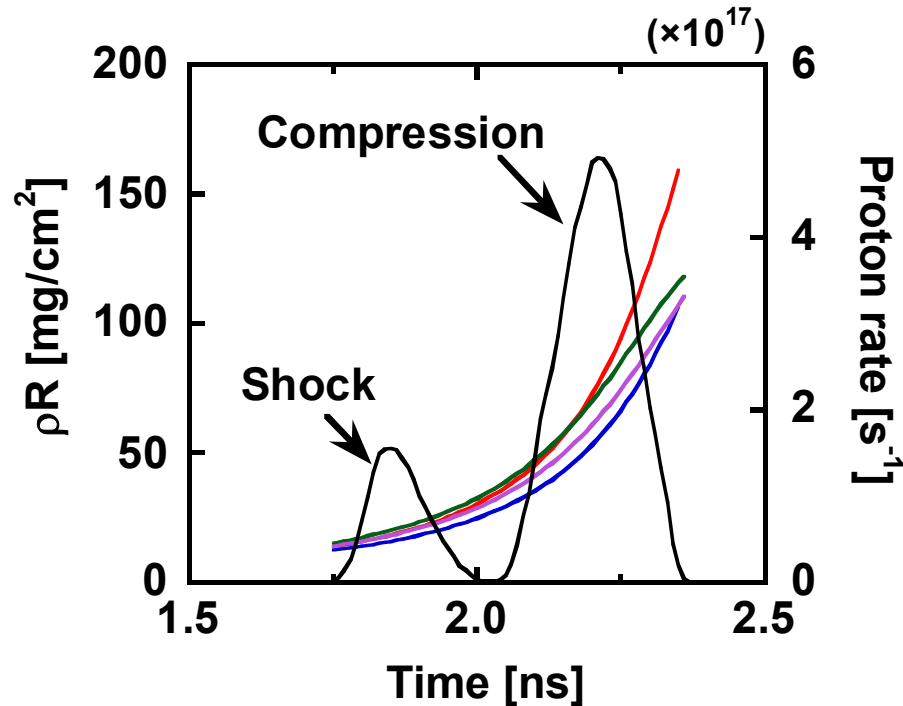
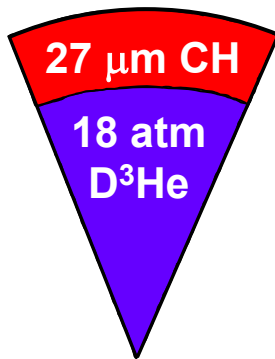


Measurements of shock timing and ρR evolution of D^3He implosions at OMEGA



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45th Annual Meeting of the
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Summary

Shock timing and ρR evolution of D^3He implosions have been measured at OMEGA

- D^3He burn history contains a shock component in addition to a compression history similar to that of DD neutrons.
- $T_i(t)$, shock time and shock-burn duration have been obtained and compared with 1-D calculations.
- Low-mode ρR asymmetries at shock time are amplified and mirrored at bang time, and correlated to laser drive asymmetry (for a large imposed $\ell = 1$).
- We are looking into 3He -seeded cryogenic D_2 implosions.

Related work

Related talks and posters at this conference:

F.J. Marshall et al., **CO2.005**
R. Epstein et al., **CO2.008**
R. D Petrasso et al., **CO2.009**
F. H. Séguin et al., **CO2.011**
C. K. Li et al., **CO2.012**
R. Rygg et al., **CO2.013**
J. DeCiantis et al., **CO2.015**
V. Yu. Glebov et al., **UP1.007**
D. Wilson et al., **B12.04**

Recent related papers:

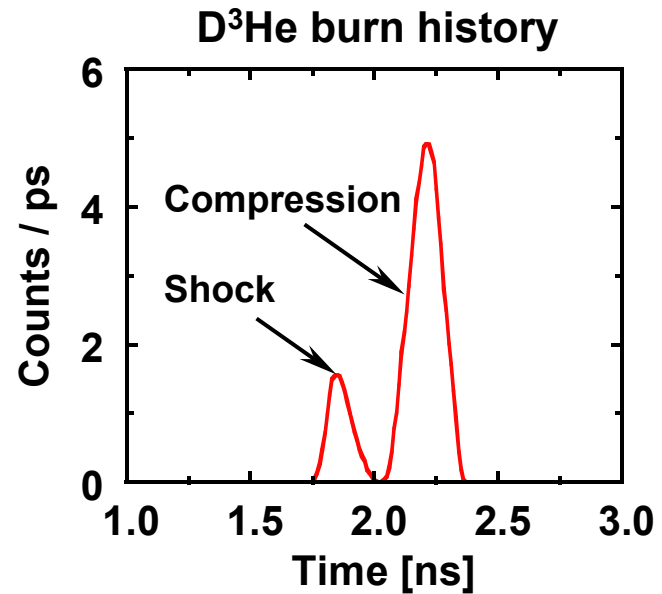
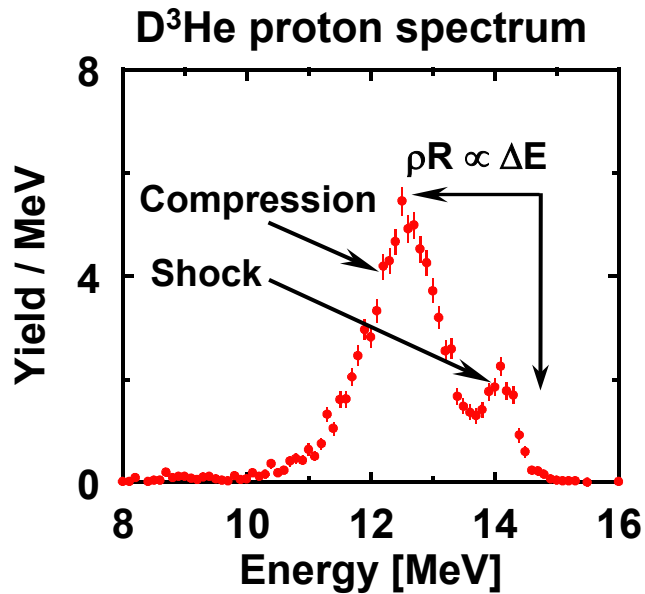
R. D. Petrasso et al., **Phys. Rev. Letters 90 (2003) 095002.**
V. A. Smalyuk et al., **Phys. Rev. Letters 90 (2003) 135002.**
C. K. Li et al., **submitted to Phys. Rev. Letters.**



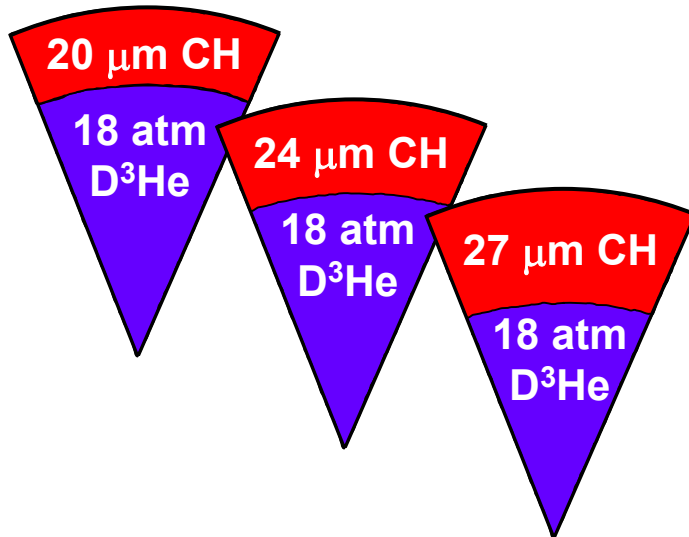
Outline

- Principle of measuring ρR evolution and $D^3\text{He}$ burn history
- Experiments
- Effects causing time dispersion in measured $D^3\text{He}$ burn history data
- Analysis method
- Results
- ^3He -seeded cryogenic D_2 implosions

$\rho R(t)$ can be inferred from $D^3\text{He}$ proton spectrum and $D^3\text{He}$ burn history



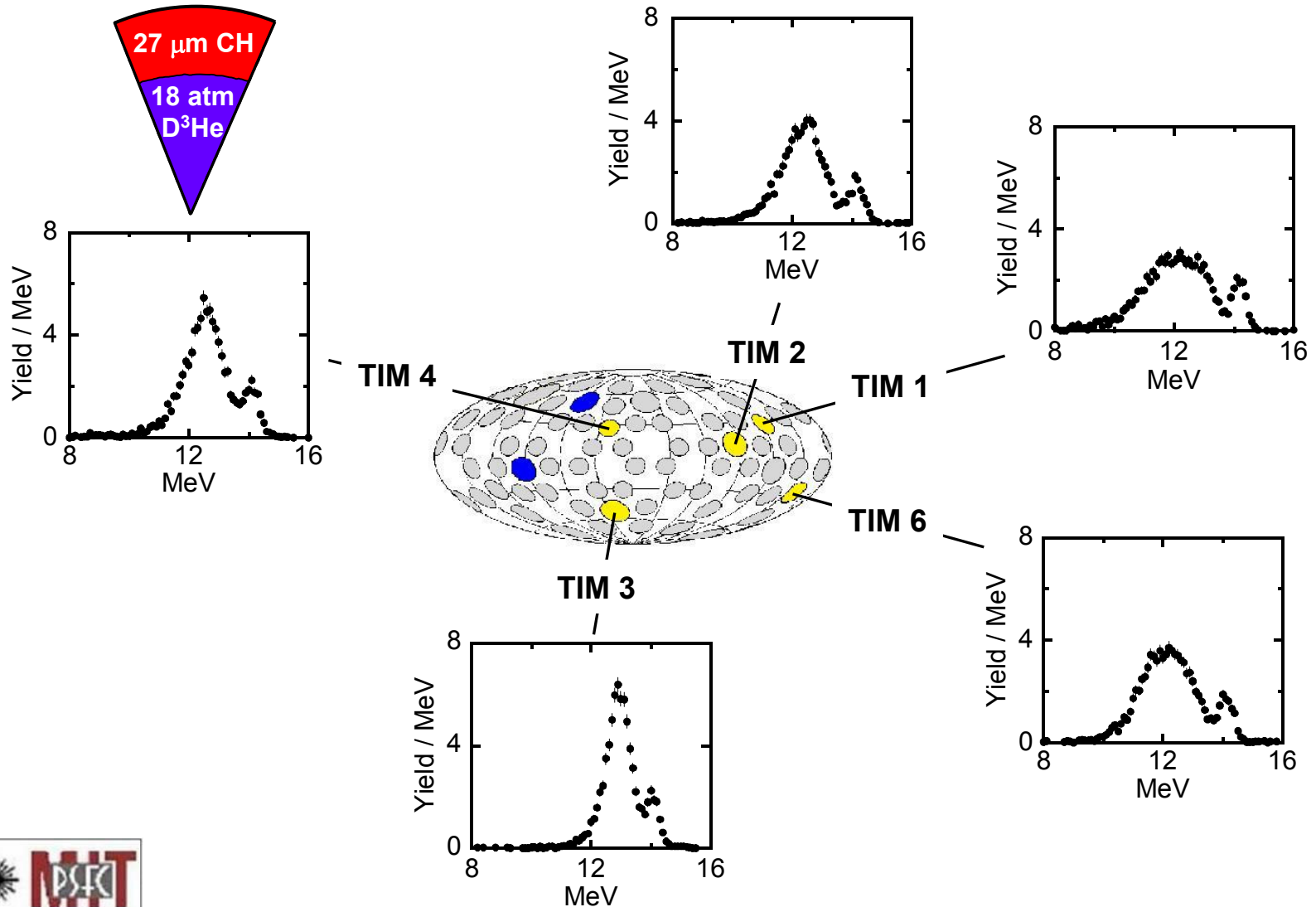
Three types of capsules were imploded



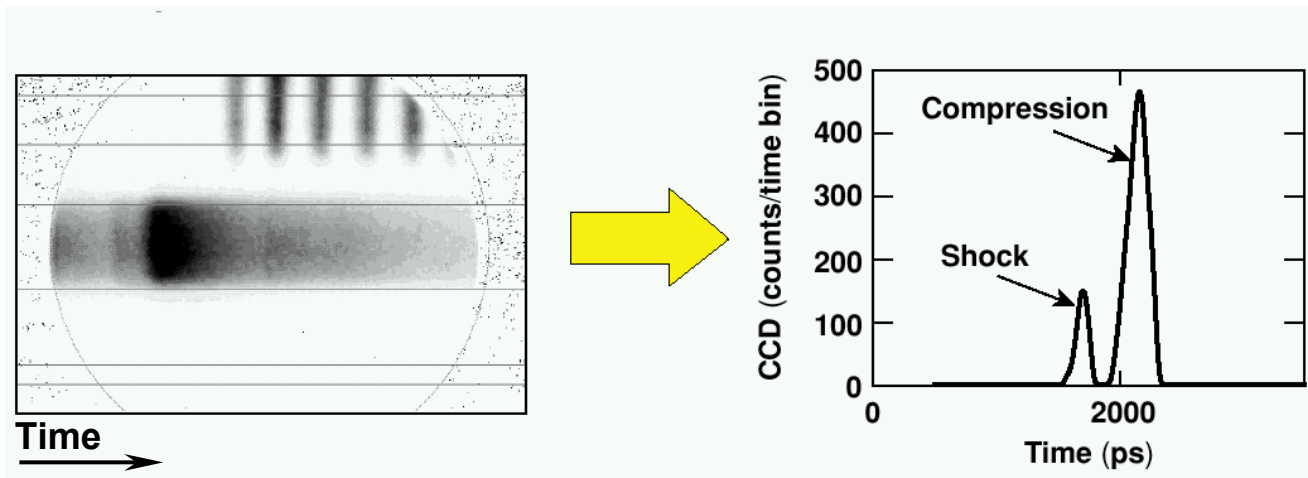
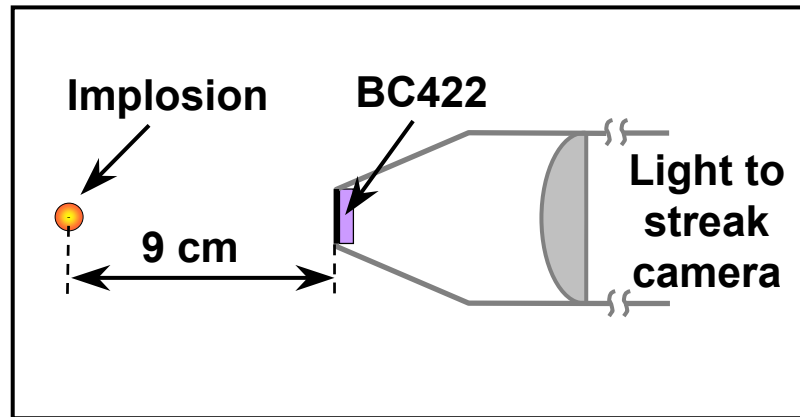
- Shock time
- Shock burn duration
- Nature of compression burn
- T_i evolution
- Evolution of ρR
- Evolution ρR asymmetries

60 laser beams
23 kJ energy
1 ns square laser pulse
1 THz, 2-D SSD+PS

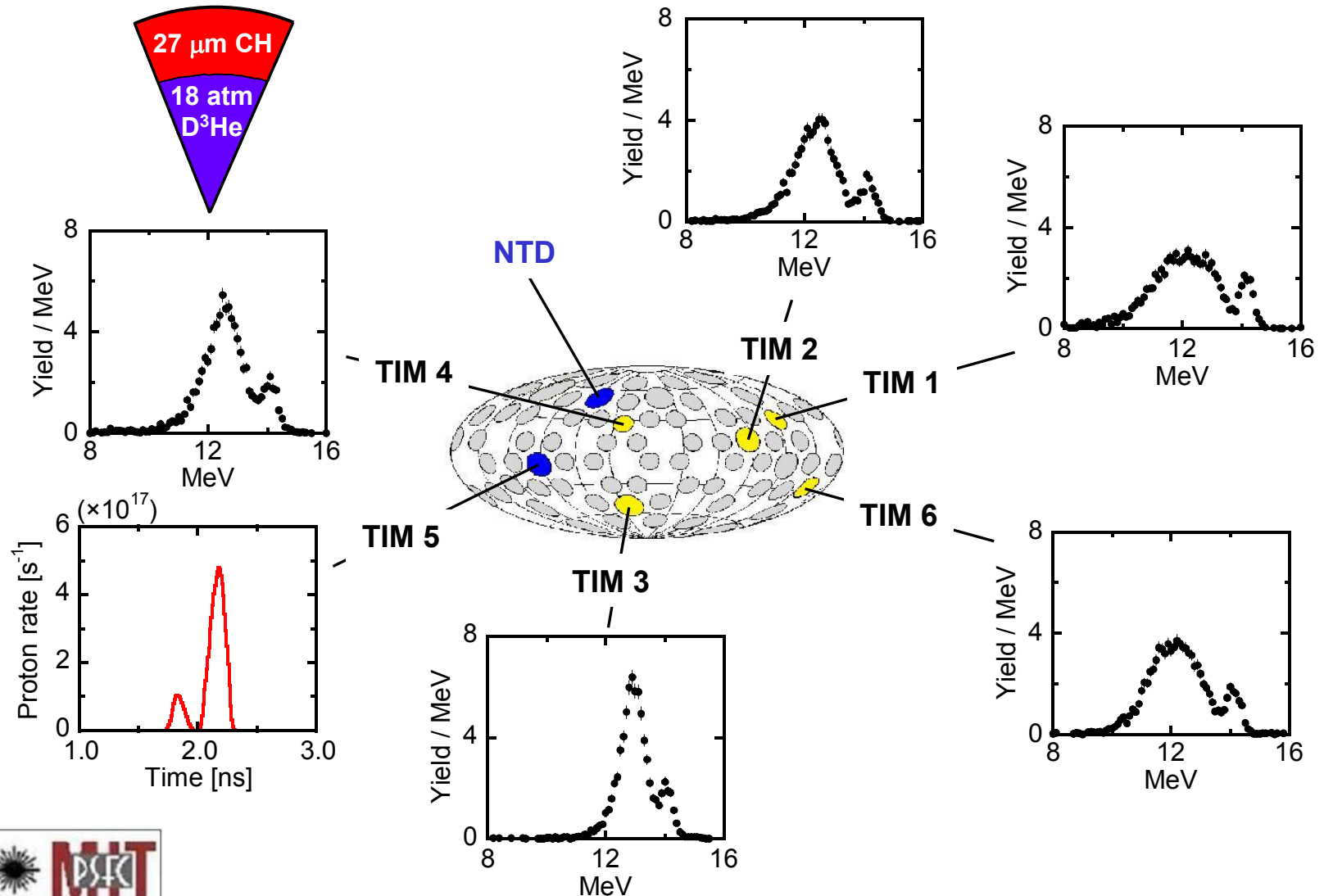
D^3He proton spectra were simultaneously measured from different directions



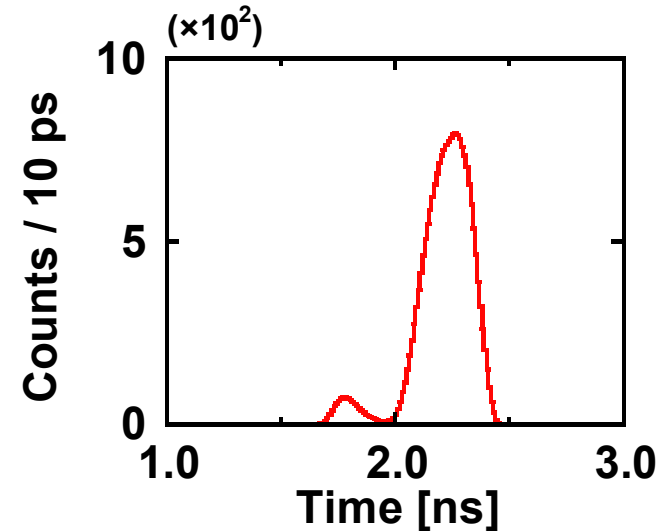
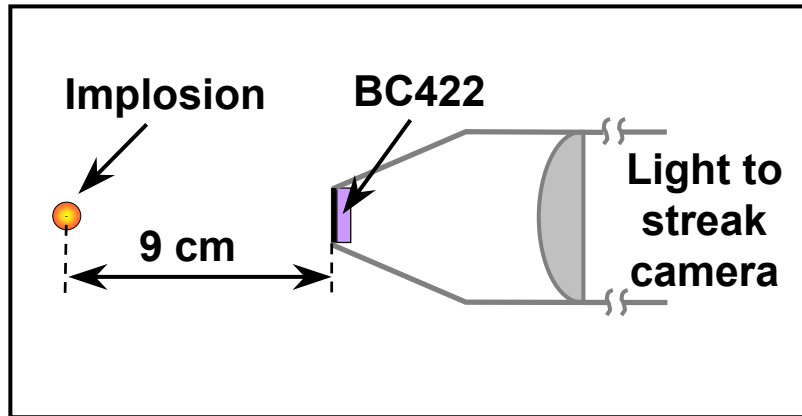
A proton temporal diagnostic (PTD) was implemented for measurements of D³He burn history



D³He burn history and D³He proton spectra were simultaneously measured



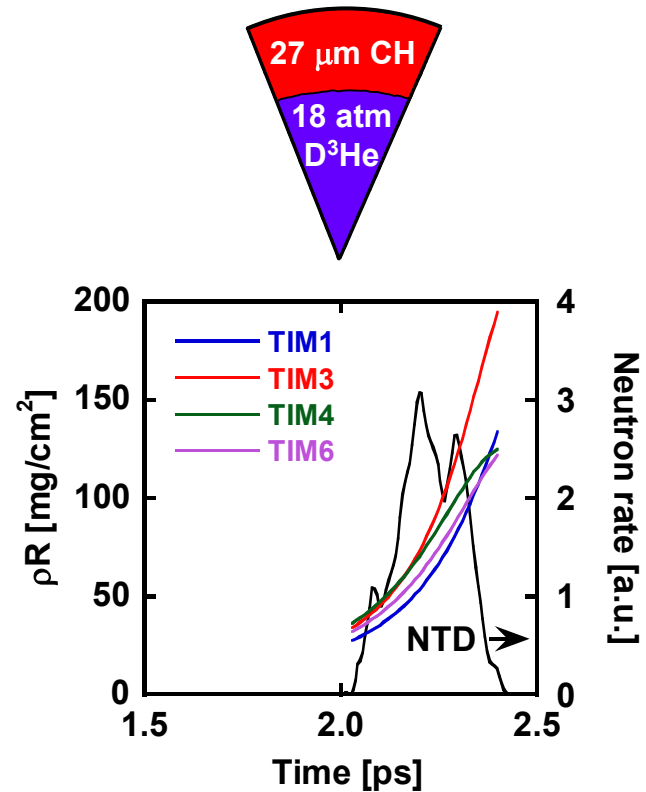
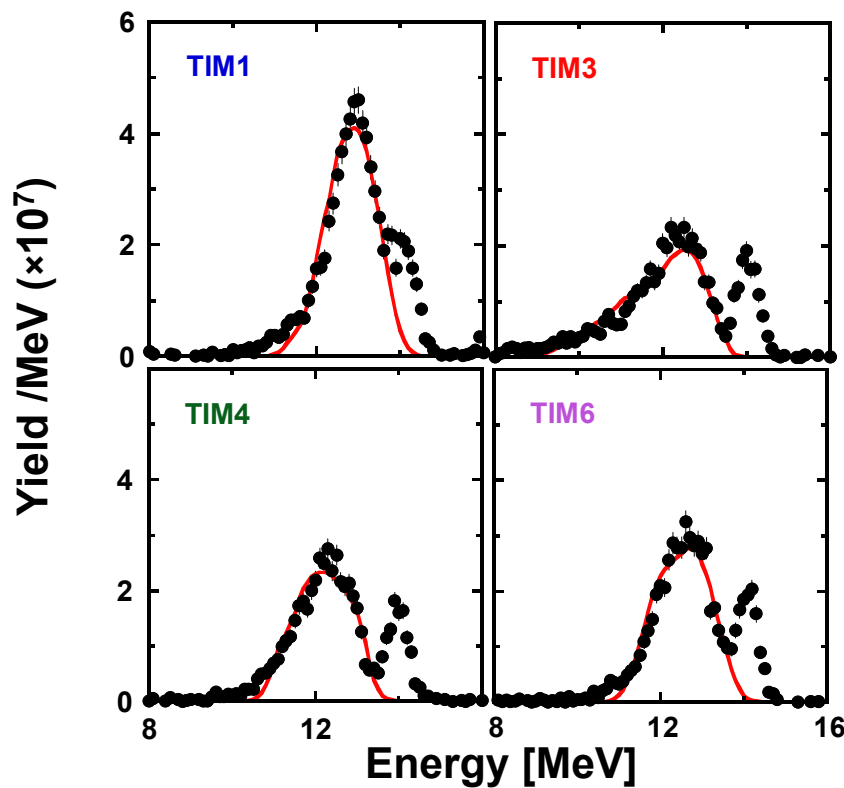
PTD data must be corrected for time dispersion



Effects causing time dispersion:

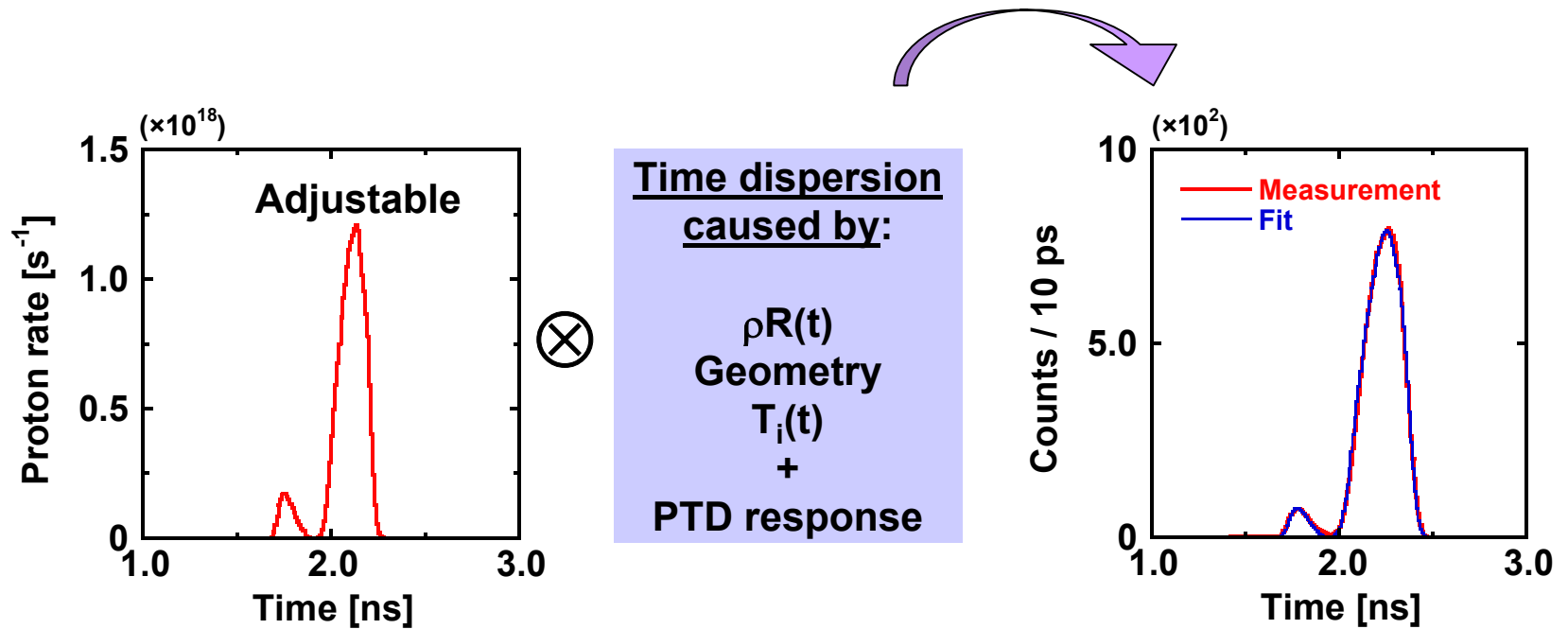
- ρR evolution: Needs to be determined.
- Source and shell geometry: From Proton Core Imaging data and X-ray imaging data.
- Doppler broadening from $T_i(t)$: From measurements.
- PTD response: From Monte-Carlo simulations.

Using DD burn history, $\rho R(t)$ was initially determined from a fit to measured D^3He -proton spectrum

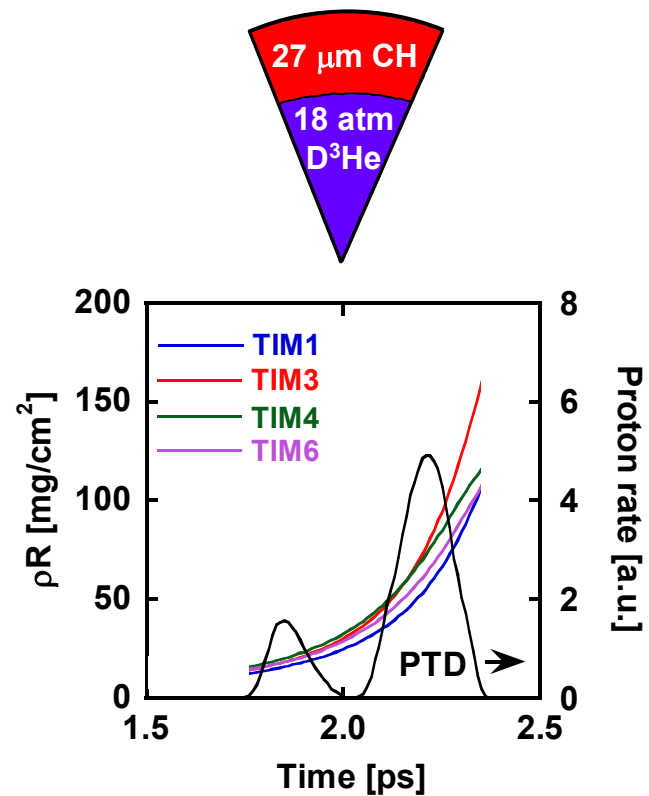
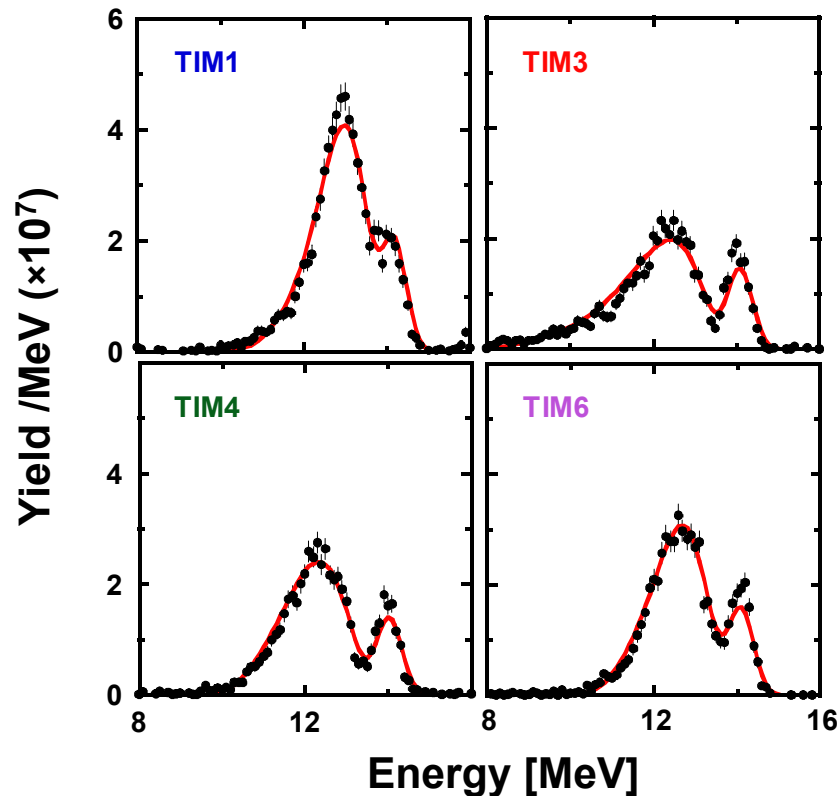


A Lorentzian function was used as a $\rho R(t)$ function in the fitting procedure.

A convolution of D^3He burn history and components causing time dispersion is fitted to measured PTD data

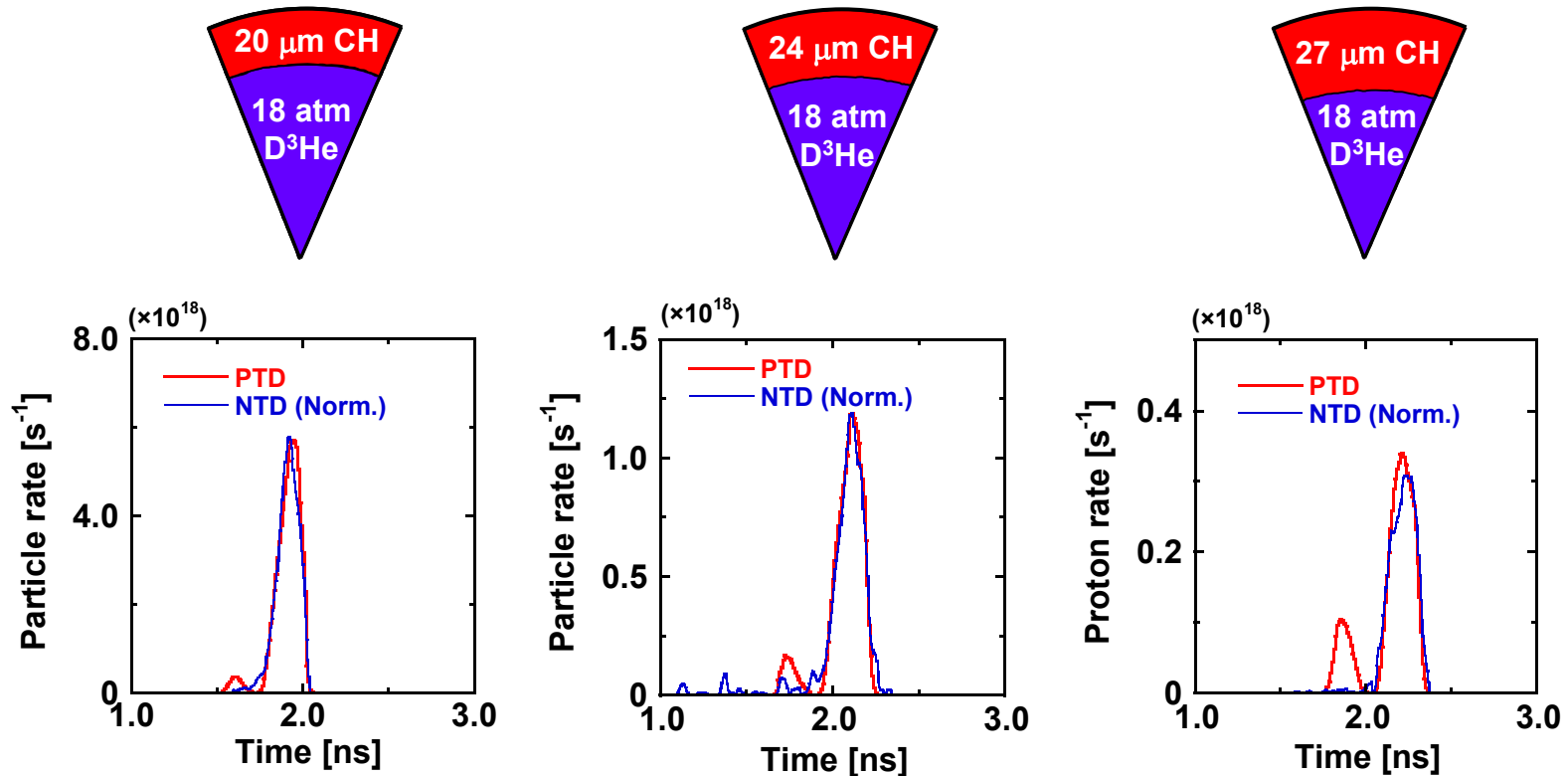


Using unfolded D^3He burn history, $\rho R(t)$ was finally determined from a fit to measured D^3He -proton spectrum

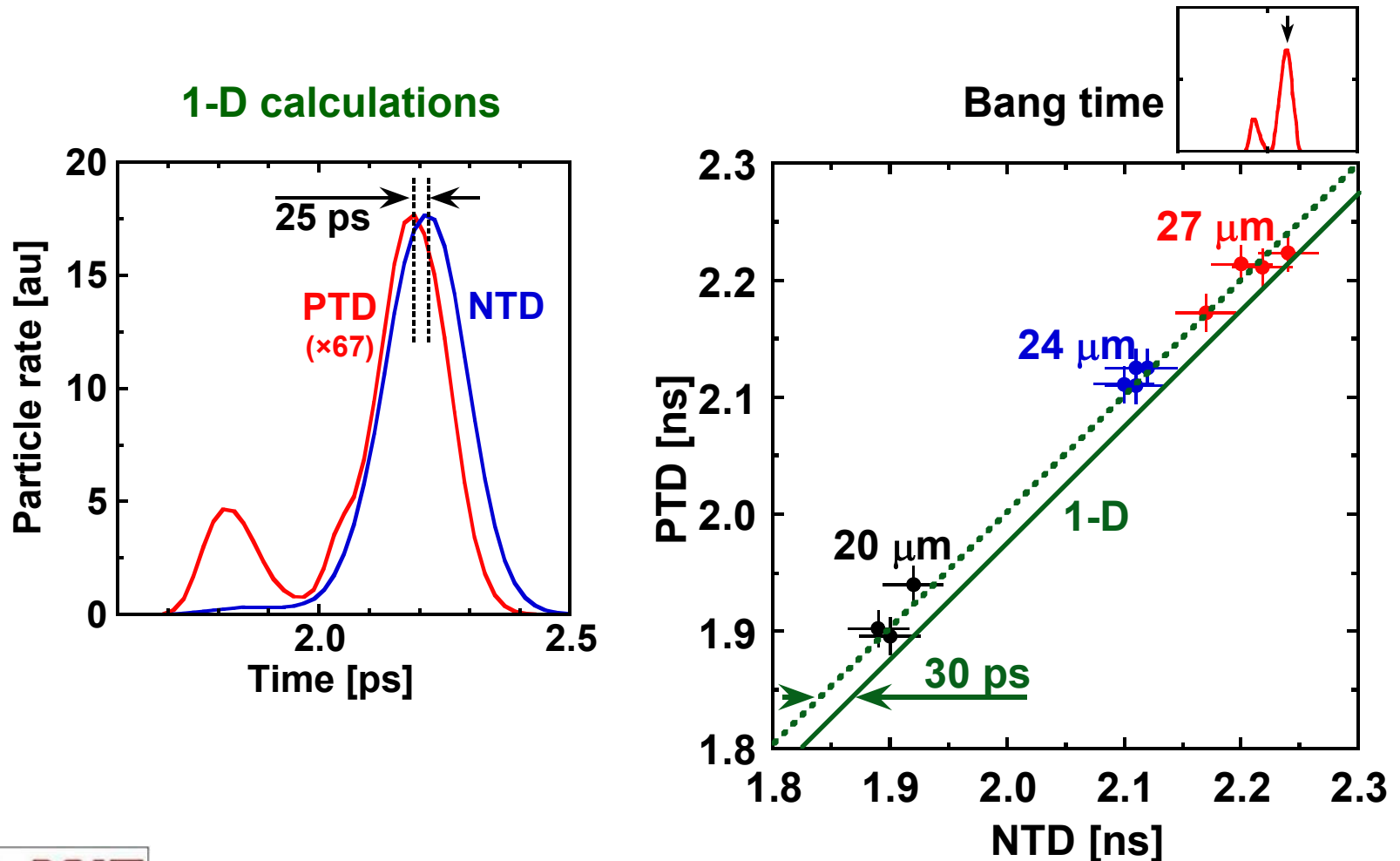


A Lorentzian function was used as a $\rho R(t)$ function in the fitting procedure.

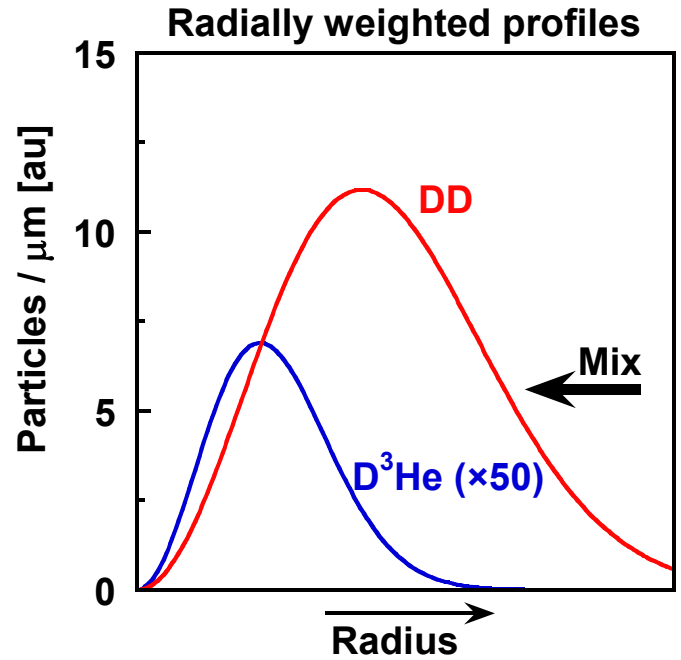
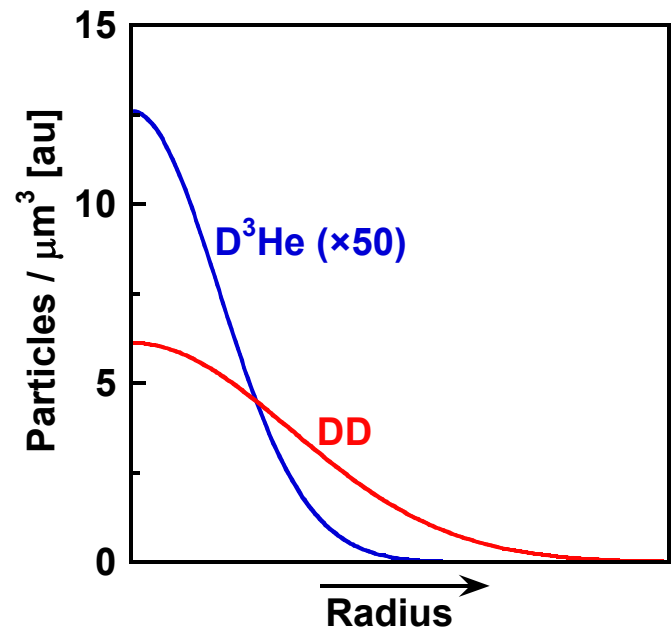
D³He burn history contains a shock component in addition to a compression history similar to that of DD neutrons



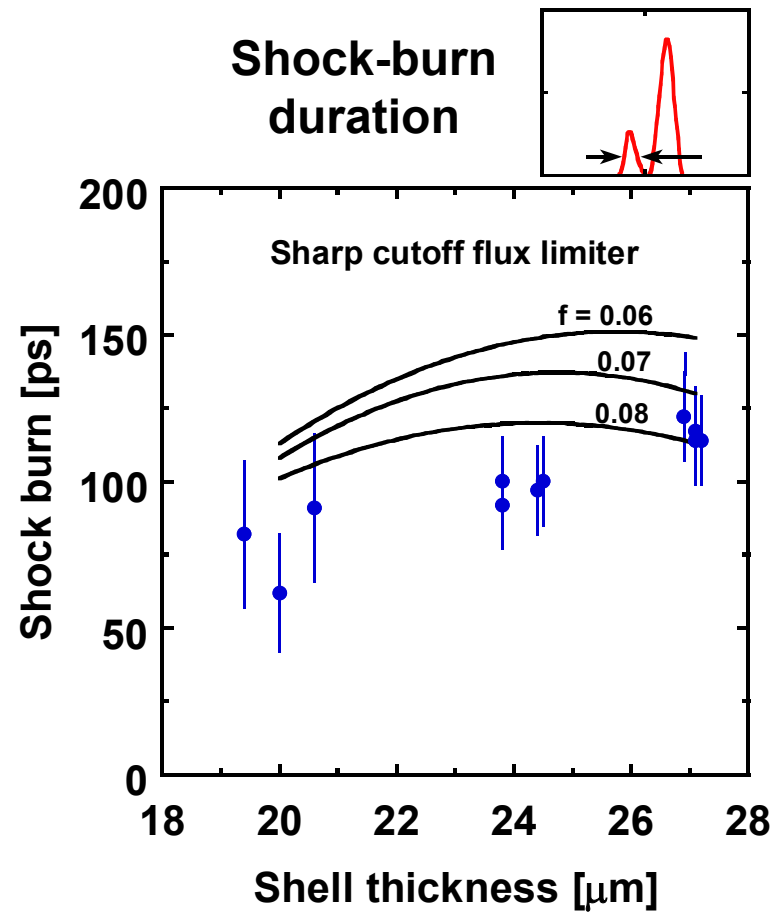
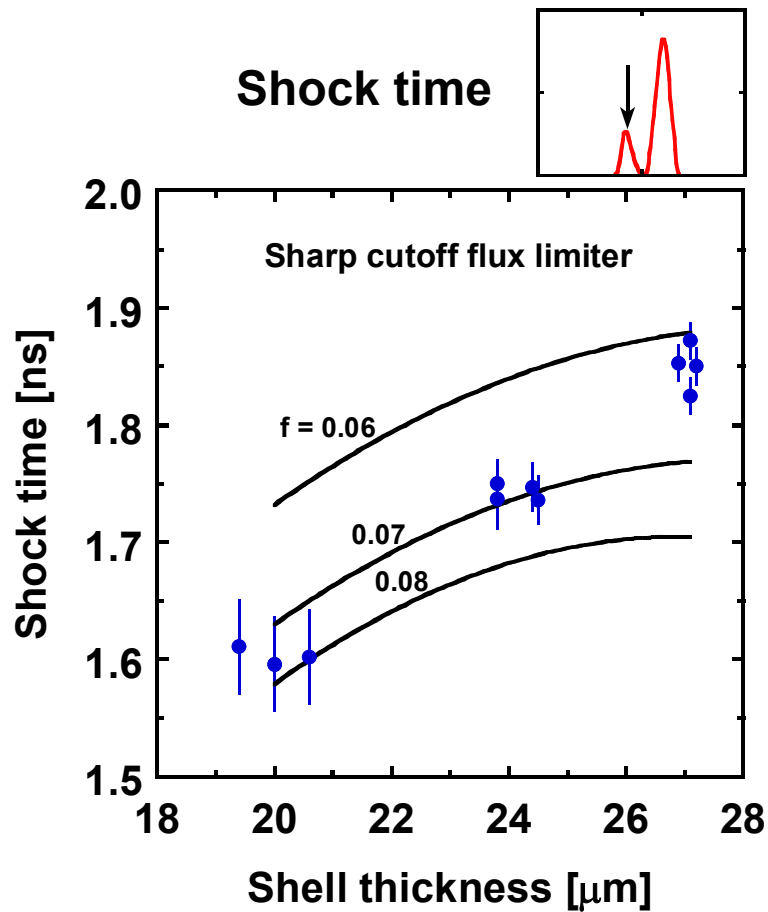
By comparing measured D^3He and DD bang times to 1-D calculations effects of mix can be addressed



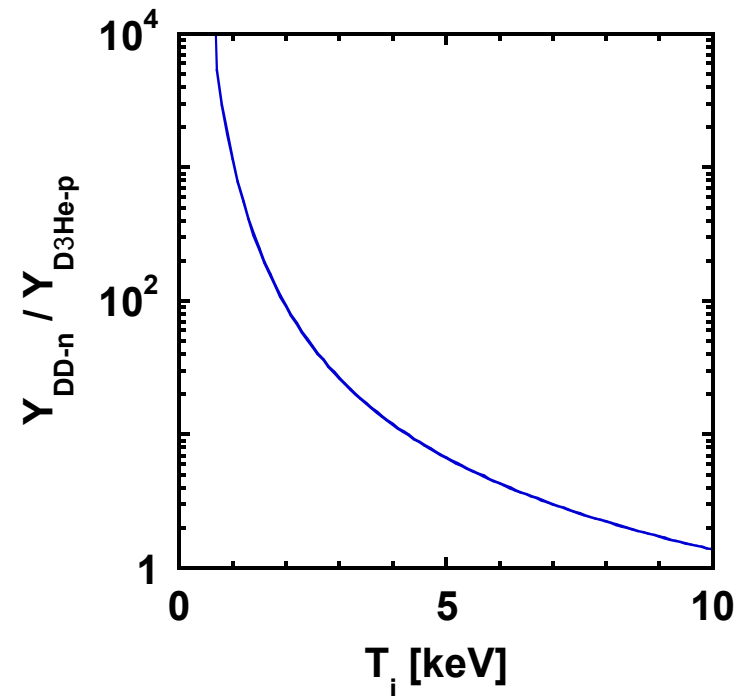
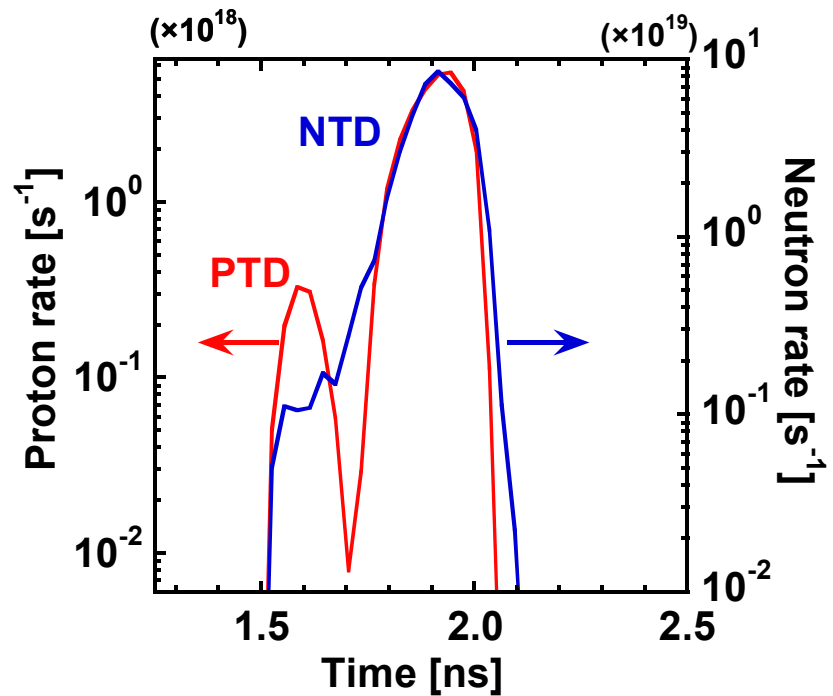
Due to a broader burn profile, DD burn history is more sensitive to mix than D³He burn history



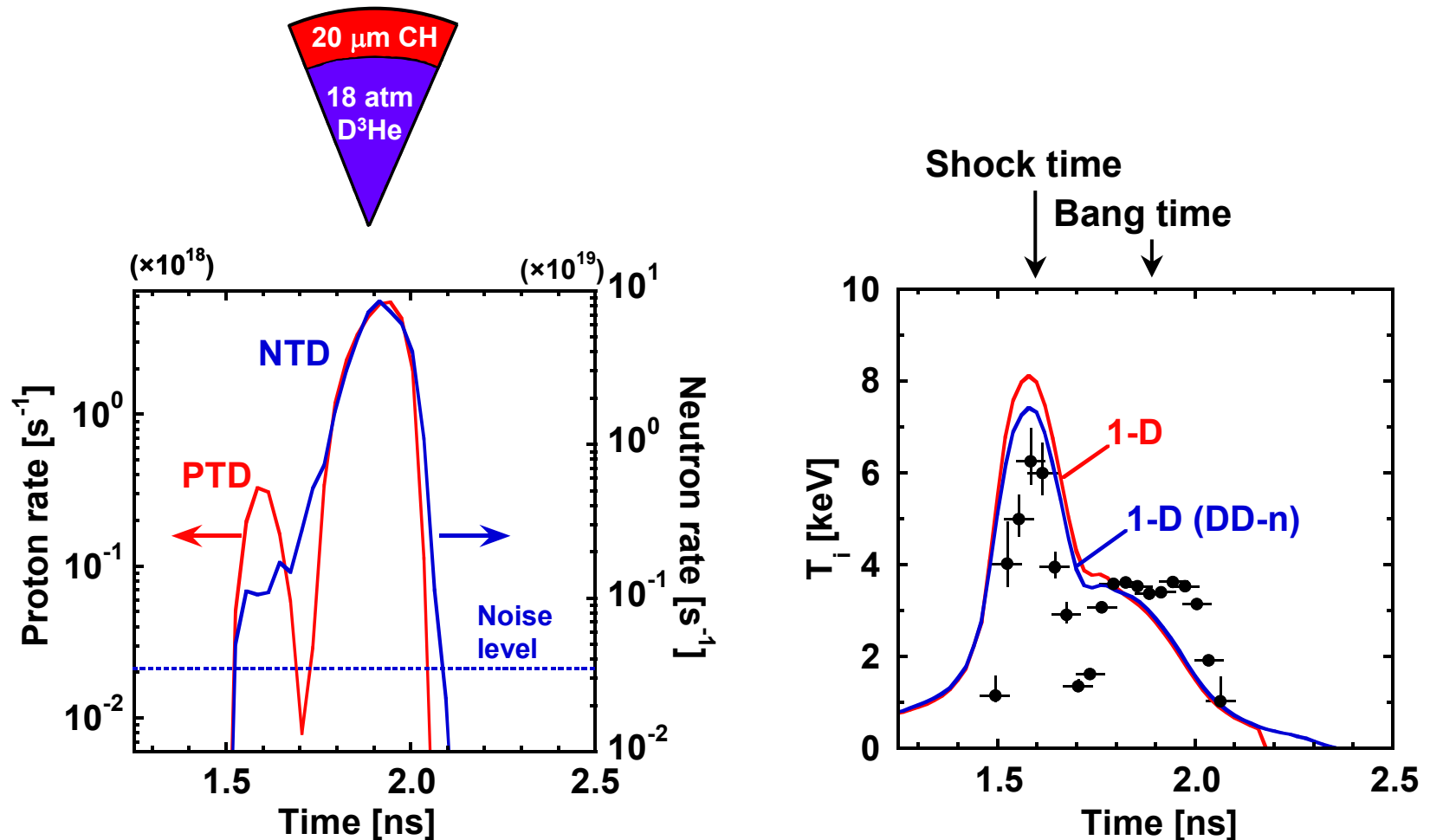
Shock time and shock-burn duration have been obtained and compared to 1-D calculations



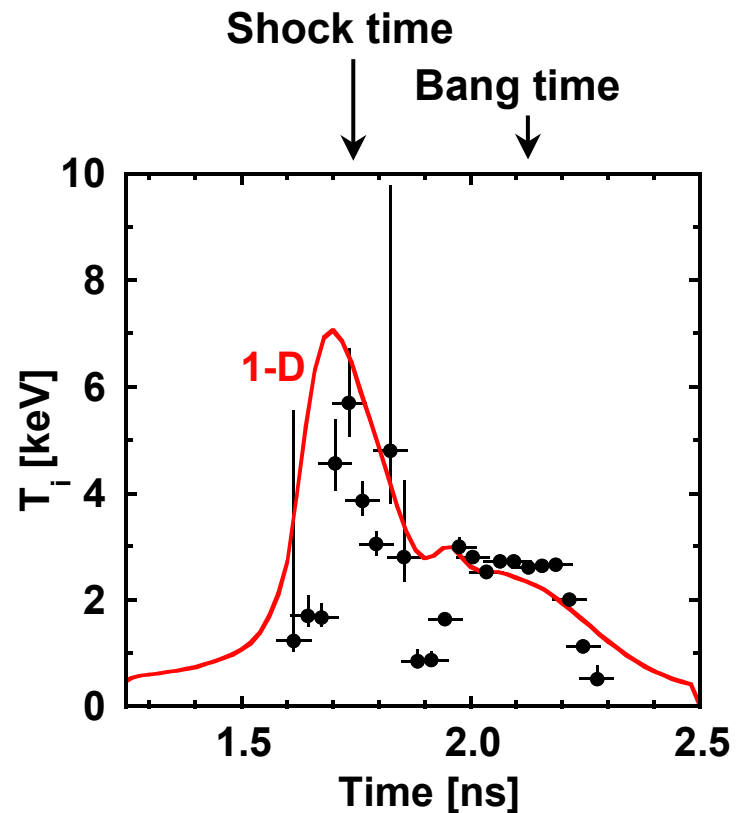
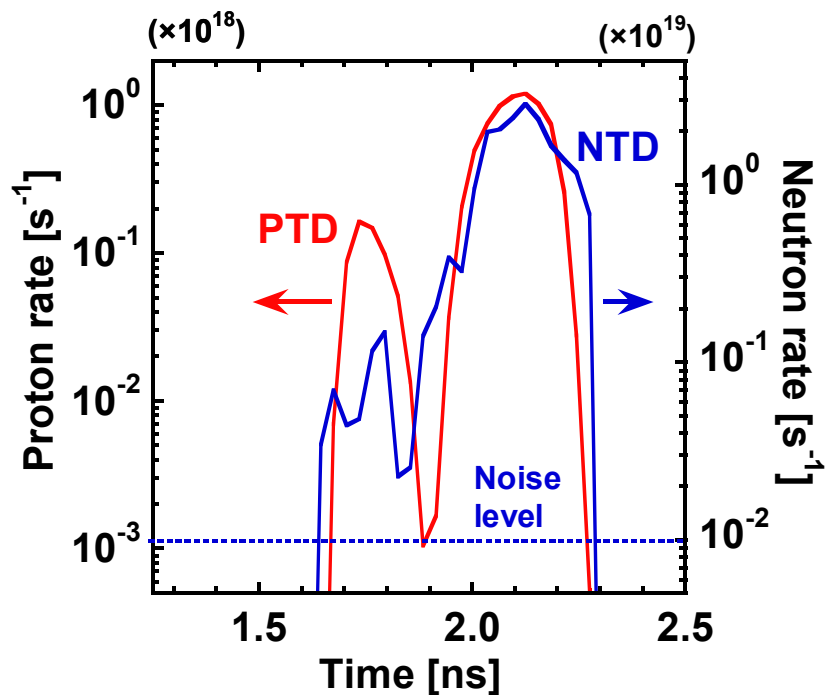
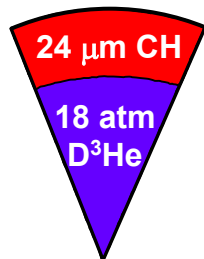
Evolution of T_i can be inferred from the ratio of D^3He and DD burn histories



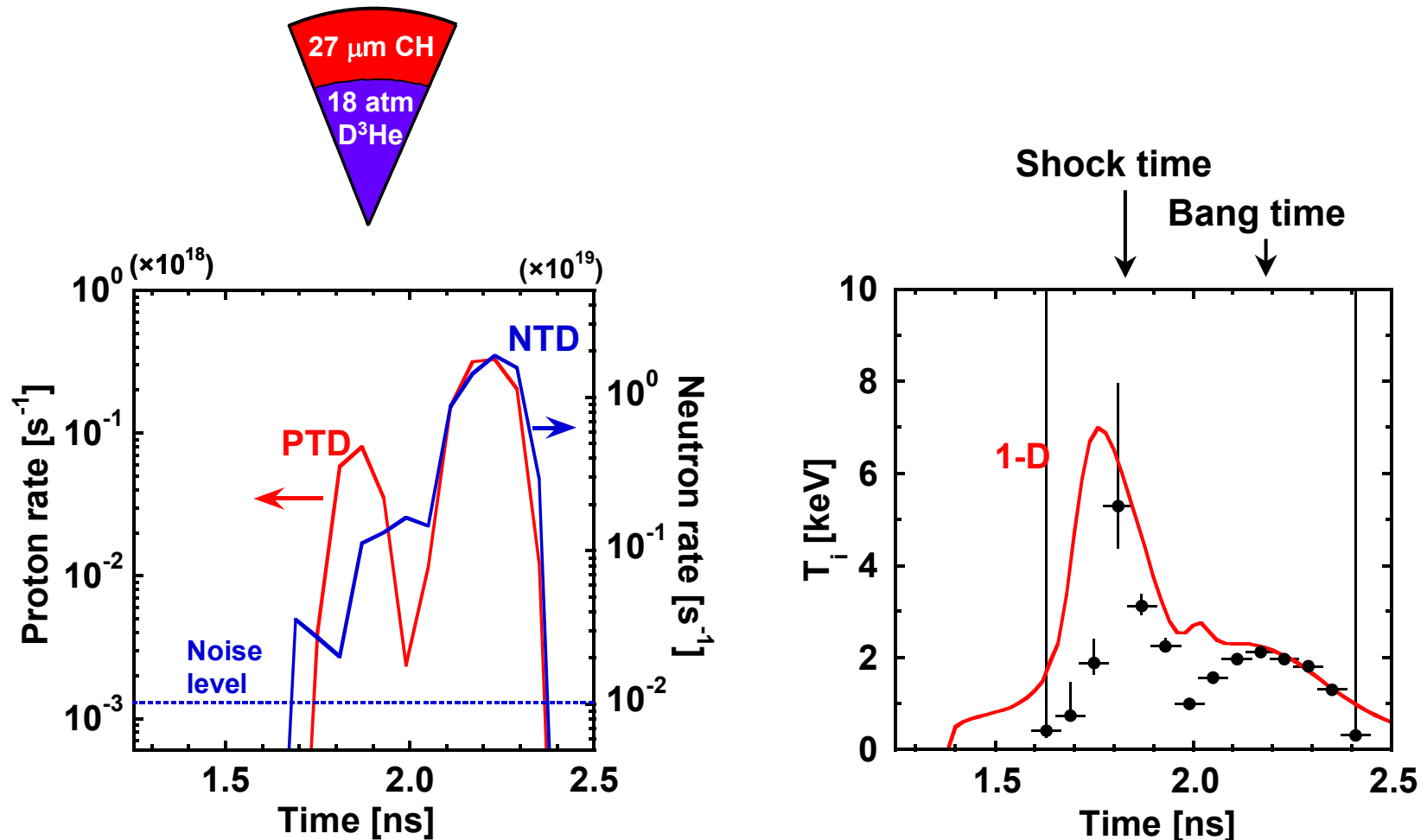
Evolution of T_i has been obtained and compared to 1-D calculations



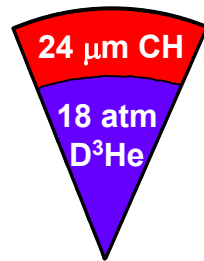
Evolution of T_i has been obtained and compared to 1-D calculations



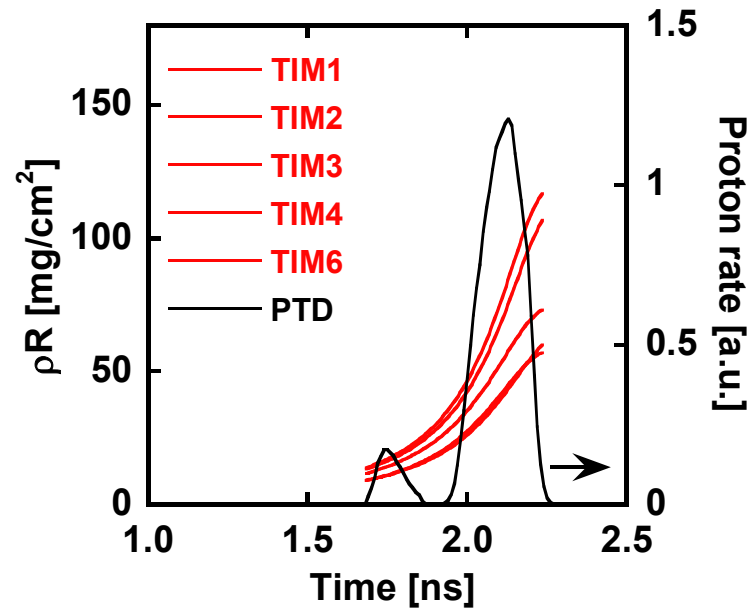
Evolution of T_i has been obtained and compared to 1-D calculations



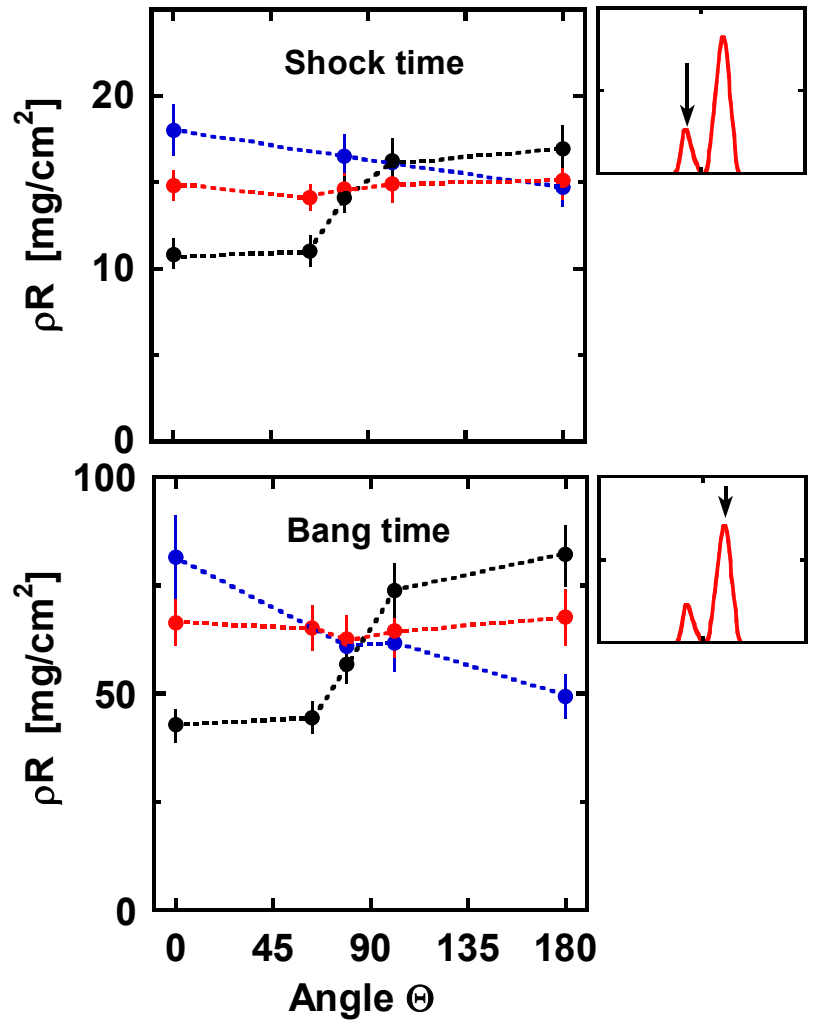
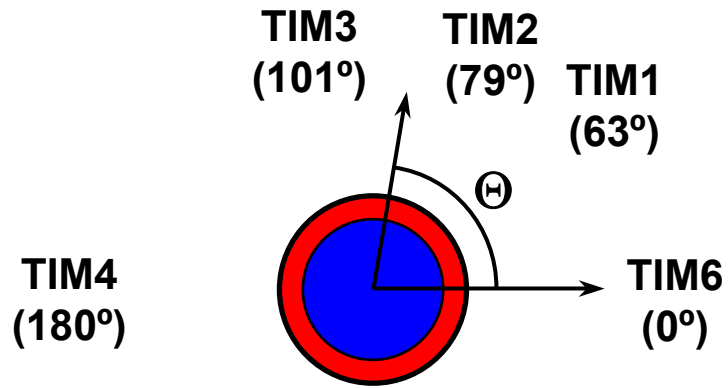
How do ρR and ρR asymmetries evolve in time?



Nominally perfect implosion



Low-mode ρR asymmetry at shock time is amplified and mirrored at bang time



Low-mode ρR asymmetry is primarily driven by capsule convergence

ρR asymmetry growth $\xi(t)$ can be expressed as

$$\xi(t) = \frac{Cr(t) - 1}{Cr_s - 1} \frac{\langle \rho R(t) \rangle}{\langle \rho R \rangle_s}$$

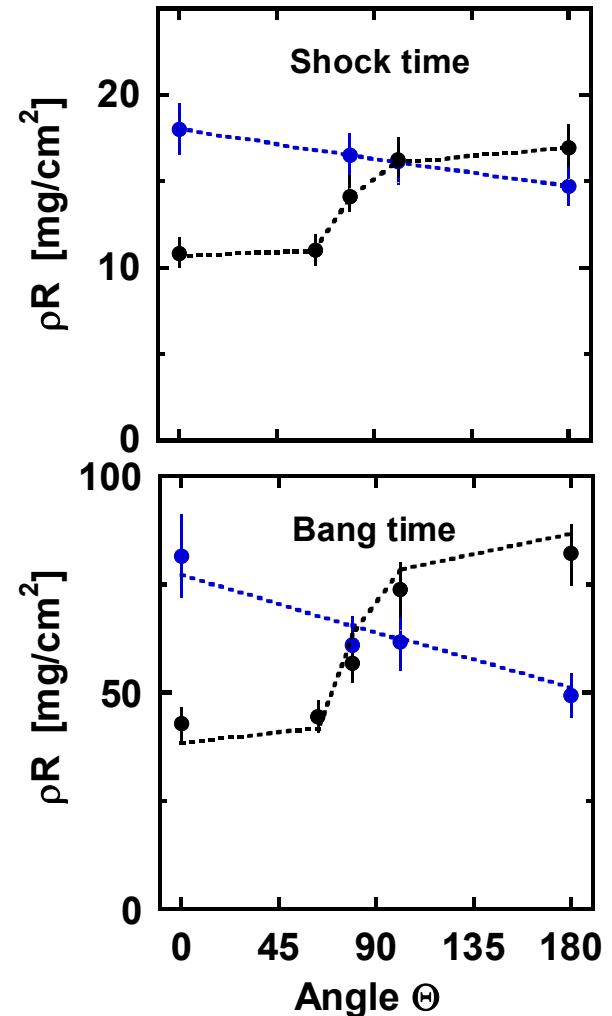
Convergence ratio $Cr(t)$ is defined as

$$Cr(t) = \sqrt{\frac{\langle \rho R(t) \rangle}{f \rho_0 R_0}}$$

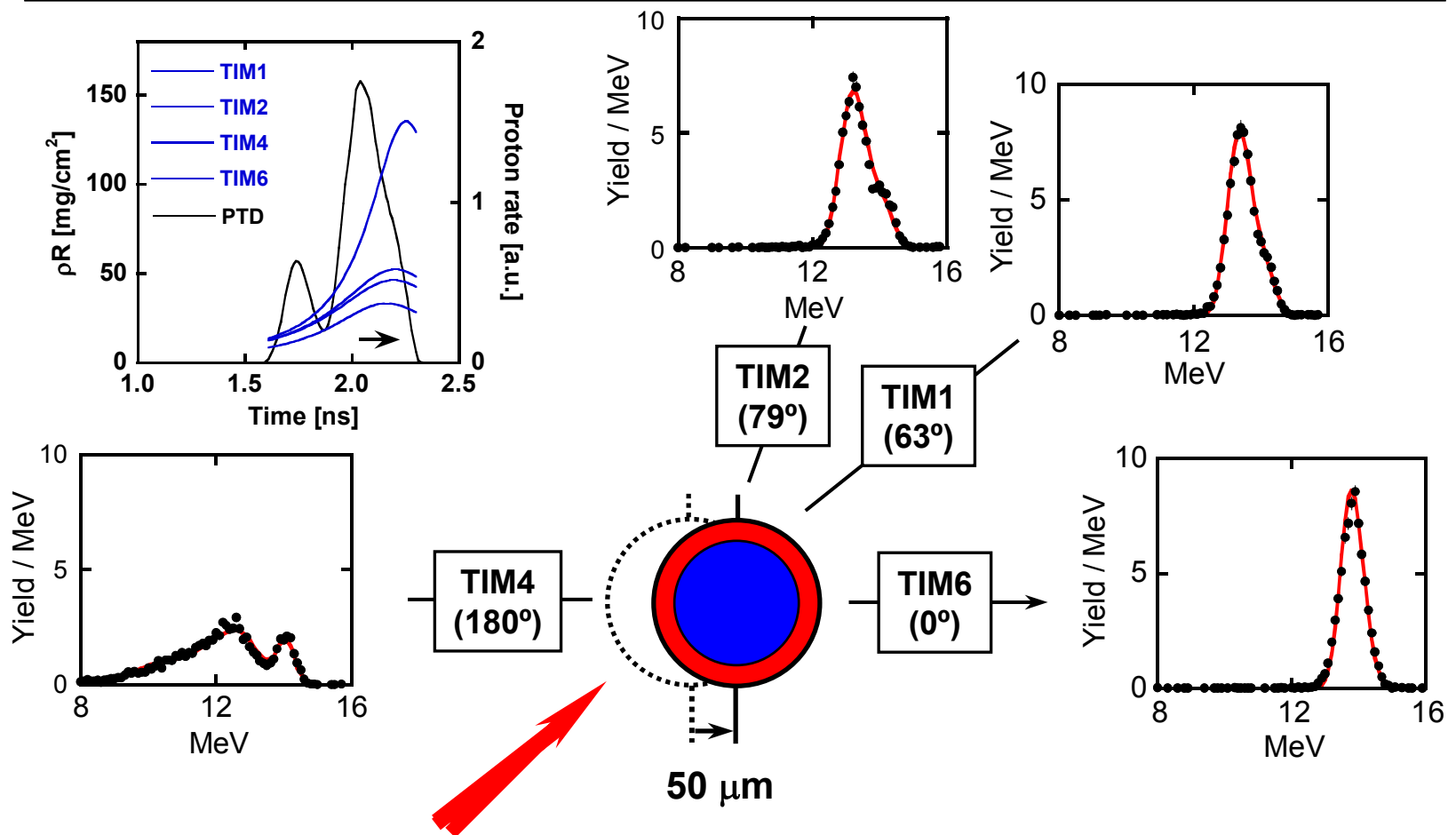
At shock time, $Cr_s \approx 5$

At bang time, $Cr_b \approx 10$

$$\xi(t) \sim 2 \frac{\langle \rho R(t) \rangle}{\langle \rho R \rangle_s}$$

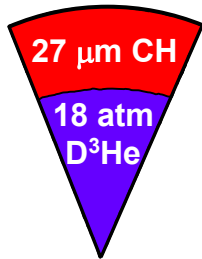


Is there a correlation between ρR asymmetry and laser drive asymmetry (for a large imposed $l = 1$)?

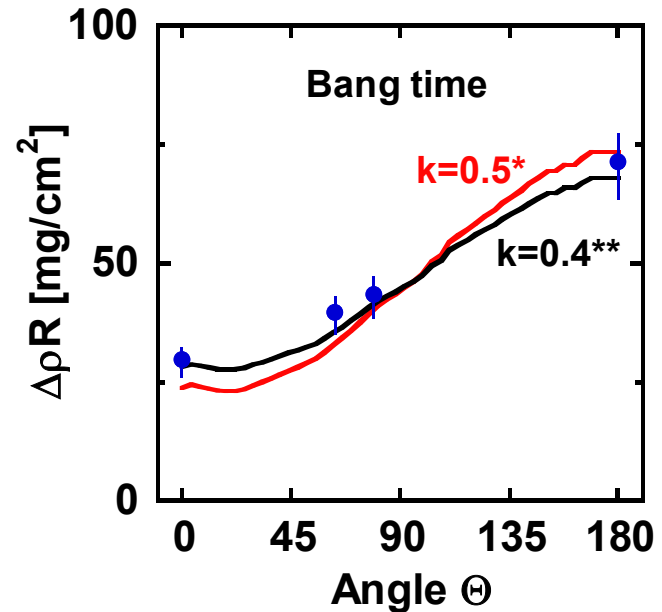
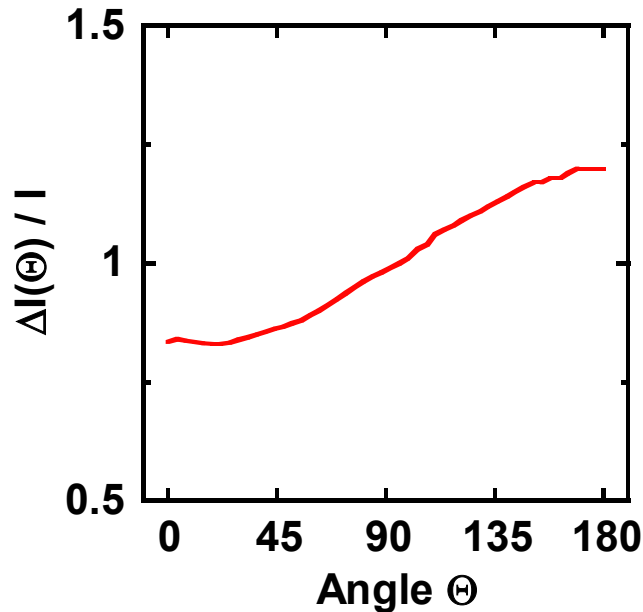


Laser beams pointed at TCC

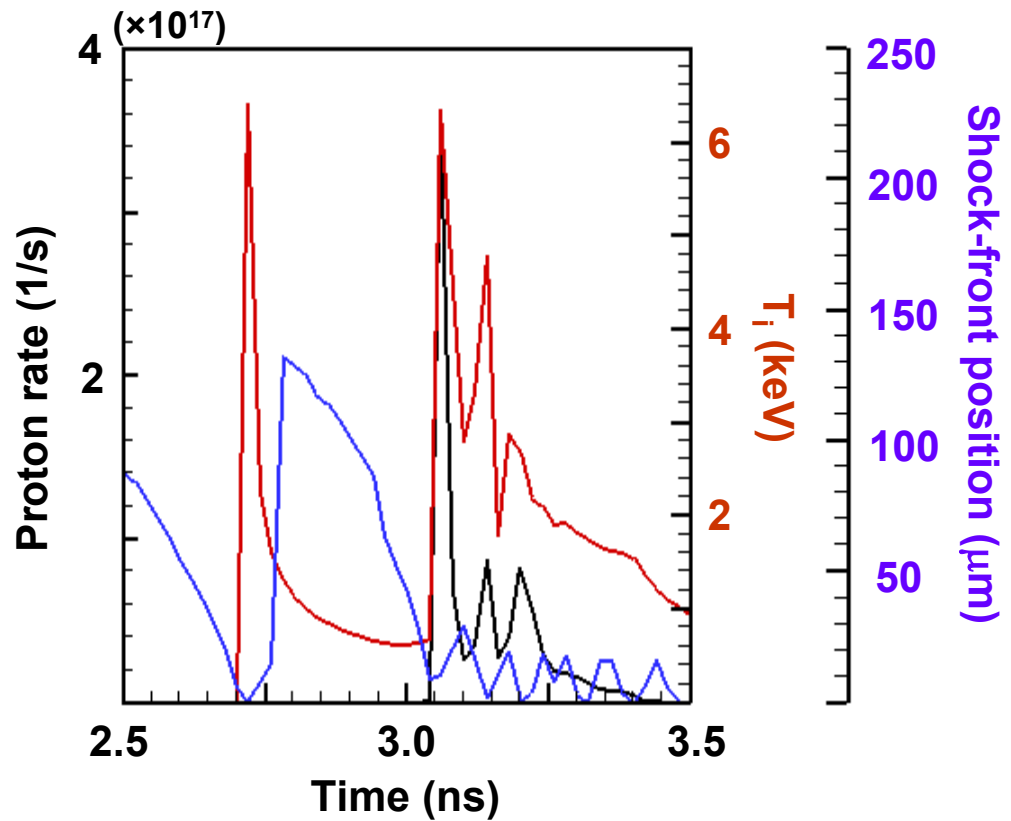
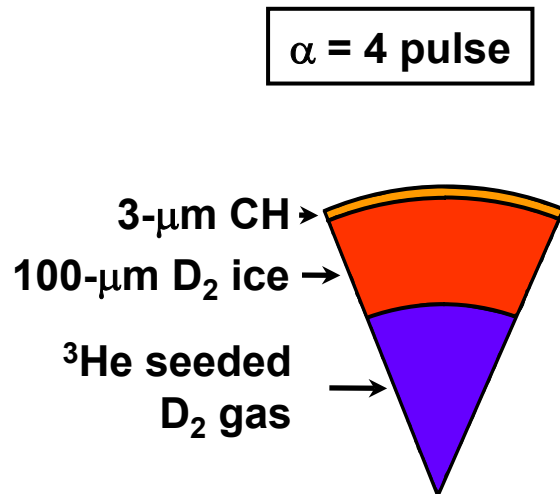
ρR asymmetry is strongly correlated to laser drive asymmetry (for a large imposed $l = 1$)



$$\frac{\Delta\rho R(t, \Theta)}{\langle \rho R(t) \rangle} = k [C_r(t) - 1] \frac{\Delta I(\Theta)}{\langle I \rangle}$$



We are looking into ^3He -seeded cryogenic D_2 implosions



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