### Theoretical Interpretation of SBS Observations in OMEGA Long-Scale-Length Plasma Experiments

352.0 SBS Wavelength (nm) 351.5 laser 351.0 350.5 350.0 1.5 2.0 2.5 3.0 3.5 0.5 1.0 Time (ns)

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UR LLE Summary

## SBS arises primarily in "hot spots" and seems to be seeded by light reflected from critical

 Polarization smoothing (PS) reduces the level of SBS to that seen at half the incident intensity without PS, implying that SBS levels are determined by hot-spot intensities.

- Overlapping beams do not seem to "cooperate" in driving SBS.
- The red-shifted portion of the spectrum appears to derive from light reflected from critical, but it is difficult to account for levels and shifts.



- Aspects of SBS spectra
- Ion-acoustic modes in multicomponent plasmas
- Strong-damping SBS model and calculation of growth factors in simulated profiles
- Summary and conclusions

#### SBS spectrum consists of distinct "red" and "blue" features



### "Blue" feature depends exponentially on hot-spot intensity



### "Red" feature depends linearly on hot-spot intensity



## Damping is fairly high, even for the least-damped ion-acoustic mode



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# In strongly damped plasmas the SBS gain may computed by integrating a local gain factor

• The equation for SBS intensity is<sup>1</sup>  $\frac{\partial ISBS}{\partial x} + \frac{ISBS}{L_{abs}} = \frac{I_{pump}ISBS}{L_{gain}}$ .

Here,  $L_{abs}$  is the aborption length and  $L_{qain}$  is the local gain length:

$$\mathbf{L}_{gain}^{-1} = \frac{\mathbf{k}_{0}}{4} \frac{\mathbf{n}_{e} / \mathbf{n}_{c}}{\sqrt{1 - \mathbf{n}_{e} / \mathbf{n}_{c}}} \frac{\mathbf{m}_{e} \mathbf{v}_{osc}^{2}}{\mathbf{T}_{e}} \left[ \left(1 + \frac{3\mathbf{T}_{i}}{\mathbf{Z}\mathbf{T}_{e}}\right) \left(\frac{\mathbf{v}_{i}}{\omega_{s}}\right) \right]^{-1} \mathbf{p}(\eta),$$

where 
$$\mathbf{p}(\eta) = \frac{\left(\frac{v_i}{\omega_s}\right)^2 \eta}{\left(\eta^2 - \mathbf{1}\right)^2 + \left(\frac{v_i}{\omega_s}\right)^2 \eta^2}$$
 and  $\eta = \frac{v_0}{c_s} + \frac{\omega_i}{\omega_s}$ .

• The simulation code SAGE is used to provide the profiles of the plasma parameters over which the above equations are integrated.

<sup>&</sup>lt;sup>1</sup>C. J. Randall, J. A. Albritton, and J. J. Thomson, Phys. Fluids <u>24</u>, 1474 (1981).

## The peak computed gain as a function of wavelength agrees well with "blue" feature



### The present model does not account for some observed features of the SBS emission

- Levels of the "red" feature lie below those expected from simple inverse bremsstrahlung absorption.
- The increasing red shift at later times is not accounted for by the SBS gain factor or the bulk hydro motion.
- These phenomena may result from hot-spot behavior near critical, e.g., enhanced localized absorption and Doppler shift.

## Even at lower intensities the time history of the red feature suggests SBS rather than simple reflection



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