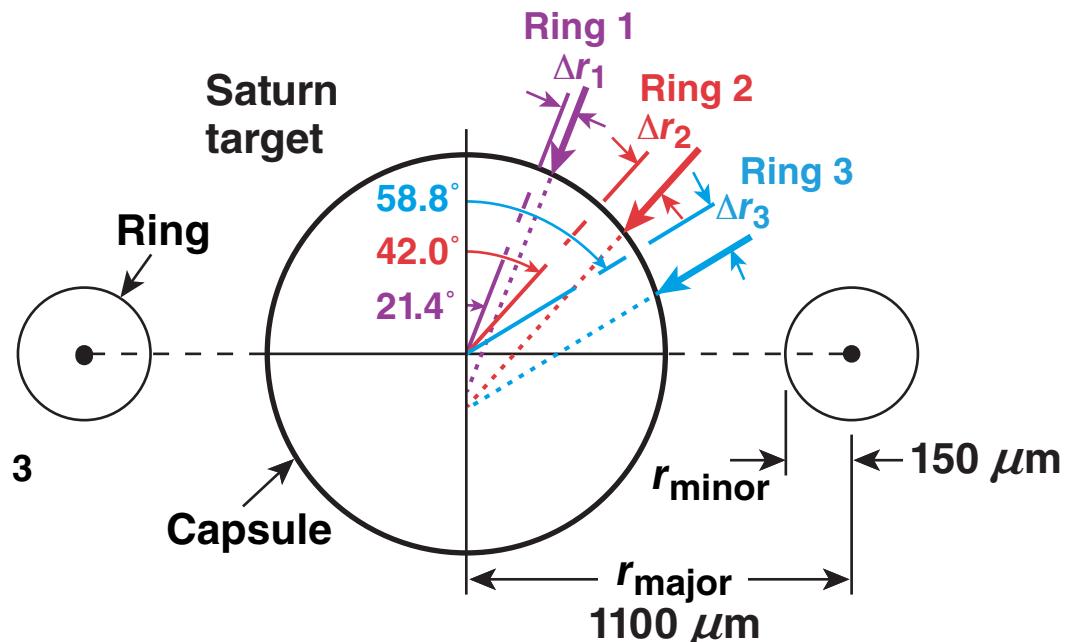
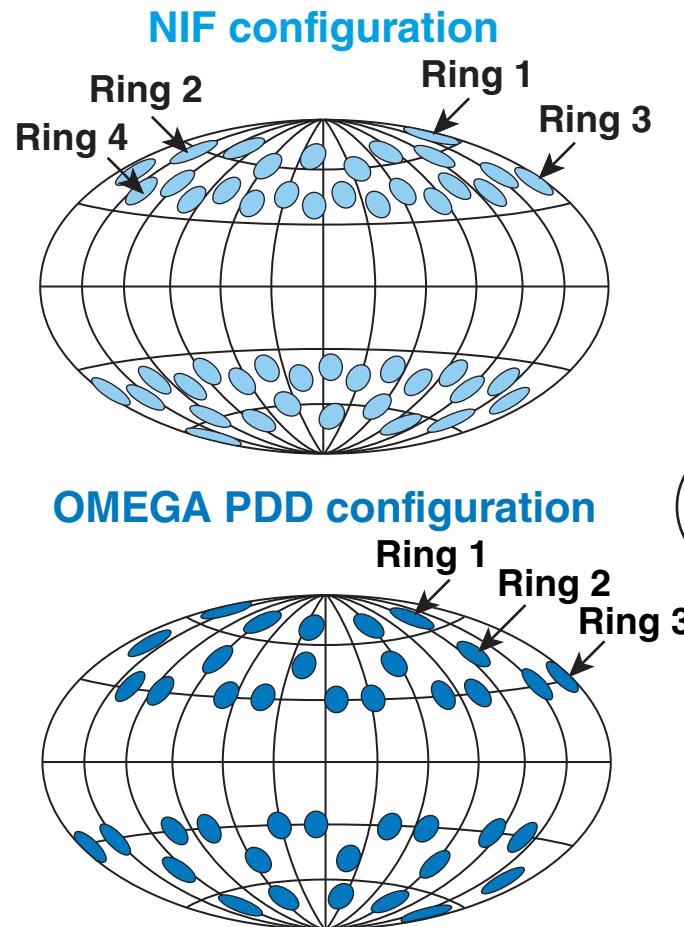
**R. S. Craxton, R. Epstein, V. Yu. Glebov, V. N. Goncharov, J. P. Knauer,
P. W. McKenty, D. D. Meyerhofer, P. B. Radha, T. C. Sangster,
W. Seka, S. Skupsky, and V. Smalyuk**

**University of Rochester
Laboratory for Laser Energetics**

J. A. Frenje, C. K. Li, R. D. Petrasso, and F. H. Séguin

**Massachusetts Institute of Technology
Plasma Fusion Center**

Forty OMEGA beams are used to emulate the NIF 192 beam (48 quad) indirect-drive configuration

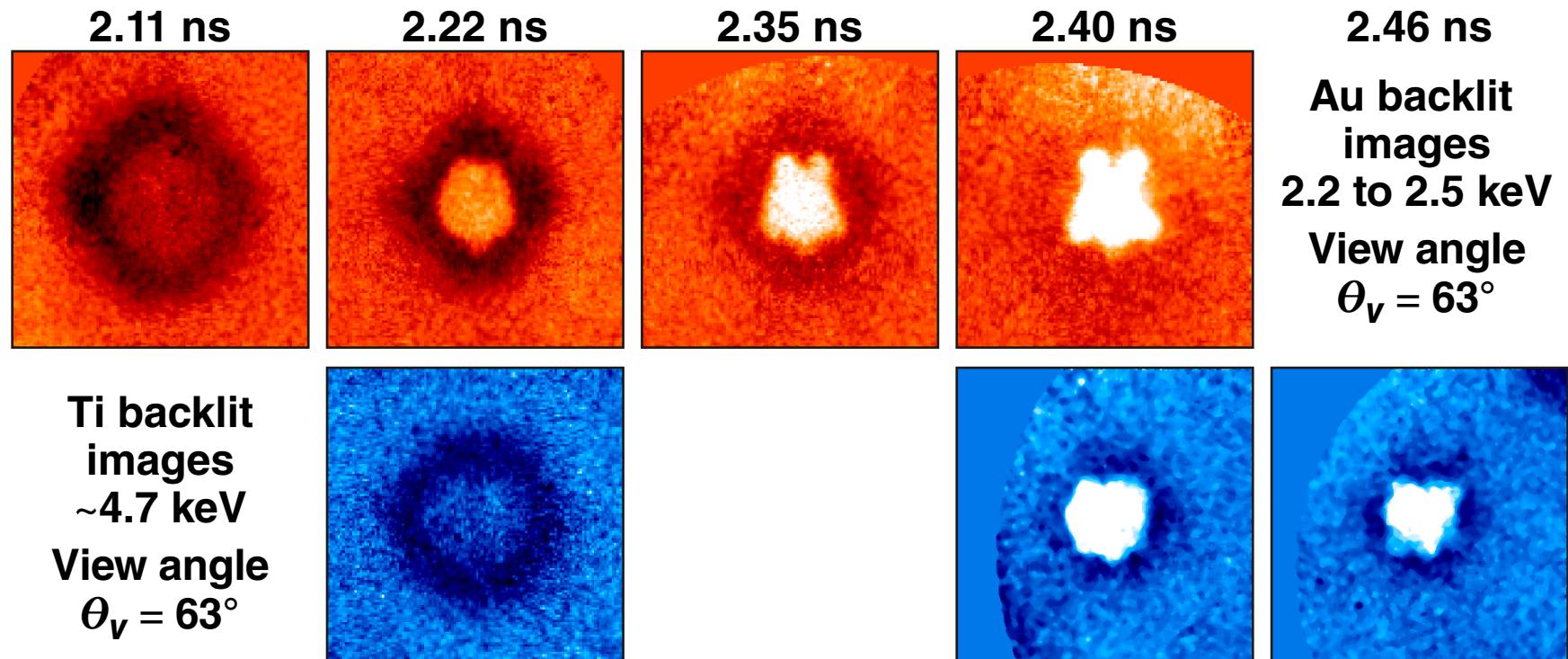


- The OMEGA beams, in six rings from 21° to 59° , are used to emulate the NIF geometry.
- Additional OMEGA beams are used for x-ray backlighting.

Comparison of low- and high-energy backlit images reveals details of the evolution of the fuel and shell



OMEGA shot 49331, polar-driven standard target, 15-atm-D₂-filled,
24-μm CH shell, LA1501 pulse shape

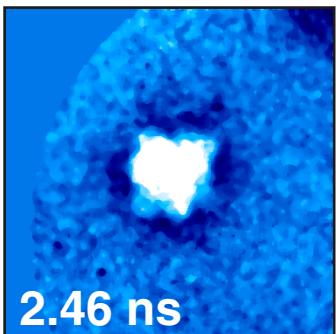
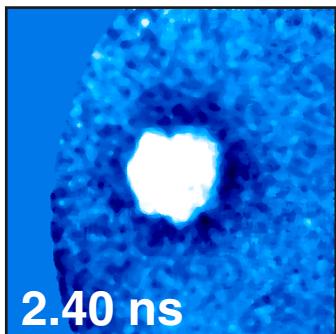
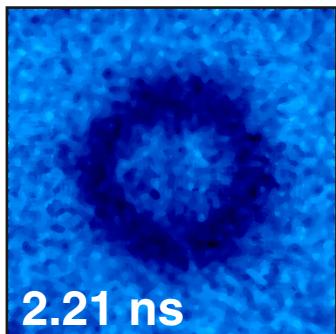
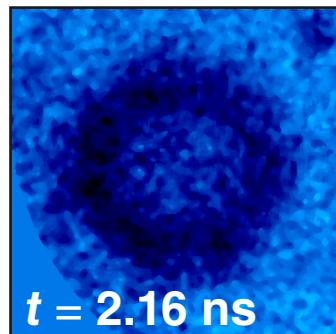


The higher-energy backlighter more clearly delineates the shell and fuel at stagnation.

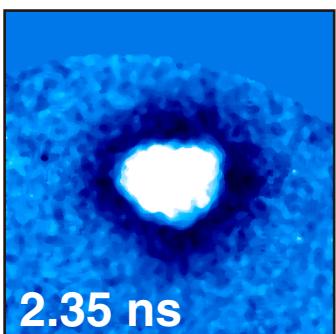
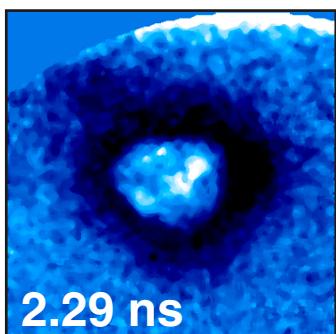
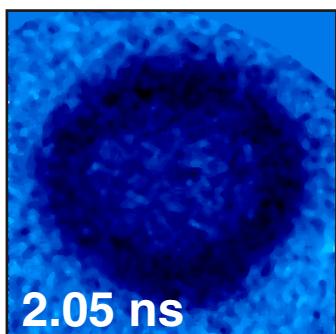
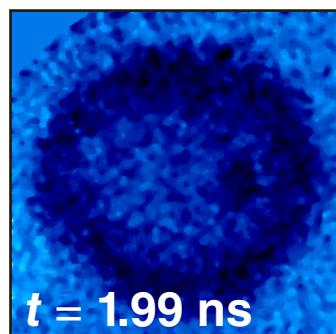
Standard targets require a larger beam offset to optimize implosion symmetry *



Polar-driven standard target, 15-atm-D₂-filled,
24-μm CH shell, LA1501 pulse shape

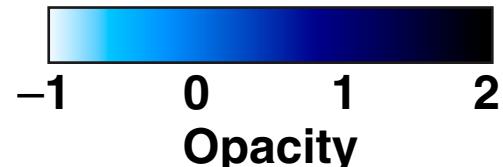


Shot 49331
90,150,150-μm
offset pointing
Ti backlit images
~4.7 keV



Shot 52135
90,120,120-μm
offset pointing
Sc backlit images
~4.3 keV

400 × 400-μm
regions



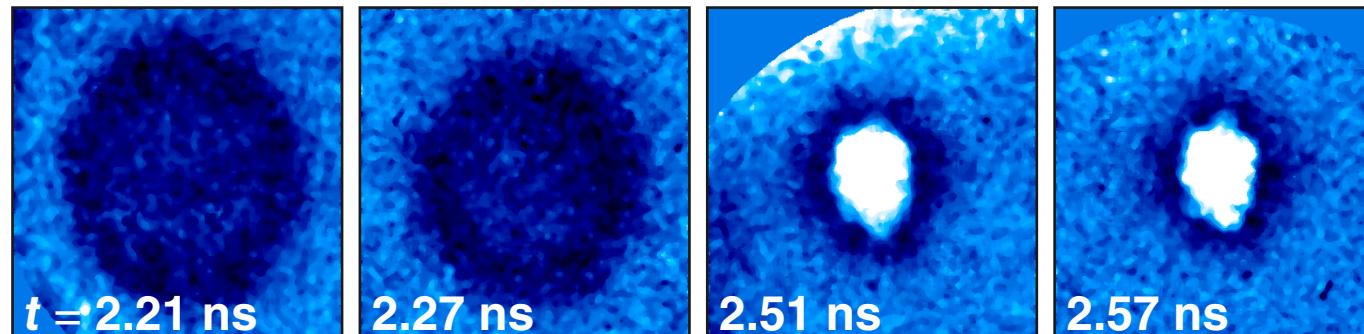
(<0 is emission)

View angle $\theta_V = 63^\circ$

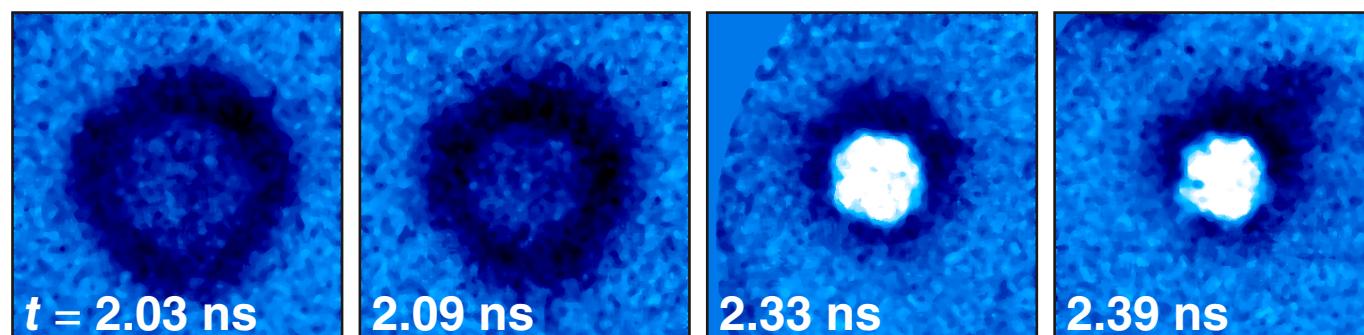
Saturn-target implosion symmetry is achieved with a smaller beam offset relative to standard targets*



Polar-driven Saturn target, 15-atm-D₂-filled,
24-μm CH shell, LA1501 pulse shape



Shot 49333
90,150,150-μm
offset pointing
Ti backlit images
~4.7 keV



Shot 52136
90,120,120-μm
offset pointing
Sc backlit images
~4.3 keV

$400 \times 400\text{-}\mu\text{m}$
regions



(<0 is emission)

View angle $\theta_V = 63^\circ$

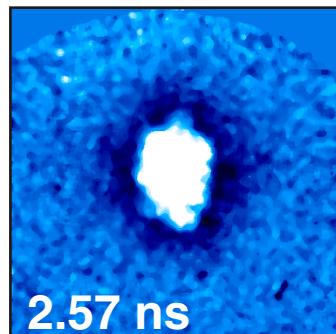
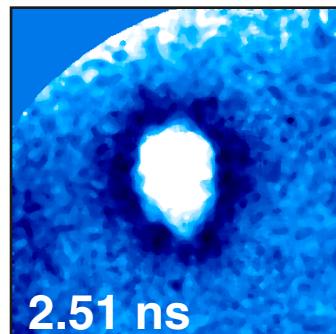
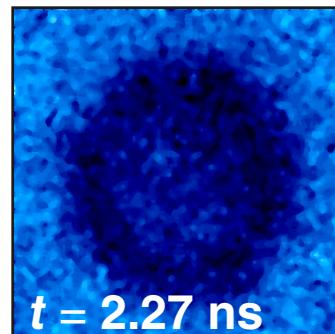
X-ray radiographs and 2-D DRACO simulations* show good agreement of both shape and size of the core



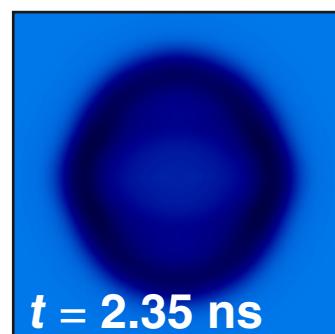
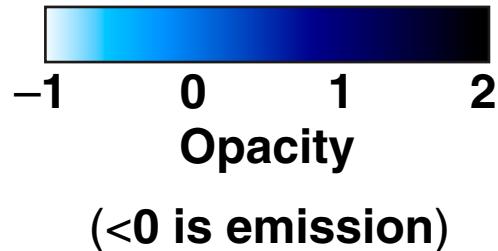
OMEGA shot 49333, polar-driven Saturn target, 15-atm-D₂-filled,
24-μm CH shell, LA1501 pulse shape

400 × 400-μm
regions

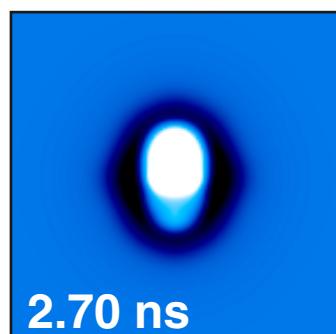
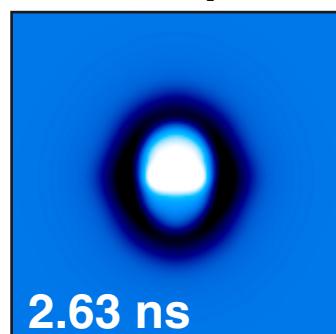
View angle $\theta_v = 63^\circ$



Ti backlit
images
~4.7 keV



DRACO/Spect3D



Simulated
backlit
images

Prolate implosion

Identifying the fuel/shell interface and the outer edge of the shell allows the areal density to be estimated from the x-ray radiographs



$$\rho R_{\text{shell}} = 3 \rho_0 r_0^2 \frac{(\Delta r_0 - \Delta r_{\text{abl}})}{(r_{\text{out}}^3 - r_{\text{in}}^3)} \times (r_{\text{out}} - r_{\text{in}})$$

$$\rho R_{\text{fuel}} = \rho_0 \left(\frac{r_0}{r_{\text{in}}} \right)^3 r_{\text{in}},$$

r_0 = initial radius

ρ_0 = initial density

Δr_0 = initial shell thickness

Δr_{abl} = ablated shell thickness

r_{in} = final fuel/shell interface radius

r_{out} = final shell outer radius

Determination of ρR from radiographs compares favorably with values determined from proton spectra



OMEGA shot 49331, polar-driven standard target

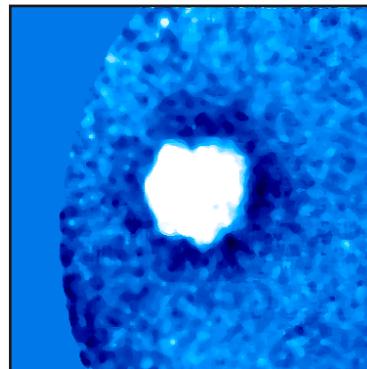
ρR values
(mg/cm²)

$$t = 2.40 \text{ ns}$$

$$\rho R_{\text{CH}} = 50 \pm 5$$

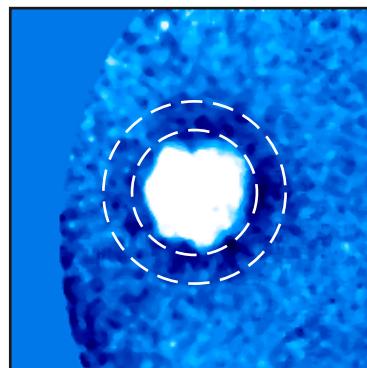
$$\rho R_{D_2} = 5 \pm 1$$

$$\rho R_{\text{total}} = 55 \pm 5$$



$$\langle \rho R \rangle_p = 58 \pm 7$$

$$\rho R_{\text{LILAC}} = 88$$



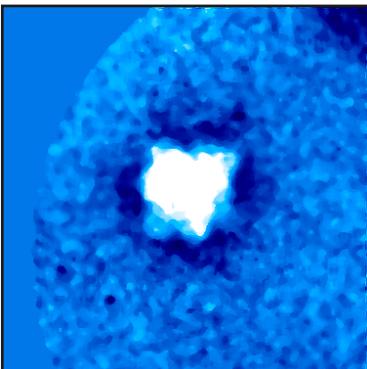
ρR values
(mg/cm²)

$$t = 2.46 \text{ ns}$$

$$\rho R_{\text{CH}} = 54 \pm 5$$

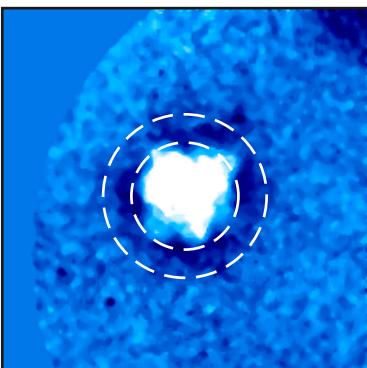
$$\rho R_{D_2} = 7 \pm 1$$

$$\rho R_{\text{total}} = 61 \pm 5$$



$$\langle \rho R \rangle_p = 58 \pm 7$$

$$\rho R_{\text{LILAC}} = 88$$



The values determined at the two times bracket the value determined from D³He protons.*

Summary/Conclusions

Low-adiabat, polar-drive experiments on OMEGA are validating the NIF designs



- Shaped pulses are used to keep the main fuel layer on a low adiabat (~3).
- Beam repointing is used to optimize the implosion symmetry on both standard and Saturn-type targets.
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