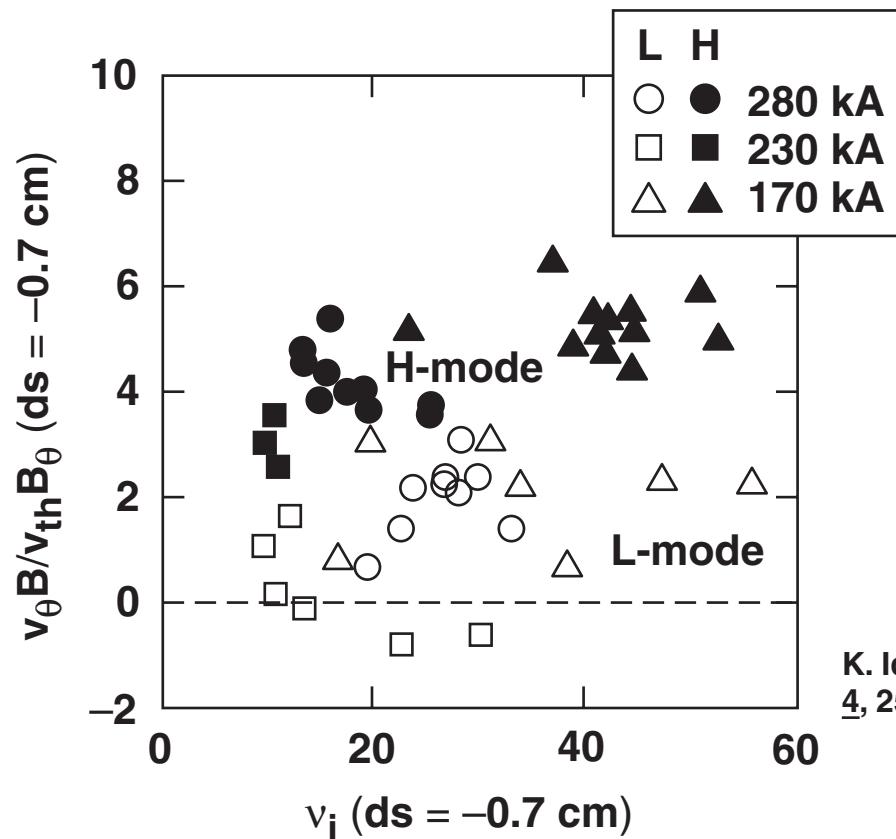


# MHD Equilibria with Poloidal and Toroidal Flow



K. Ida et al., Phys. Fluids B  
4, 2552 (1992).

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Laboratory for Laser Energetics

46th Annual Meeting of the  
American Physical Society  
Division of Plasma Physics  
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# Acknowledgements

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- **R. Betti, University of Rochester**
- **J. Manickam and S. Kaye, PPPL**
- **J. P. Freidberg, MIT**
- **J.-L. Gauvreau, UCLA**

# Outline

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- **Review of theory and numerics (with the code FLOW<sup>1</sup>) of tokamak equilibria with flow**
- **Equilibria with toroidal flow, effects of flow and anisotropy (application to NSTX)**
- **Equilibria with poloidal flow:**
  - **Equilibria with transonic poloidal flow**
  - **Equilibria with super-Alfvénic poloidal flow**
    - inward-shifted equilibria
    - quasi-omnigenous equilibria

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<sup>1</sup>L. Guazzotto, R. Betti, J. Manickam and S. Kaye, Phys. Plasmas **11**, 604 (2004).  
FLOW web site: [http://www.me.rochester.edu/~guazzott/FLOW\\_manual.htm](http://www.me.rochester.edu/~guazzott/FLOW_manual.htm)

# MHD equilibrium equations with flow

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- **Continuity:**

$$\nabla \cdot (\rho \vec{v}) = 0$$

- **Momentum:**

$$\rho \vec{v} \cdot \nabla \vec{v} = \vec{J} \times \vec{B} - \nabla \cdot \vec{P}$$

$$\vec{P} \equiv p_{\perp} \vec{I} + \Delta \vec{B} \vec{B} \quad \Delta \equiv (p_{||} - p_{\perp}) / B^2$$

- **Maxwell equations (and Ohm's law):**

$$\nabla \times (\vec{v} \times \vec{B}) = 0 \quad \nabla \cdot \vec{B} = 0$$

# The MHD equations are reduced to a “Bernoulli” and a “Grad–Shafranov” equation



- Plasma flow

$$\vec{v} = \mathbf{M}_{A\theta} \vec{v}_A + R \Omega(\Psi) \hat{\mathbf{e}}_\phi \quad \mathbf{M}_{A\theta} = \mathbf{v}_\theta / v_{A\theta} = \Phi(\Psi) / \sqrt{\rho}$$

- “Bernoulli” equation

$$\frac{1}{2} \frac{(\mathbf{M}_{A\theta} \mathbf{B})^2}{\rho} - \frac{1}{2} [R \Omega(\Psi)]^2 + w = H(\Psi)$$

- “GS” equation

Enthalpy

$$\nabla \cdot \left[ \left( 1 - \mathbf{M}_{A\theta}^2 - \Delta \right) \left( \frac{\nabla \Psi}{R^2} \right) \right] =$$

$$- \frac{\partial p_{||}}{\partial \Psi} - \frac{B_\phi}{R} \frac{d F(\Psi)}{d \Psi} - \vec{v} \cdot \vec{B} \frac{d \Phi(\Psi)}{d \Psi} - R \rho v_\phi \frac{d \Omega(\Psi)}{d \Psi} - \rho \frac{d H(\Psi)}{d \Psi} + \rho \frac{\partial w}{\partial \Psi}$$

E. Hameiri, Phys. Fluids 26, 230 (1983).

R. Iacono et al., Phys. Fluids B 2, 1794 (1990).

# The input of the code FLOW uses quasi-physical free functions



- Each physical quantity reduces to the corresponding free function in the cylindrical limit.
- The input functions can be supplied as analytical expressions or numerical tables.

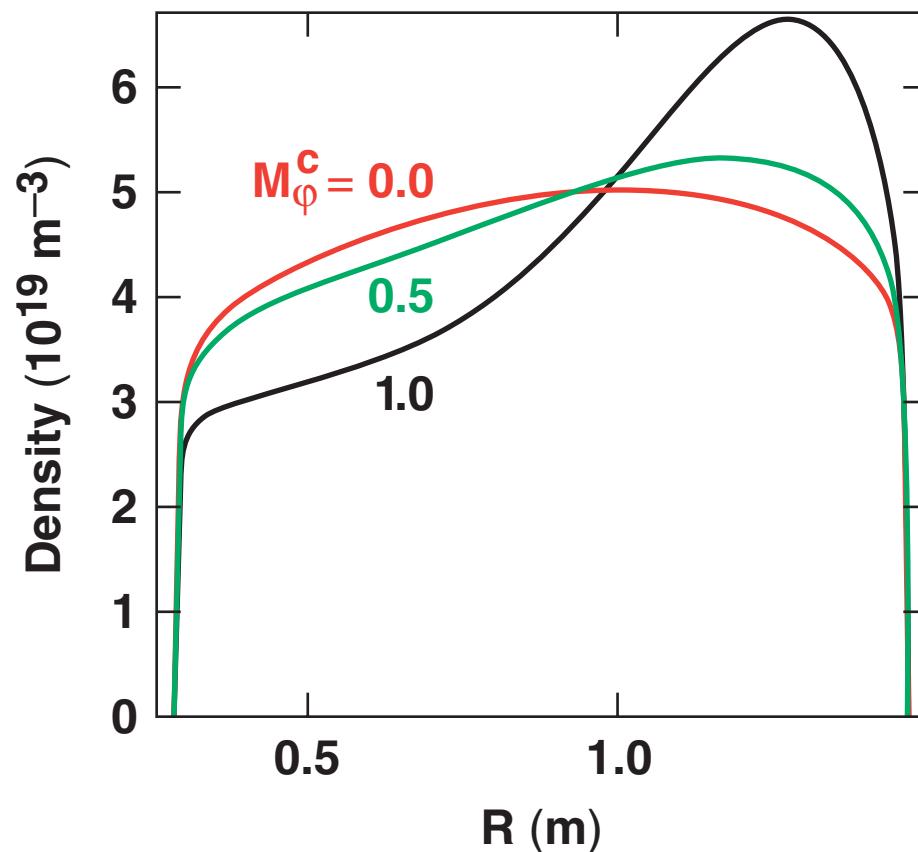
$D(\Psi)$	→ Quasi-density
$P_{  }(\Psi)$	→ Quasi-parallel pressure
$P_{\perp}(\Psi)$	→ Quasi-perpendicular pressure
$B_0(\Psi)$	→ Quasi-toroidal magnetic field
$M_\theta(\Psi)$	→ Quasi-poloidal sonic Mach number
$M_\phi(\Psi)$	→ Quasi-toroidal sonic Mach number

# Equilibria with Purely Toroidal Flow. Applications to NSTX



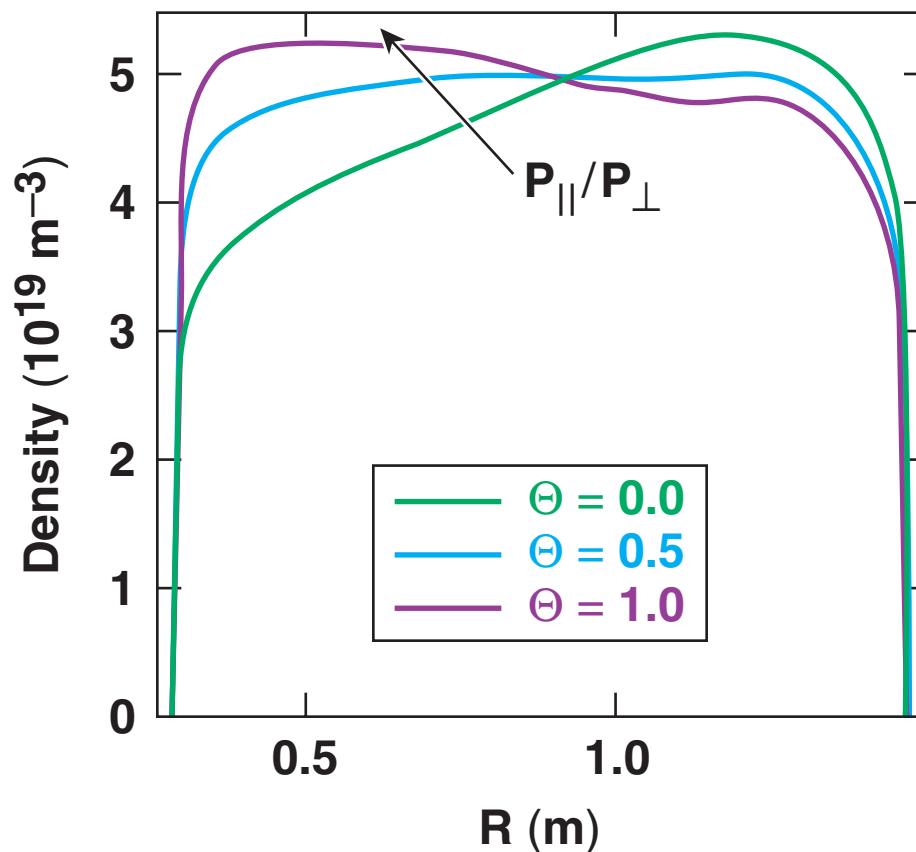
$$\begin{aligned}\mathbf{v}_\phi &= \Omega(\Psi)\mathbf{R} \\ \mathbf{v}_\theta &= \mathbf{0}\end{aligned}$$

# The centrifugal force causes an outward shift of the plasma



**Effect of increasing rotation with constant plasma total mass for NSTX-like equilibria**

# The parallel anisotropy ( $p_{||} > p_{\perp}$ ) causes an inward shift



$$\Theta = \frac{P_{||}(\Psi) - P_{\perp}(\Psi)}{P_{\perp}(\Psi)}$$

(Toroidal Mach number)  
 $M_{\phi}^c = 0.5$

Total energy is conserved.

NSTX-like parameters

# **Equilibria with Poloidal Flow**



# Viscosity is reduced in supersonic flows and omnigenous B-field



- Poloidal flows in tokamaks are damped.
- Poloidal viscosity  $\nu_\theta \sim 1/M_p^2$

$$M_p = \frac{v_\theta}{C_{s\theta}}$$

$$C_{s\theta} = C_s \frac{B_\theta}{B}$$

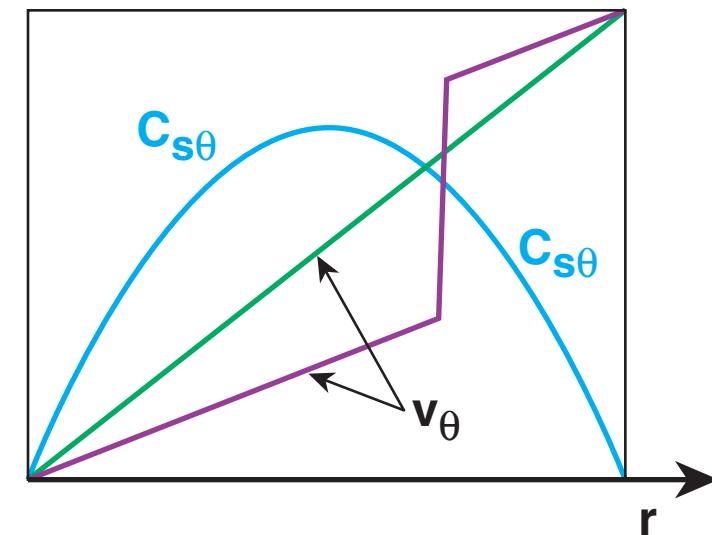
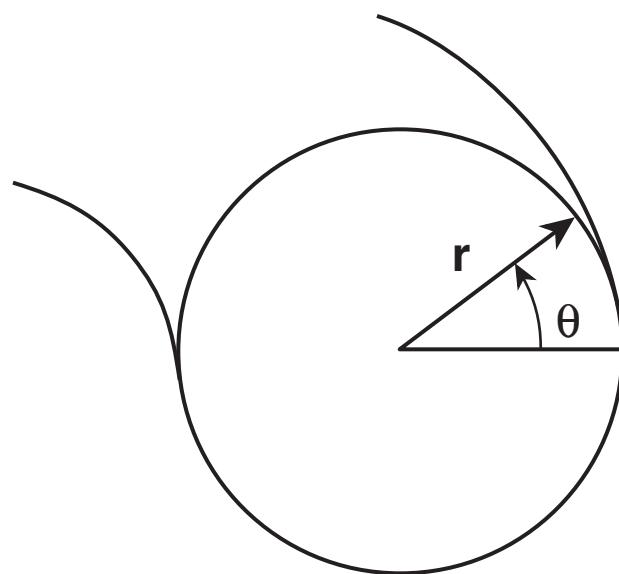
Poloidal  
sound speed

- Equilibria with supersonic poloidal flow have reduced viscosity.
- Omnigenous equilibria have low neoclassical viscosity.
- Specific applications shown for the UCLA Electric Tokamak;  
**Results have general applicability.**

# Equilibria with transonic poloidal flow: flow profile ranging from subsonic to supersonic

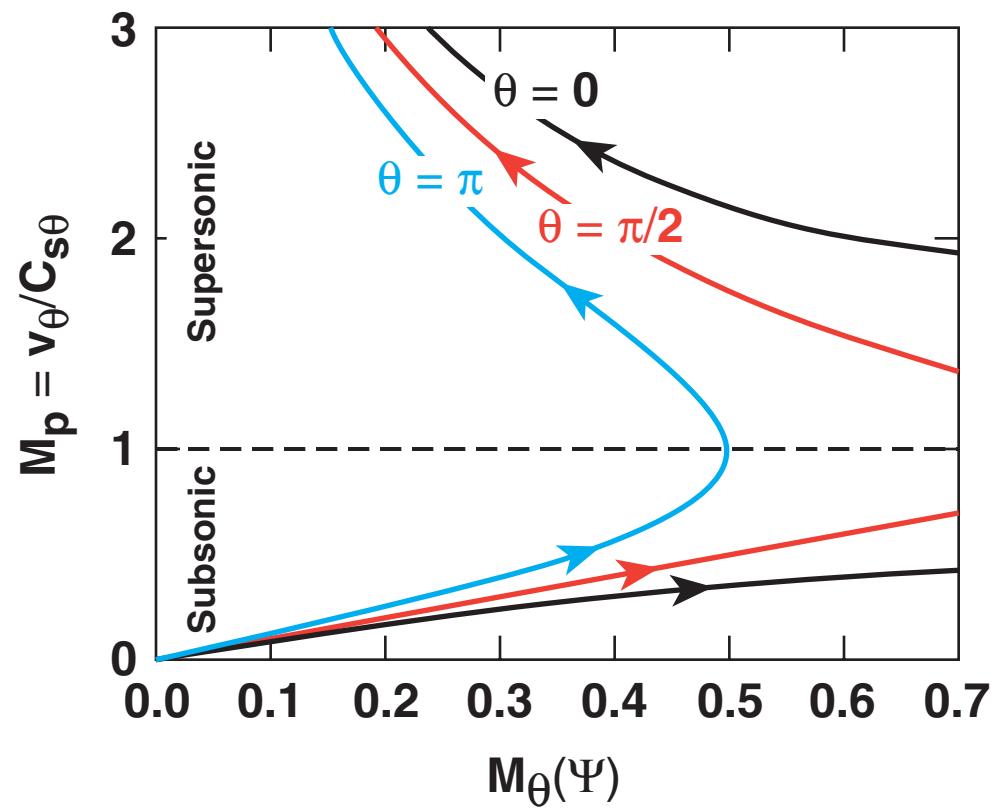
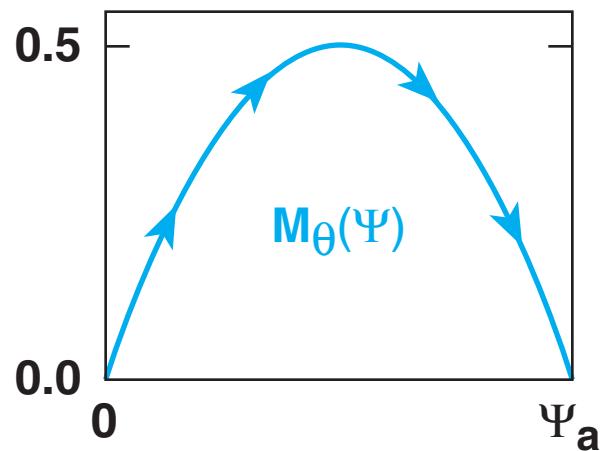


$$v_\theta \sim C_{s\theta} = C_s \epsilon / q$$



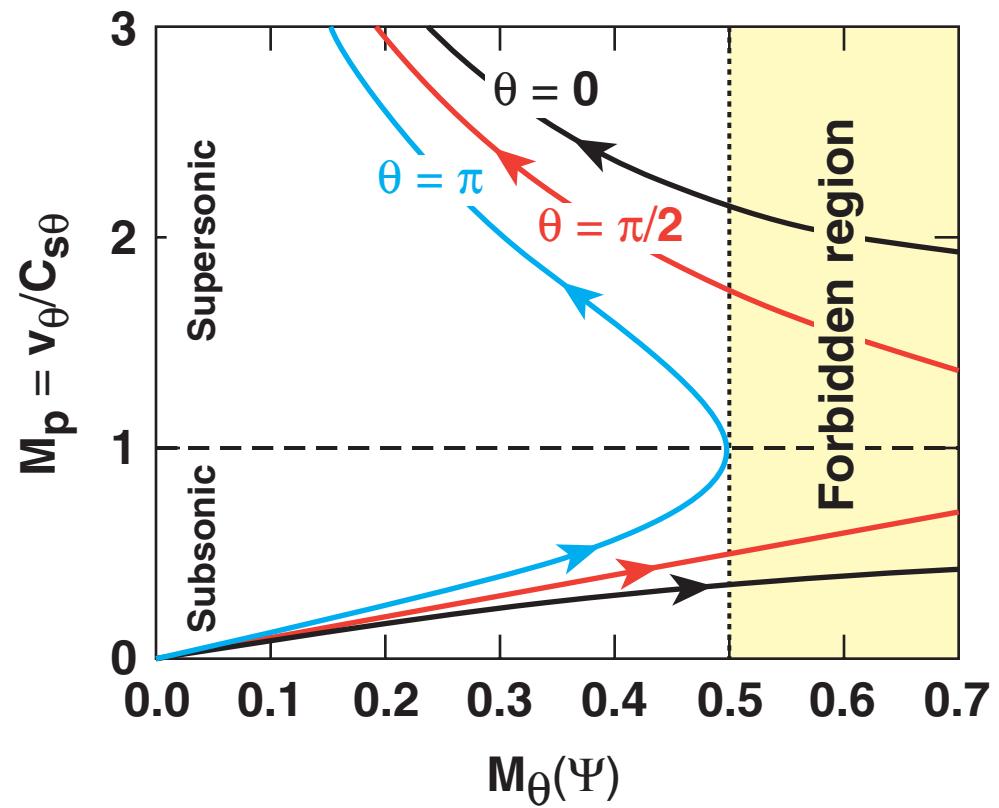
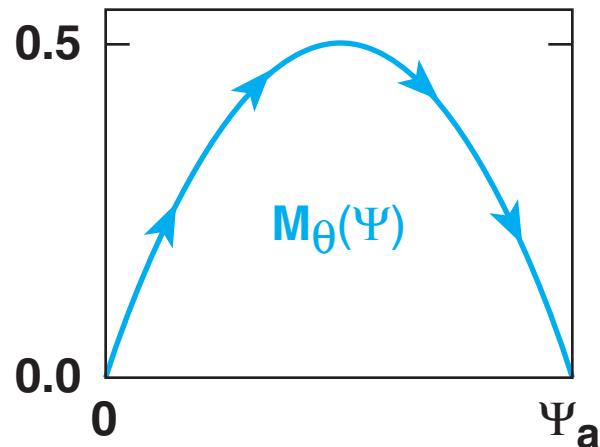
# Transonic solution of Bernoulli equation is discontinuous and imposes constraints on the free functions

- $M_\theta(\Psi)$  cannot be chosen arbitrarily.



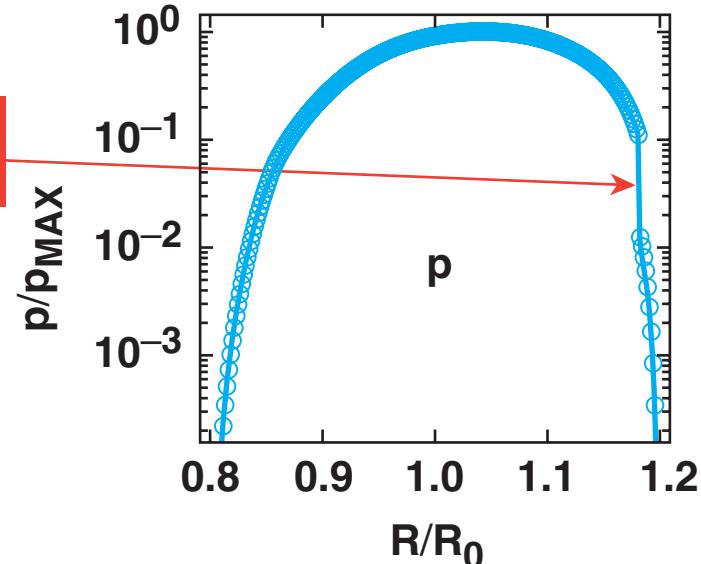
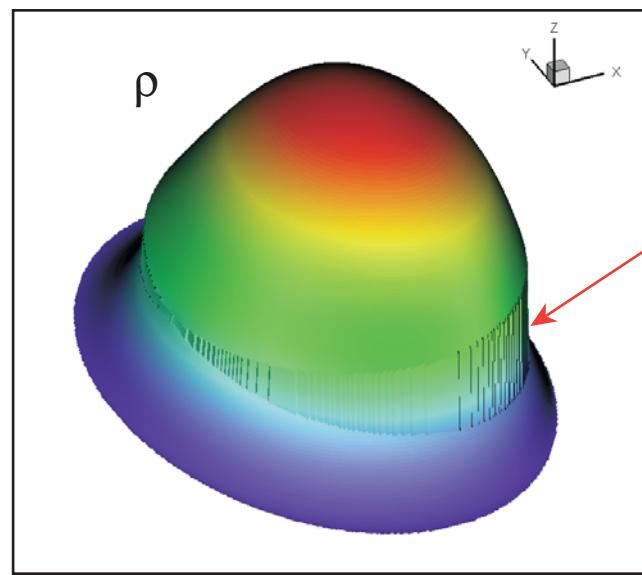
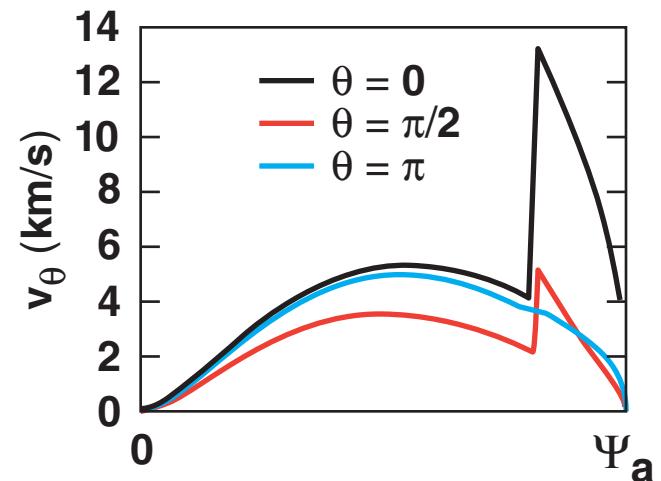
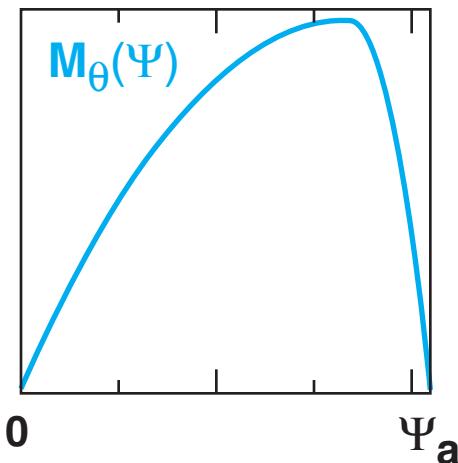
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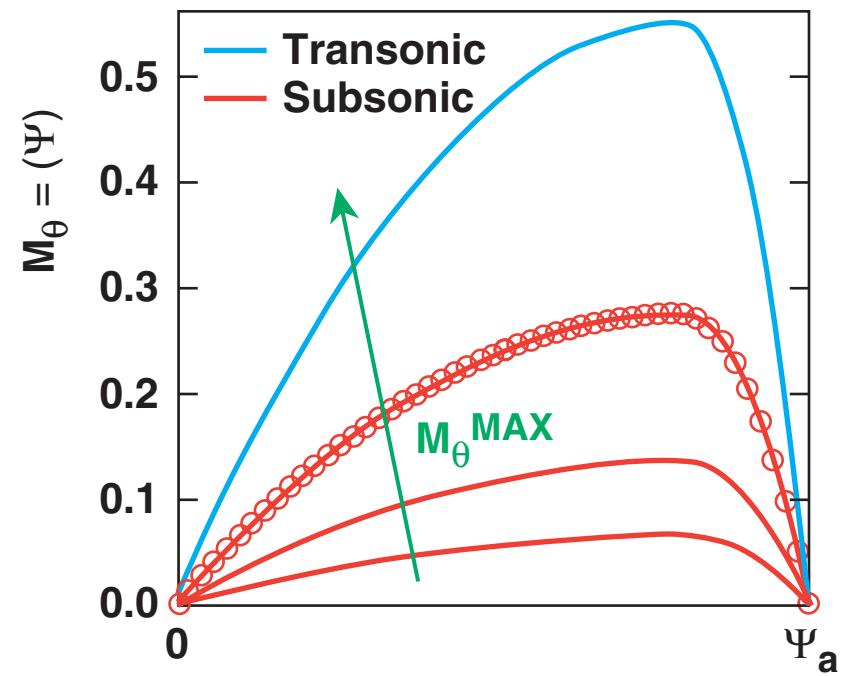
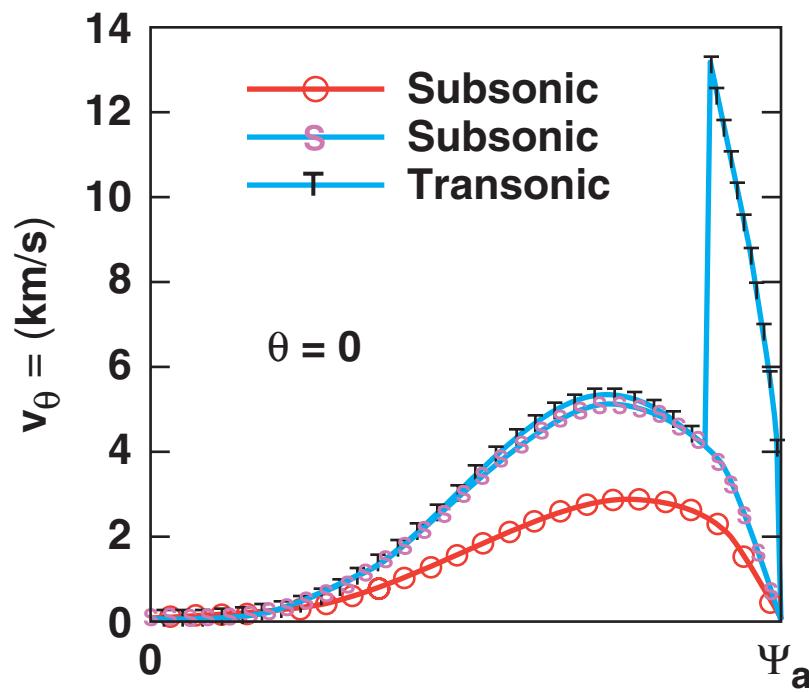
# Numerical transonic equilibria exhibit an edge pedestal structure in the pressure profile

- Low- $\beta$  ET equilibrium
- Results are general



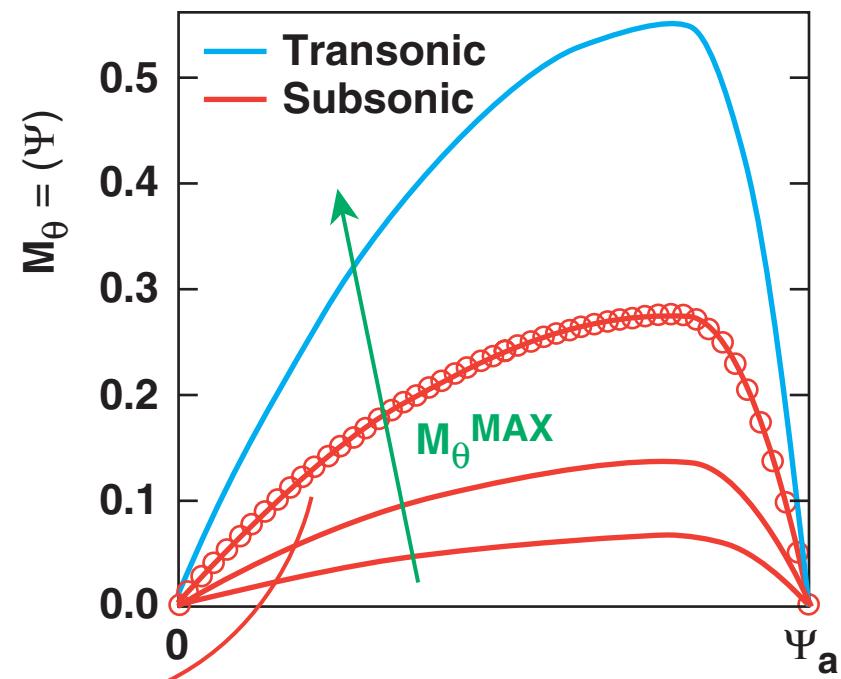
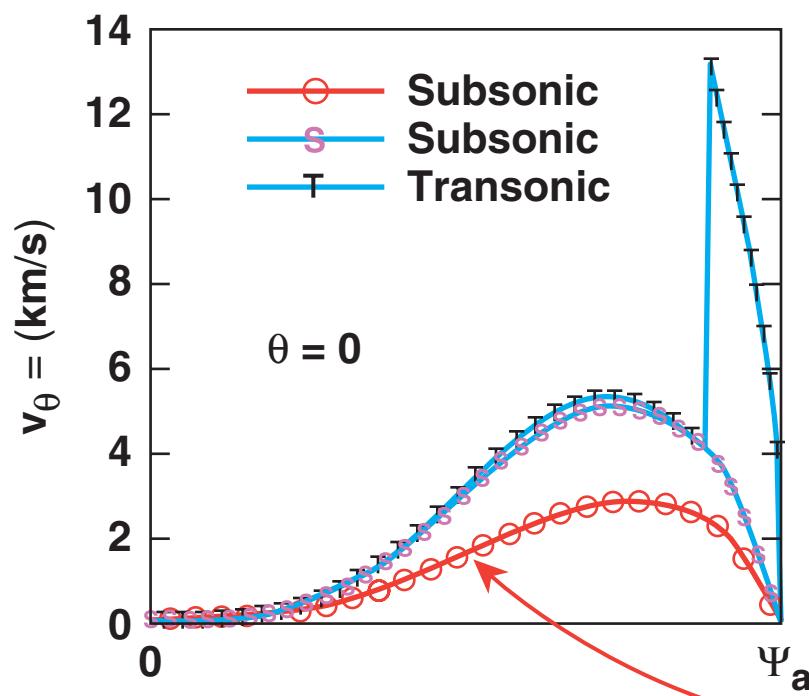
# A bifurcated equilibrium exists for a critical poloidal velocity

A bifurcated equilibrium exists when the free function  $M_\theta(\Psi)$  reaches the critical value.



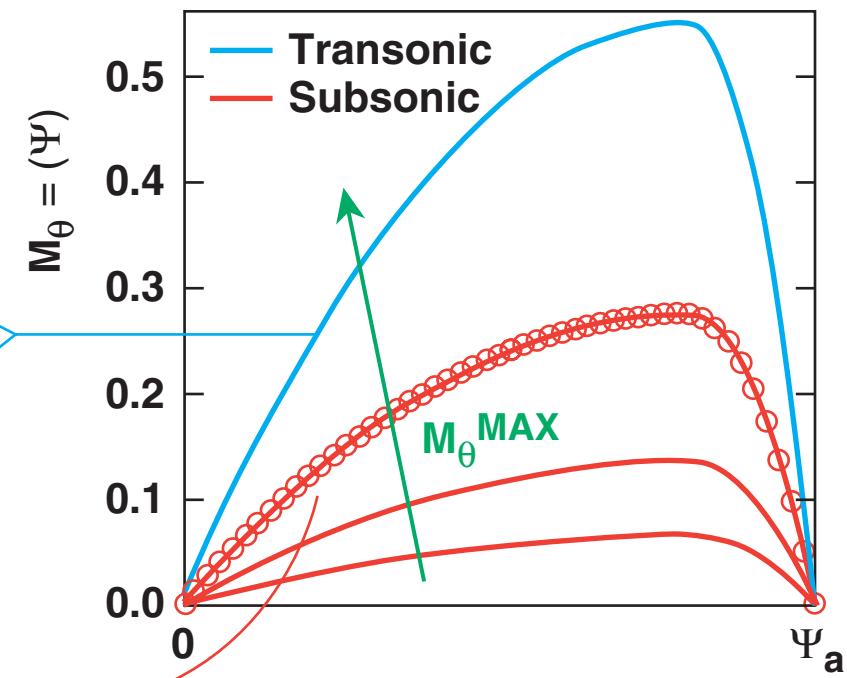
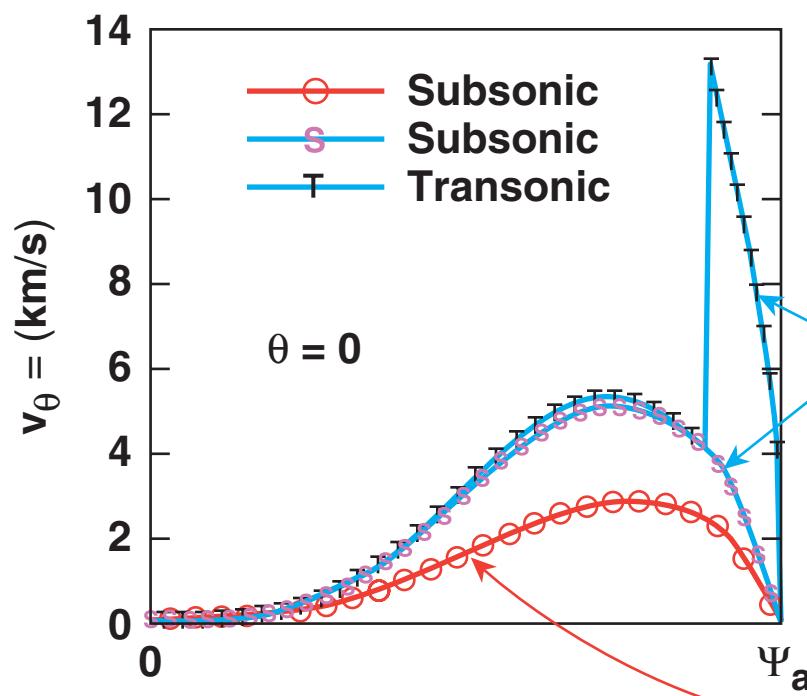
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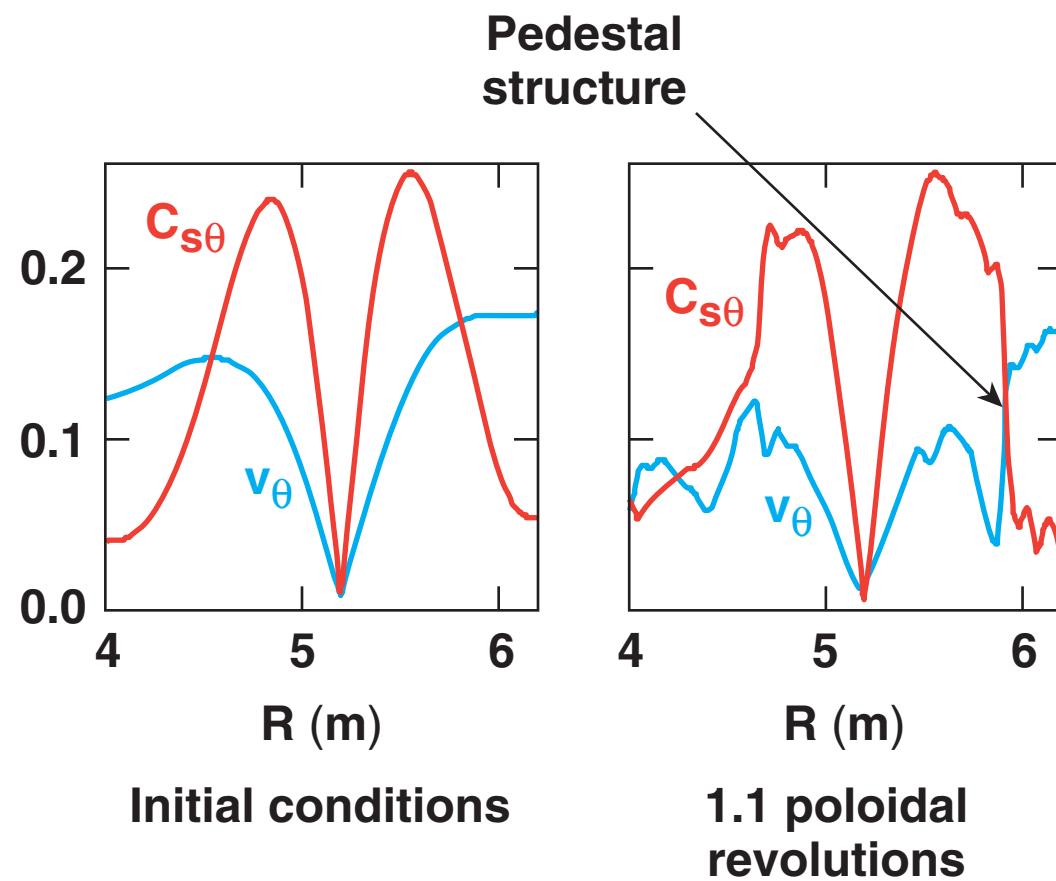
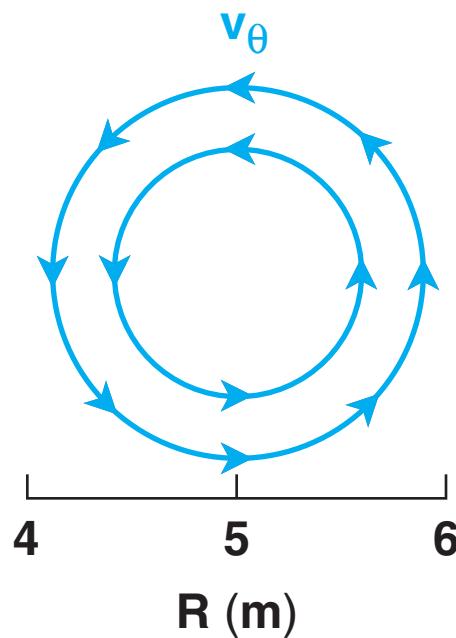
A bifurcated equilibrium exists when the free function  $M_\theta(\Psi)$  reaches the critical value.



# Initial value 2-D MHD simulations of transonic flow show the generation of discontinuities



- Poloidal sound speed (red) and poloidal velocity (blue) evolve to a discontinuous state.

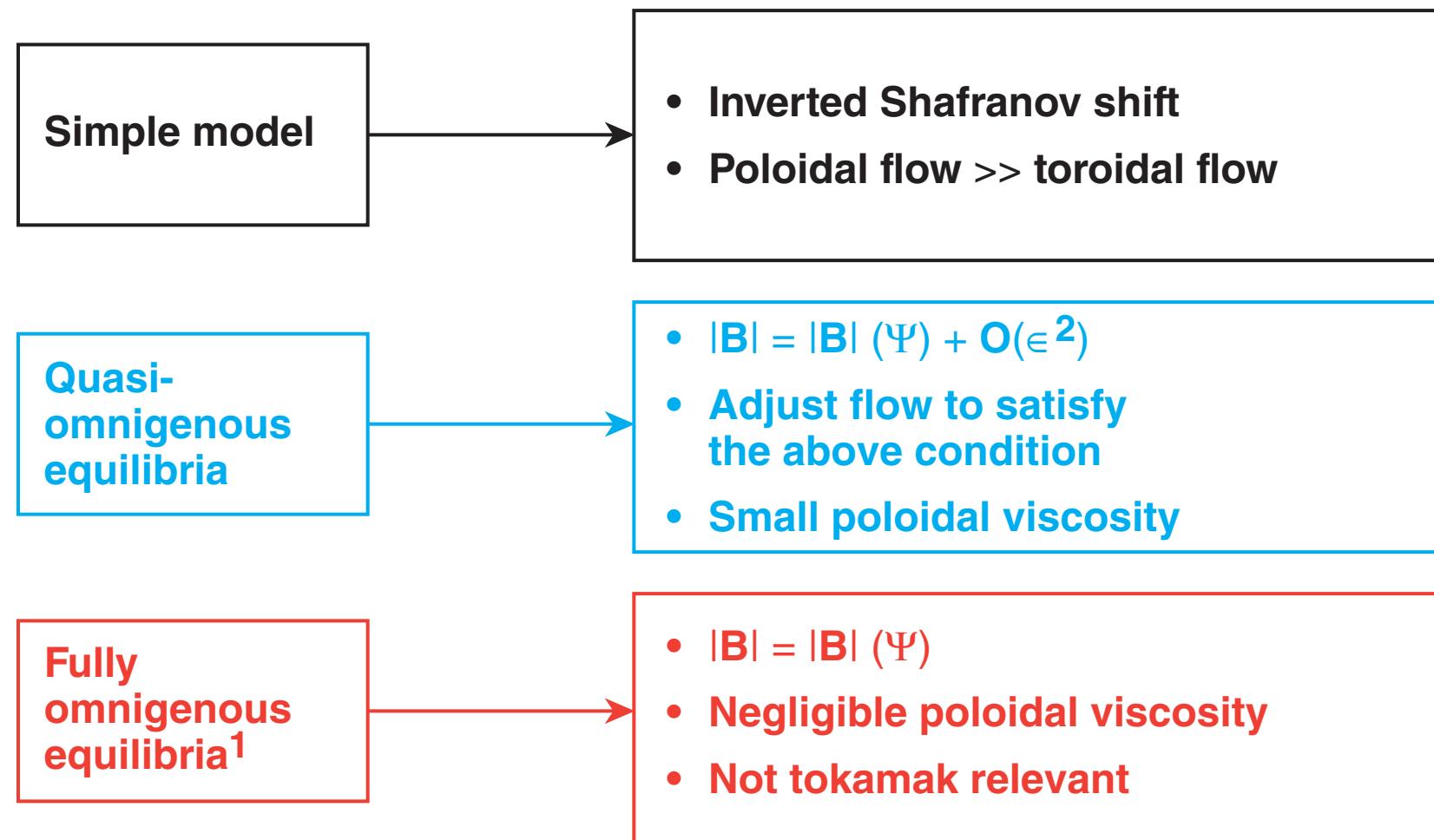


# **Equilibria with Super-Alfvénic Poloidal Flow**



$$\mathbf{v}_\theta \gtrapprox \mathbf{V}_{A\theta}$$

# Equilibria with super-Alfvénic poloidal flow (with respect to the poloidal Alfvén speed)

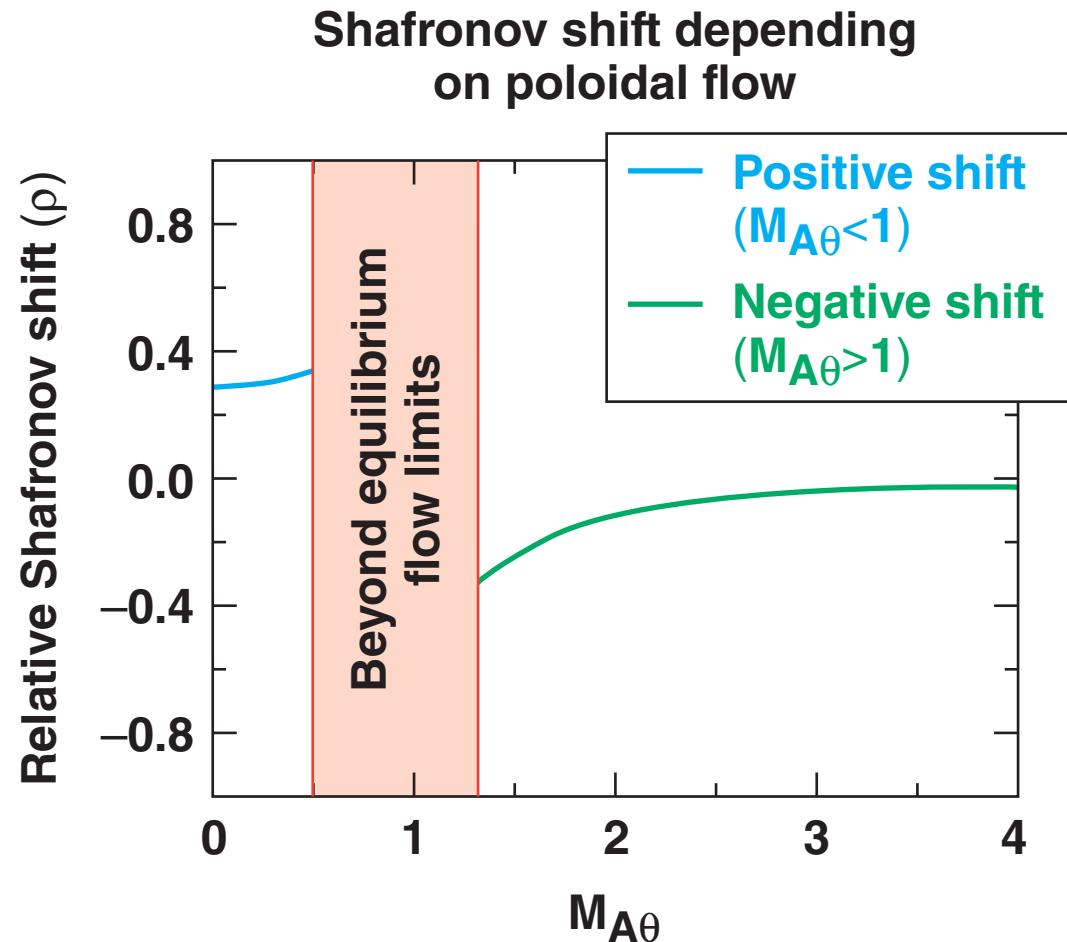


# An analytic model gives inverted Shafranov shift for Super-Alfvénic equilibria

**Assumptions:**

$$\epsilon \ll 1 \quad \beta \sim \epsilon$$
$$M_{A\theta} \approx \text{const.} \quad v_\phi \ll v_\theta$$

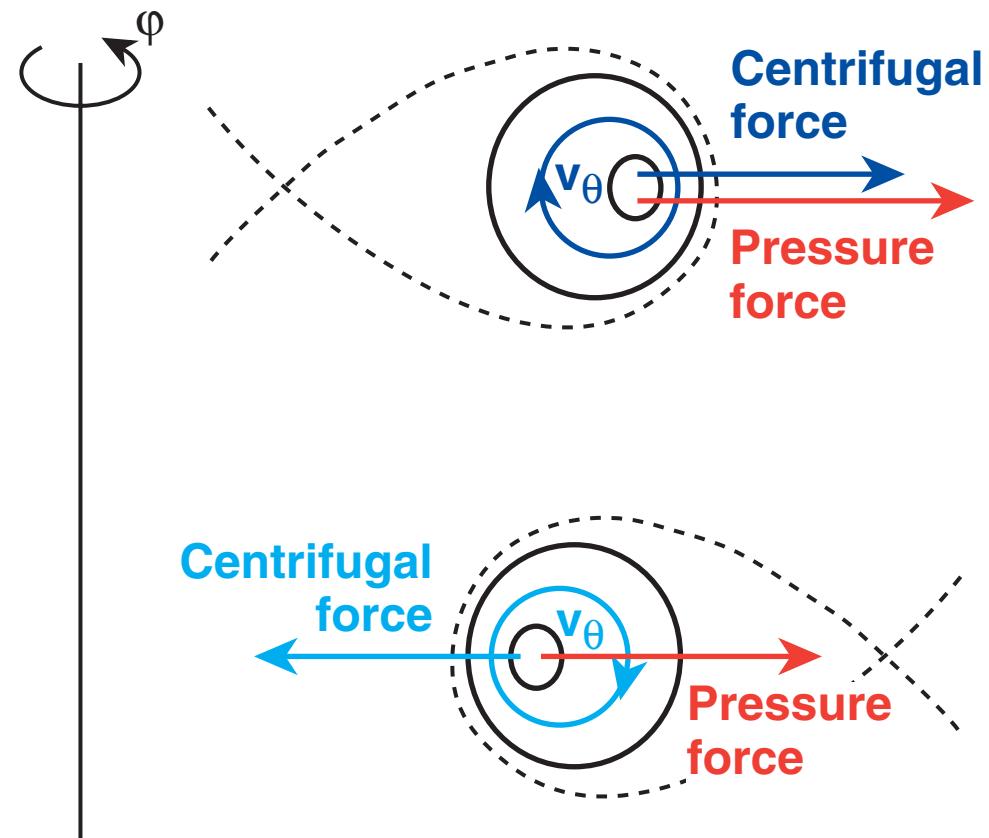
For a fixed  $\beta$  the Shafranov shift is computed as a function of  $M_{A\theta}$ .



J. P. Freidberg, *Ideal Magnetohydrodynamics* (Plenum Press, New York, 1987), p. 138.  
F. A. Haas, Phys. Fluids 15, 141 (1972).

# For inward shifted equilibria the pressure forces are balanced by the centrifugal forces

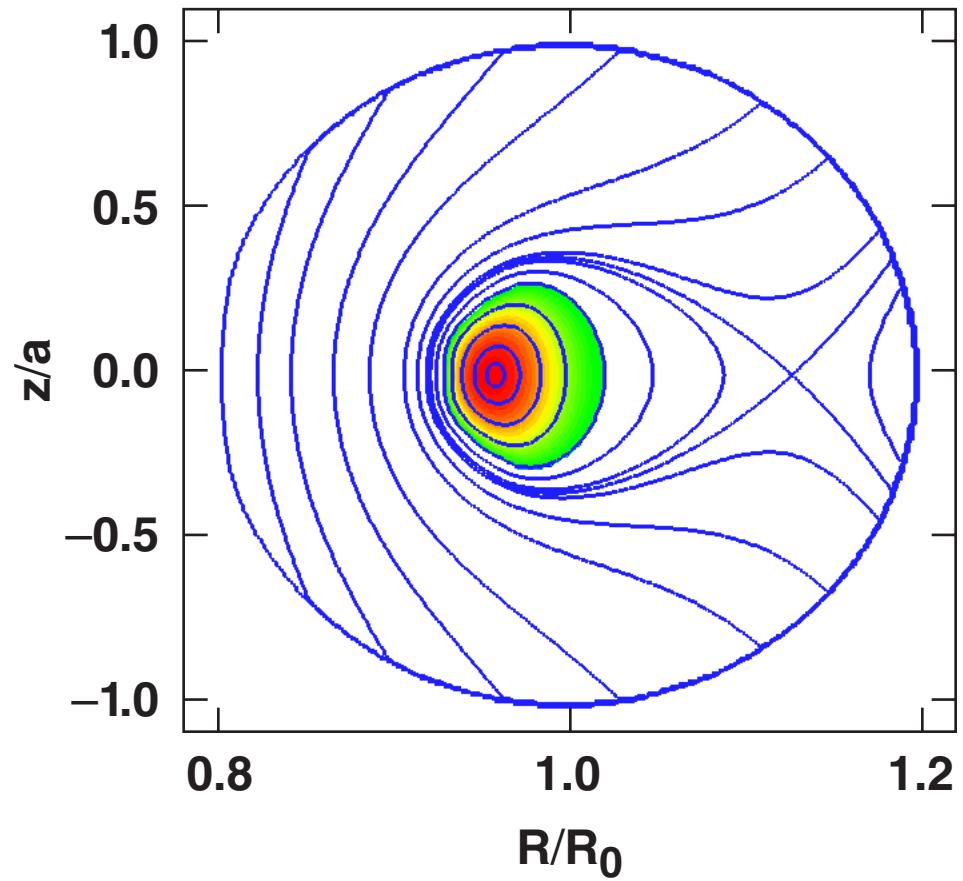
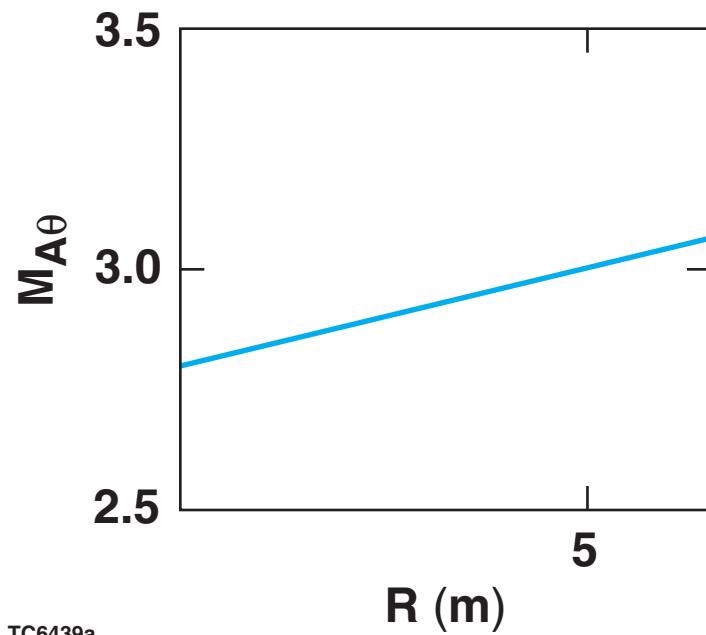
- Pressure forces cause outward Shafranov shift.
- Poloidal flow produces additional force.
- If the Shafranov shift is outward, pressure and centrifugal forces are aligned.
- If the Shafranov shift is inward, pressure and centrifugal forces are opposite.



# FLOW confirms the existence of inward shifted equilibria for $M_{A\theta} > 1$

## Numerical example:

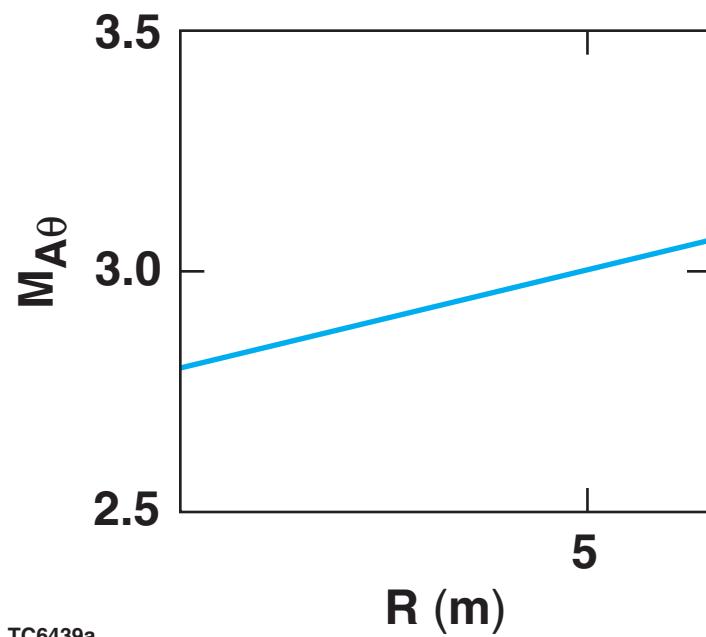
- $M_\varphi (\Psi) = 0$
- Flat density
- Peaked pressure



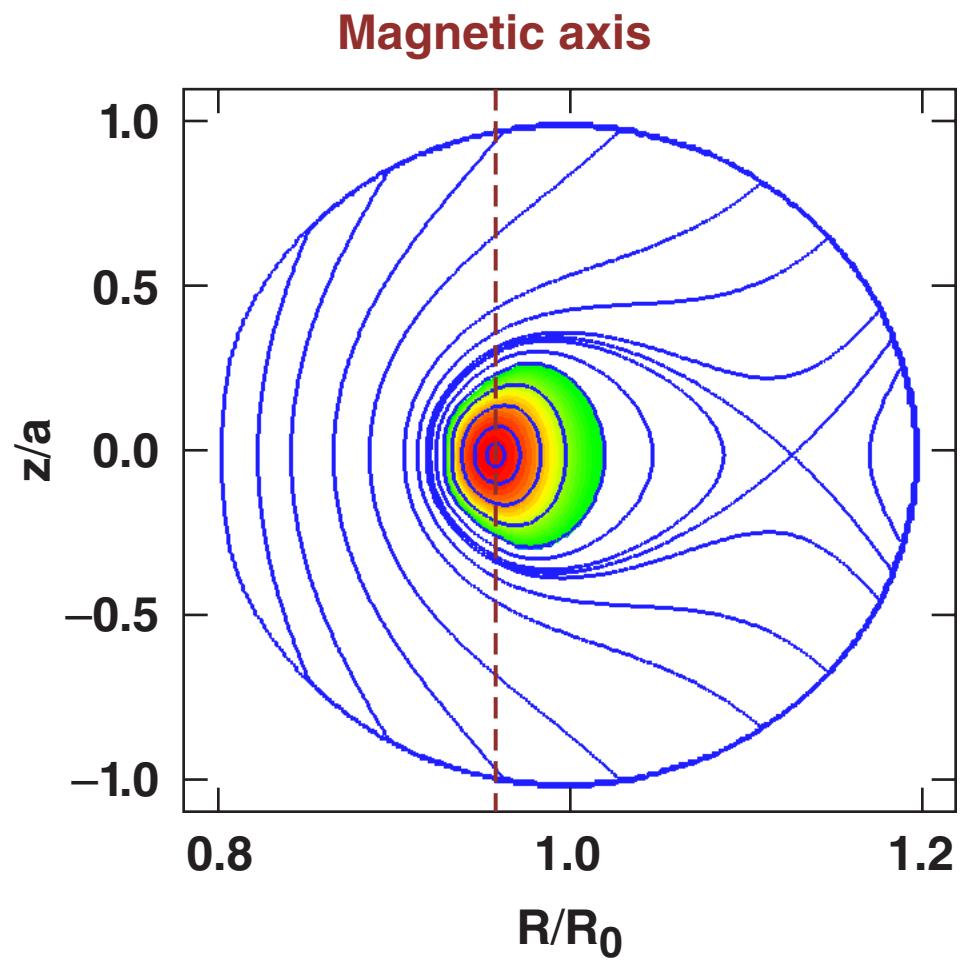
# FLOW confirms the existence of inward shifted equilibria for $M_{A\theta} > 1$

## Numerical example:

- $M_\varphi (\Psi) = 0$
- Flat density
- Peaked pressure



TC6439a



# Quasi-omnigenous [ $|B| = |B|(\Psi) + O(\epsilon^2)$ ] equilibria with fast poloidal flow



- The solution uses an  $\epsilon$  expansion, assuming the ordering:

$$\underbrace{B_\theta \sim \epsilon B_\phi}_{\downarrow} \quad \left( M_{A\theta}^2 - 1 \right) \sim 1 \quad \beta \sim \epsilon$$
$$|B|^2 = B_\phi^2 + B_\theta^2 \approx B_\phi^2 + O(\epsilon^2)$$

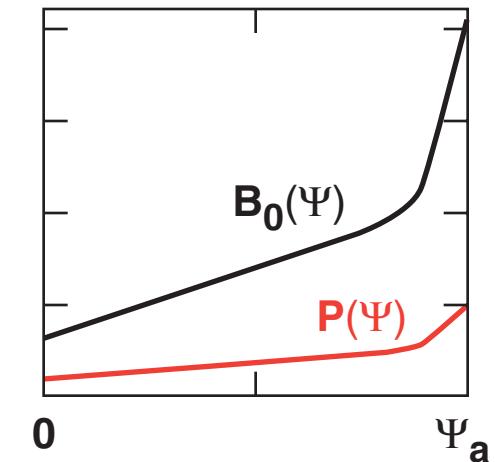
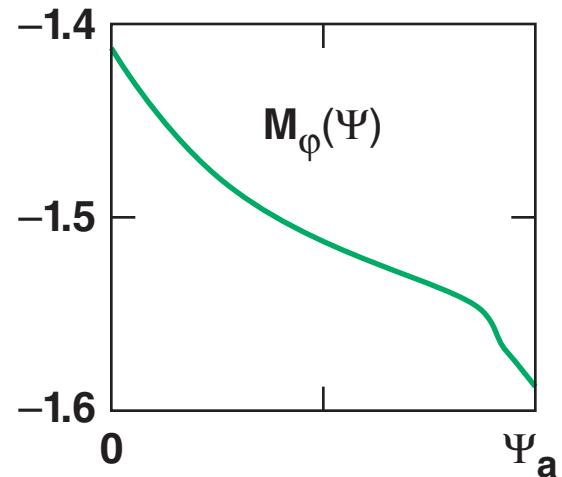
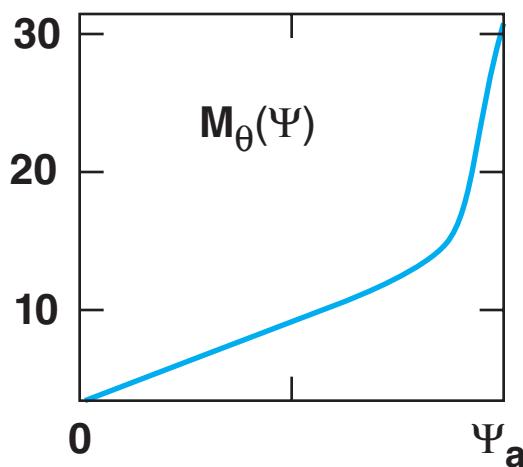
- $B_\phi$  is imposed to be a function of  $\Psi$  only up to  $O(\epsilon^2)$  corrections.

# Fast poloidal flows are used to make the magnetic field quasi-omnigenous

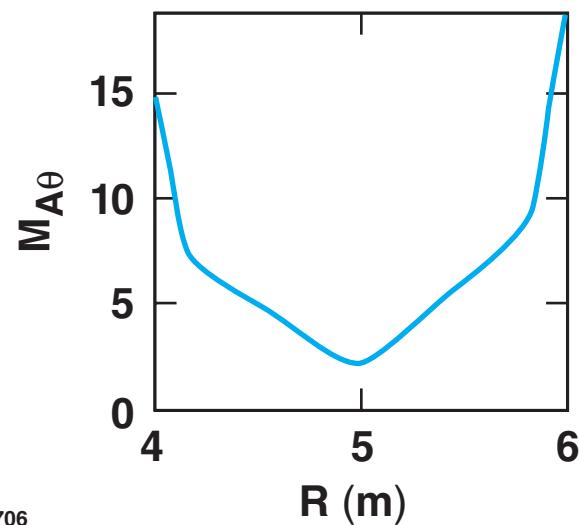
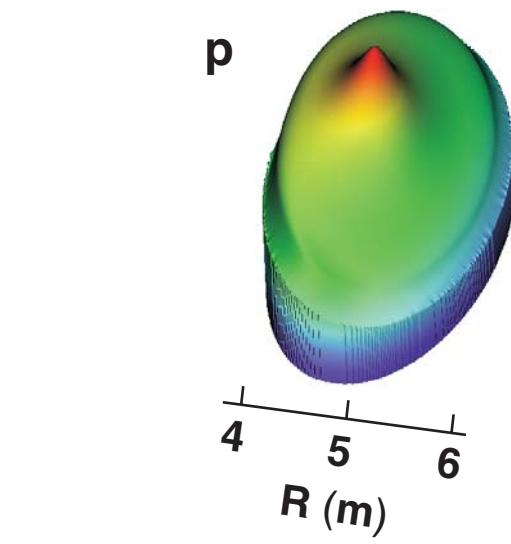
In the static case  $B_\phi = \frac{F(\Psi)}{R} \longrightarrow B_\phi = B_\phi(\Psi) + O(\epsilon).$

With flow:

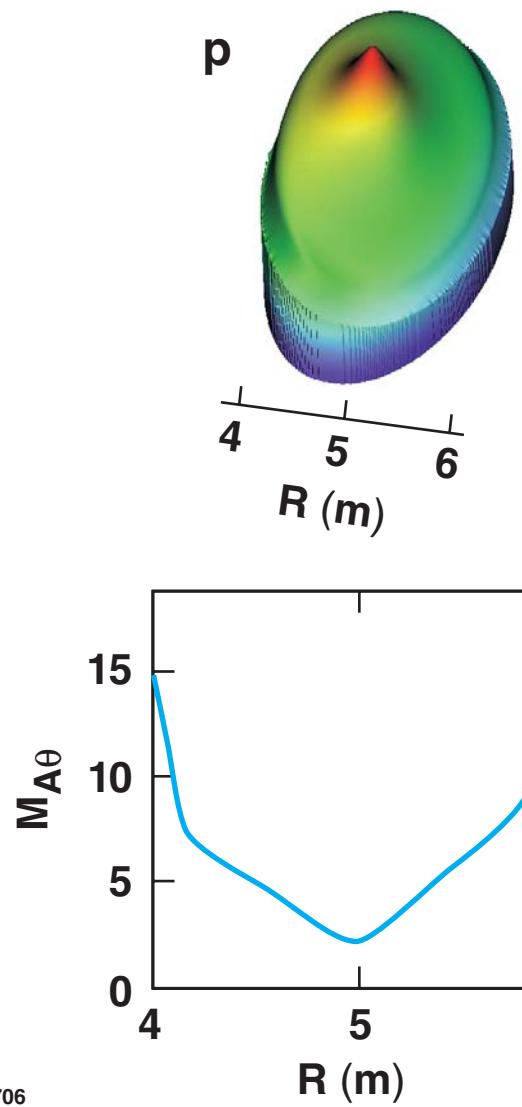
$$B_\phi = \frac{1}{1 - M_{A\theta}^2} \left[ \frac{F(\Psi)}{R} + RM_{A\theta}\sqrt{\rho}\Omega(\Psi) \right] \longrightarrow \text{adjust flow to impose } B_\phi = B_\phi(\Psi) + O(\epsilon^2)$$



# FLOW is used to compute the quasi-omnigenous equilibria of arbitrary shape

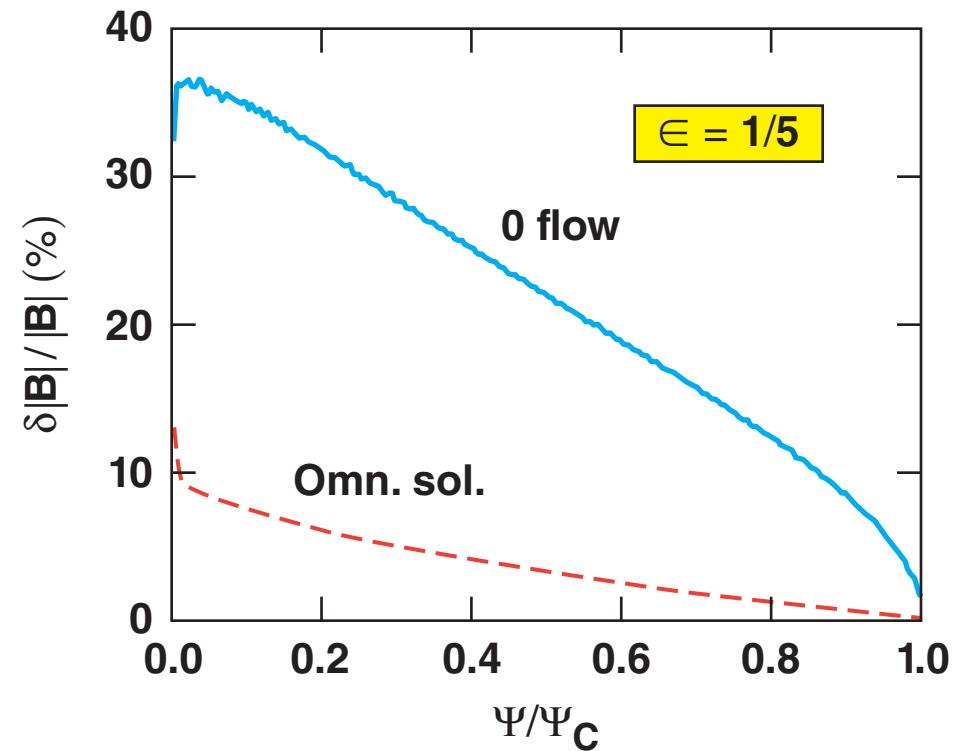


# FLOW is used to compute the quasi-omnigenous equilibria of arbitrary shape



TC6706

Magnetic field relative variations are computed for a quasi-omnigenous and a static equilibrium.



## Conclusions

# Tokamak equilibria with flow are very different from static equilibria



- Equilibria with macroscopic flows have been studied analytically. The numerical results of the code FLOW confirm the results of theory.
- Equilibria with poloidal flow in the range of the poloidal sound speed  $C_s B_\theta / B$  develop a pedestal structure due to the transition from subsonic to supersonic regime.
- Equilibria with poloidal flow in the super-Alfvénic regime show inverted Shafranov shift. The existence of a new class of quasi-omnigenous equilibria in such regime has been discussed.