Spectroscopic Measurements of Fuel–Pusher Mix in Direct-Drive Implosions on OMEGA



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The estimated core mass composition in the mix region of OMEGA direct-drive implosions with a CR \sim 15 is \sim 1/2 deuterium and \sim 1/2 CH

- Plastic shells with an Ar-doped deuterium fill gas were driven with a 23-kJ, 1-ns square laser pulse with full smoothing.
- The emissivity-averaged core electron temperature and density were inferred from the measured time-dependent Ar K-shell spectral line shapes.
- Ar K-shell emission probes the mantle of fuel surrounding the central hot spot (i.e., the mix region).
- The estimated core deuterium density ρ_f is 3.4(±0.9) g/cm³, and the estimated core CH density in the mix region is 3.6(±1.2) g/cm³.

Plastic shells with an Ar-doped deuterium fill gas were driven with a 23-kJ, 1-ns square laser pulse



- Predicted convergence ratio at stagnation ~15
- Laser irradiation with 1-THz SSD, PS, and on-target beam-to-beam power imbalance < 5% rms.

Streaked x-ray spectroscopy is used to measure time-dependent Ar K-shell spectral line shapes



 Absolute timing of peak x-ray signal is established with a slower streak camera.

Time-resolved Ar K-shell spectroscopy provides a direct measure of the core electron density to diagnose core-mix

- Ar K-shell emission from hot, dense plasma (n_e > 1 × 10²³ cm⁻³ and T_e > 1 keV) has line shapes that depend strongly on density and are insensitive to variations in electron temperature.
- The relative intensities of the Ar K-shell lines and their associated L-shell satellites are sensitive to variations in electron temperature and density.
- Stark-broadened Ar K-shell resonance lines and satellites are calculated with MERL^{1,2} assuming uniform core conditions.

 ¹R. C. Mancini et al. Comput., Phys. Commun. <u>63</u>, 314 (1991).
²D. A. Haynes et al. Phys Rev., E <u>53</u>, 1042 (1996).

Time history of emissivity-averaged core n_e and T_e is inferred from the Ar K-shell spectroscopy



- Significant changes in the Ar K-shell emission spectrum occur during the implosion.
- Peak neutron production is measured ~170 ps±100 ps before peak x-ray production.
- LILAC simulations indicate that peak neutron production occurs at the same time as peak emissivity-averaged T_e.

At peak neutron production, $n_e = 2.5 \times 10^{24} \text{ cm}^{-3}$ and $T_e = 2.0 \text{ keV}$



1-D simulations reveal the Ar K-shell emission probes the mantle of fuel surrounding the central hot spot



Emissivity average:

weighting function = $\epsilon(\mathbf{r}) \cdot \mathbf{V}_{shell}(\mathbf{r}) / \Sigma \epsilon(\mathbf{r}) \cdot \mathbf{V}_{shell}(\mathbf{r})$

Spectroscopic results are compared with nuclear measurements of $\rho_f R$ to estimate the amount of core-mix

• The $\rho_f R$ measurement is obtained from knock-on deuteron spectra recorded on similar implosions with a DT fill gas.

 $\rho_f \mathbf{R} \text{ (knock-ons)} = 15 \text{ mg/cm}^2$

• The radius of the imploding core at peak neutron production is estimated from $M_f = 4/3 \pi (\rho_f R) R^2$.

R = 44.6 μ**m**

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$$\rho_{f} = 3.4 \ (\pm 0.9) \ g/cm^{3}$$

Ar dopant has a limited effect on target performance



Shot(s)	Target	$Y_n \times 10^{11}$	Y _{2n} /Y _n	$\rho_{f} \mathbf{R} (\mathbf{from} \mathbf{Y_{2p}})$	$\rho_{f} \mathbf{R} \ (\mathbf{from} \mathbf{Y_{2n}})$
22507	DD (15), Ar (0.18%), CH (20)	1.0	1.8 × 10 ^{−3}	>8.5 mg/cm ²	<18.1 mg/cm ²
Multiple	DD (15), CH (20)	1.3(±0.08)	$2.1(\pm 0.28) imes 10^{-3}$	—	_

• $\rho_f R (knock-ons) = 15 mg/cm^2$

Direct-drive implosions with CR \sim 15 have an estimated core mass composition in the mix region of \sim 1/2 deuterium and \sim 1/2 CH

DD (15), Ar (0.054),	CH [20]
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MER

 $n_e(CH) = n_e - n_e(D) - n_e(Ar)$

n _e	2.2 × 10 ²⁴ cm ⁻³ (averaged over 170-ps neutron burnwidth)
Т _е	1.9 keV (averaged over 170-ps neutron burn width)
n _e (D)	$1.0 imes 10^{24} \text{ cm}^{-3}$
n _e (Ar)	$3.1 imes 10^{22} ext{ cm}^{-3}$
n _e (CH)	$1.2 imes 10^{24} \text{ cm}^{-3}$
р <mark>сн</mark>	3.6 (±1.2) g/cm ³
₽ f	3.4 (±0.9) g/cm ³

- Pressure (electron + ion) ~ 11 Gbars
- * $n\tau T\sim 7\times 10^{20}~m^{-3}s~keV$

Summary/Conclusions

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