

Hard x-ray imaging on NIF

CEA-NNSA Joint Diagnostic Meeting

Rochester, NY, June 2016

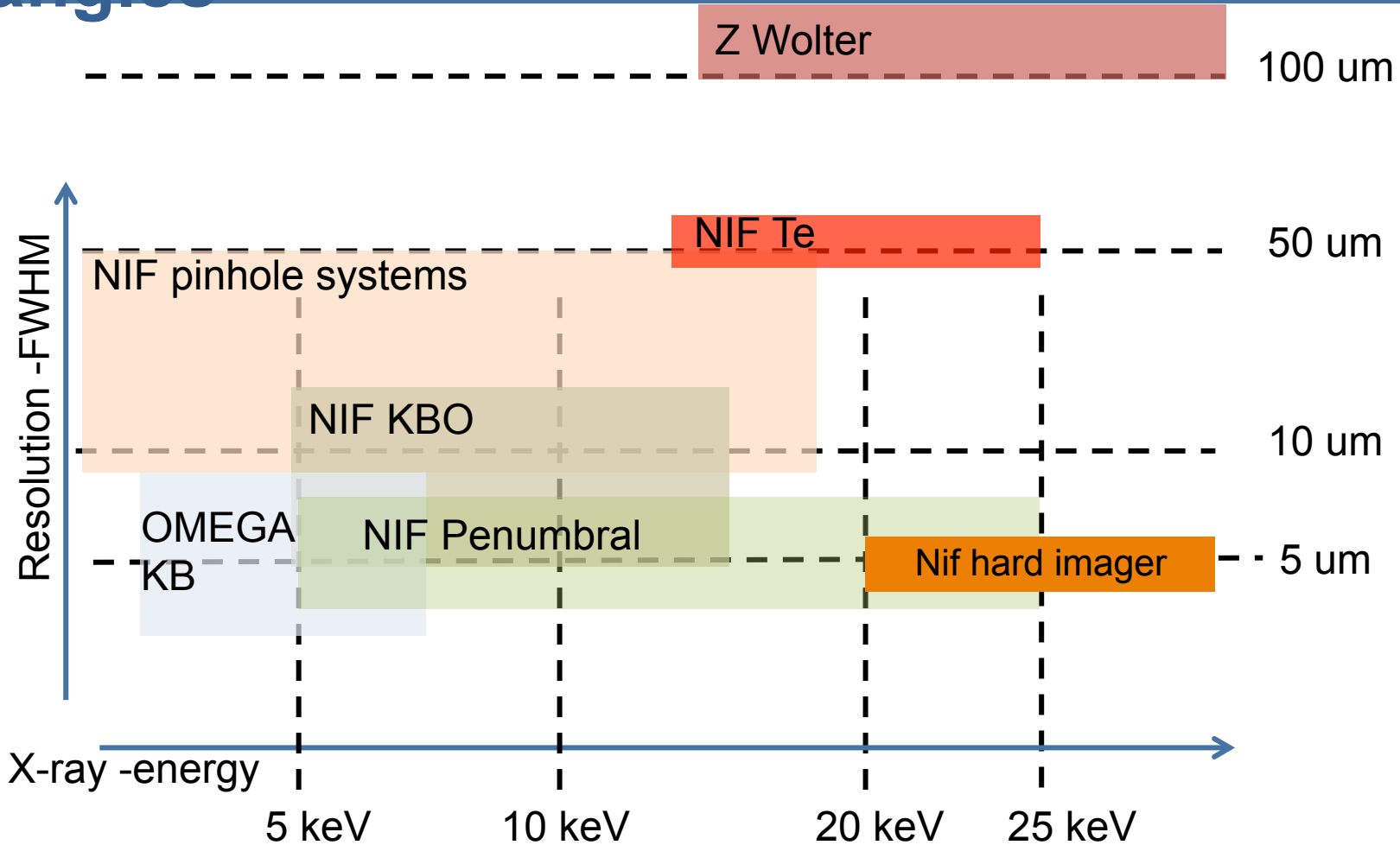
Louisa Pickworth



Overview of hard X-ray imaging

- Summary of existing imaging systems for ICF and HED plasmas
- Motivational experiments that require hard x-ray imaging
- Simple specification of the ideal system
- Issues common in hard x-ray imaging. It's all about the collection efficiency!
- Implementations at NIF (and Z) of high throughput (large solid angle) x-ray microscopes for hard x-ray imaging
- Replicated shell Wolters and the R&D challenge of reaching 5 μm resolution
- What other microscope systems fit the bill?

To image ICF plasmas above 20 keV with suitable resolution requires large solid angles



Detectors, sources and back lighters are weaker/ less efficient at high energy, demanding better collection efficiency per res. element from the imaging system

Motivation to pursue hard X-ray imaging on NIF

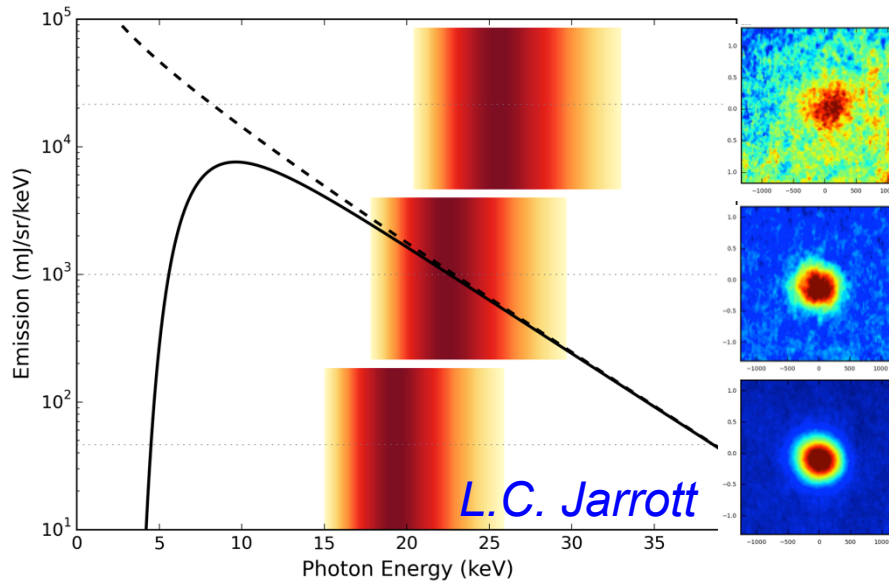
- ICF
 - Hot spot temperature (>15 keV, $5 \mu\text{m}$)
 - Hot spot shape (>15 keV, $5 \mu\text{m}$)
 - Close in radiograph (>16 keV, monochromatic, $5 \mu\text{m}$)
 - Shell density
 - Compton (>50 keV, broad band, better served with point projection)
 - Self compton (50 keV, mono chromatic?)
 - Double shell? (>20 keV self emission?)
- Material strength (>20 keV, monochromatic, $<10 \mu\text{m}$)

Difficulties of imaging Hard X-ray in ICF/ HED

- Weak sources (comparatively)
- Short life time (< 1 ns, requiring high speed detectors)
- Low DQE detectors ($\sim 1\%$)
- Low resolution detectors (FWHM > 50 μm)
- Bright low energy emission which needs to be filtered out

We have begun measuring spatially and time integrated hot spot temperature

T_e measurement, (N160411-001)
(50 μm pinholes @ 10 cm)



Measured source:

$$I = 2e^{-\left(\frac{E-11}{kT_e}\right)} \quad \text{J/sr/keV}$$

Where kT_e and $h\nu$ in keV. $kT_e \sim 4$ keV

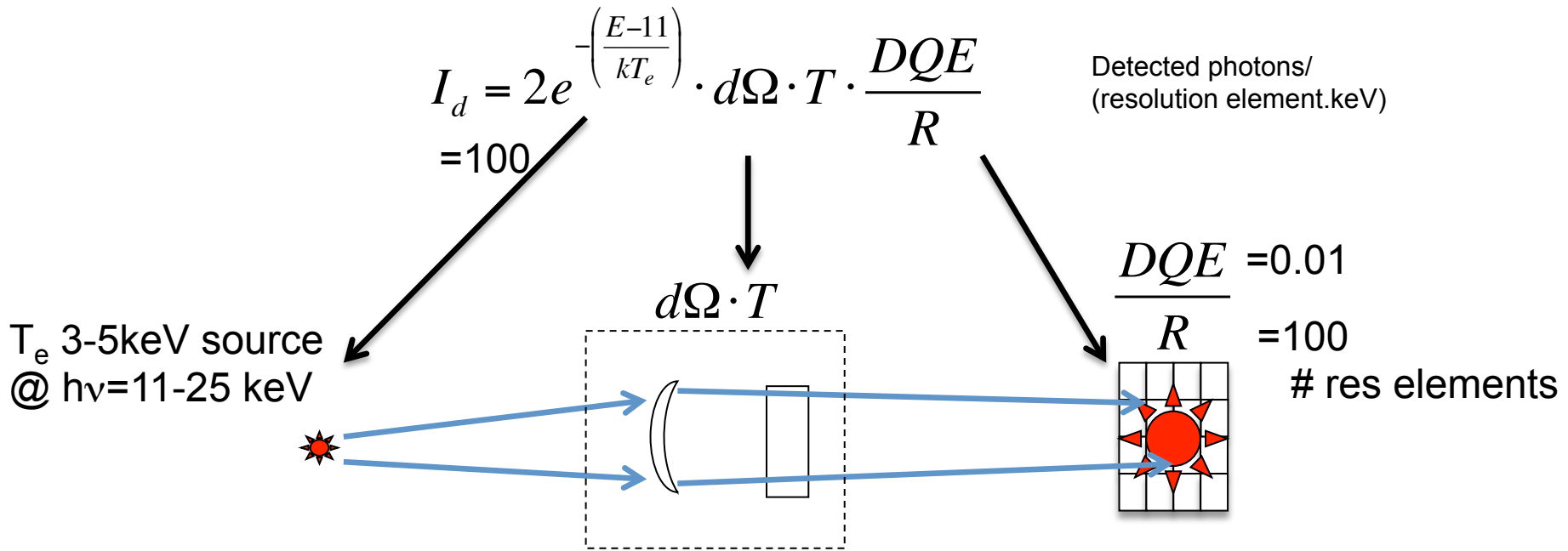
Existing resolved imaging- not shown
is at ~ 10 keV

Goal: Spatial resolved T_e measurement.

Requires 5 μm PSF, ~ 50 μm image, $S/N > 10$ ~ 1 keV bands @ 15, 20, 25 keV

The “ideal” specification for a T_e measurement

An example for Hot spot T_e measurement:

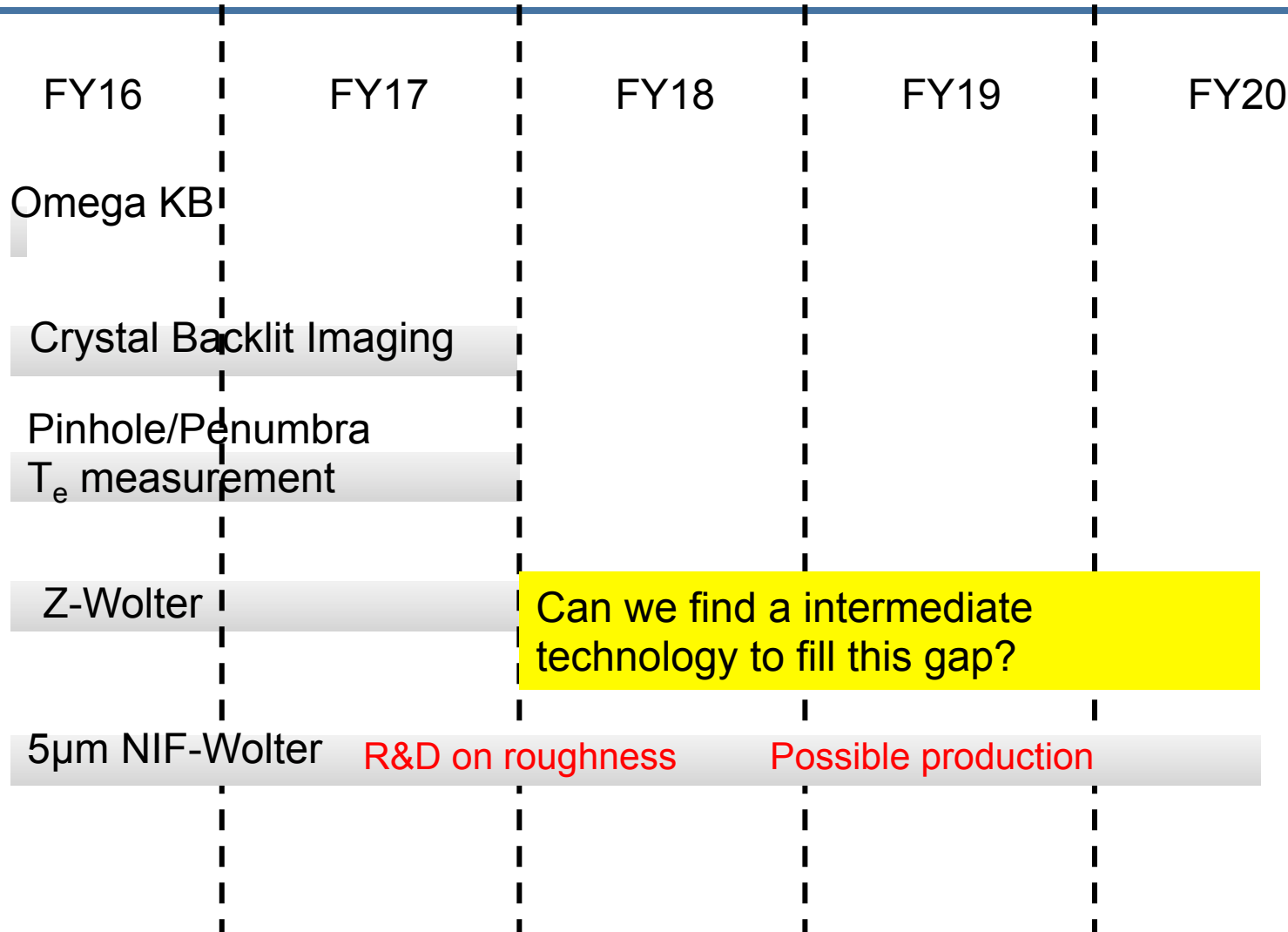


Solid angle * Throughput should be $>10^{-7}$, we need $>\sim 10^{-6}$ to allow shot to shot variation in shape, T_e and size and time resolution

We have started investigating how to meet the requirement with new microscope systems

- To date, (with some notable exceptions) NIF has focused on variations of pinhole and slit imaging
 - We have recently built KB
- Crystal backlight imaging (CBI)
- Penumbra
- Wolters

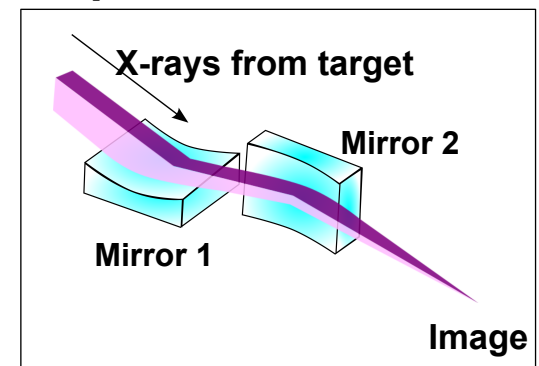
Time line to achieve a high energy imaging system for HED /ICF



KB's and Toroidal optics play a role for hotspot imaging at lower energies

- We have recently developed and implemented a Kirkpatrick-Baez imager (KBO)
 - Limited in solid angle
- We are in discussions exploring similar 2-mirror imaging systems (Toroids) that have a larger solid angle.. See P. Troussel et al.,
 - ? How many could we get
 - ? How quickly can we make it happen

Kirkpatrick-Baez Scheme



Penumbral imaging on DIXI could reveal detailed hotspot structure

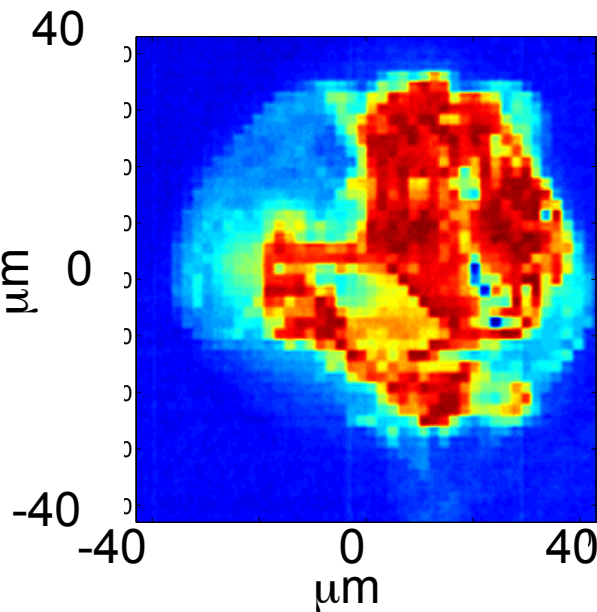
Possible ICF setup w/ DIXI diagnostic:

- Pinhole diameter: $100\mu\text{m}$
- Magnification: $M=64$
- DIXI-Point Spread Function: $270\mu\text{m}$ FWHM Gauss
- 10% of pixels affected by neutron noise (10% dead pixels)

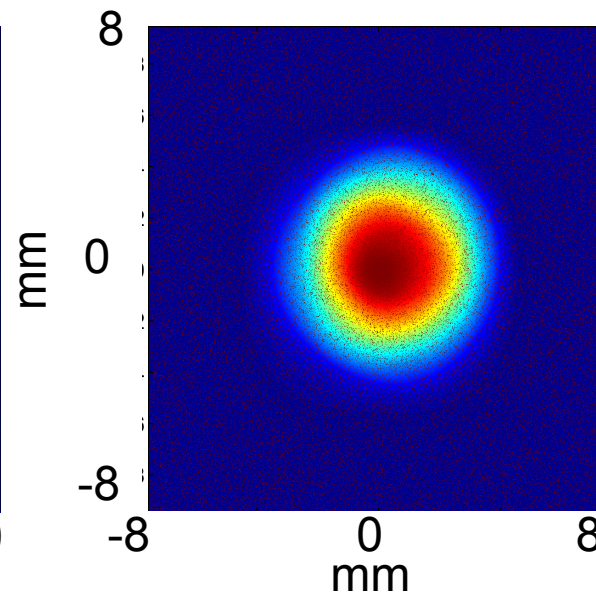
Expected SNR > 10

Resolution: $\sim 5\mu\text{m}$

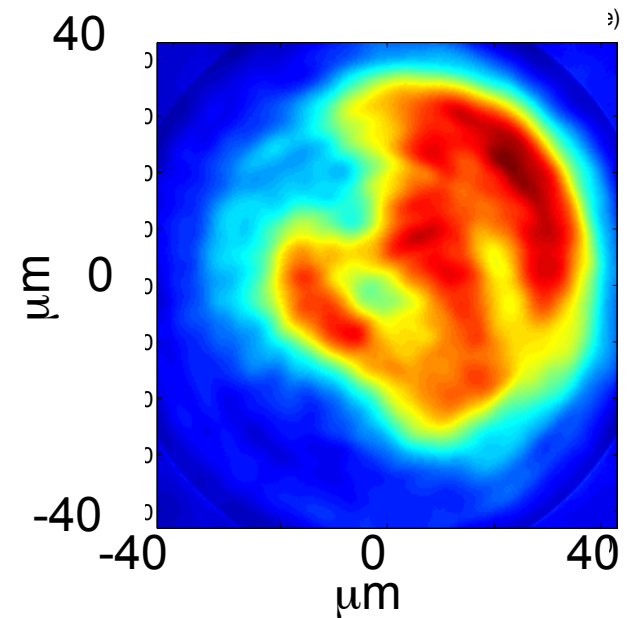
Fresnel limit ($a=10\text{cm}$): $3.6\mu\text{m}$



Simulated hotspot

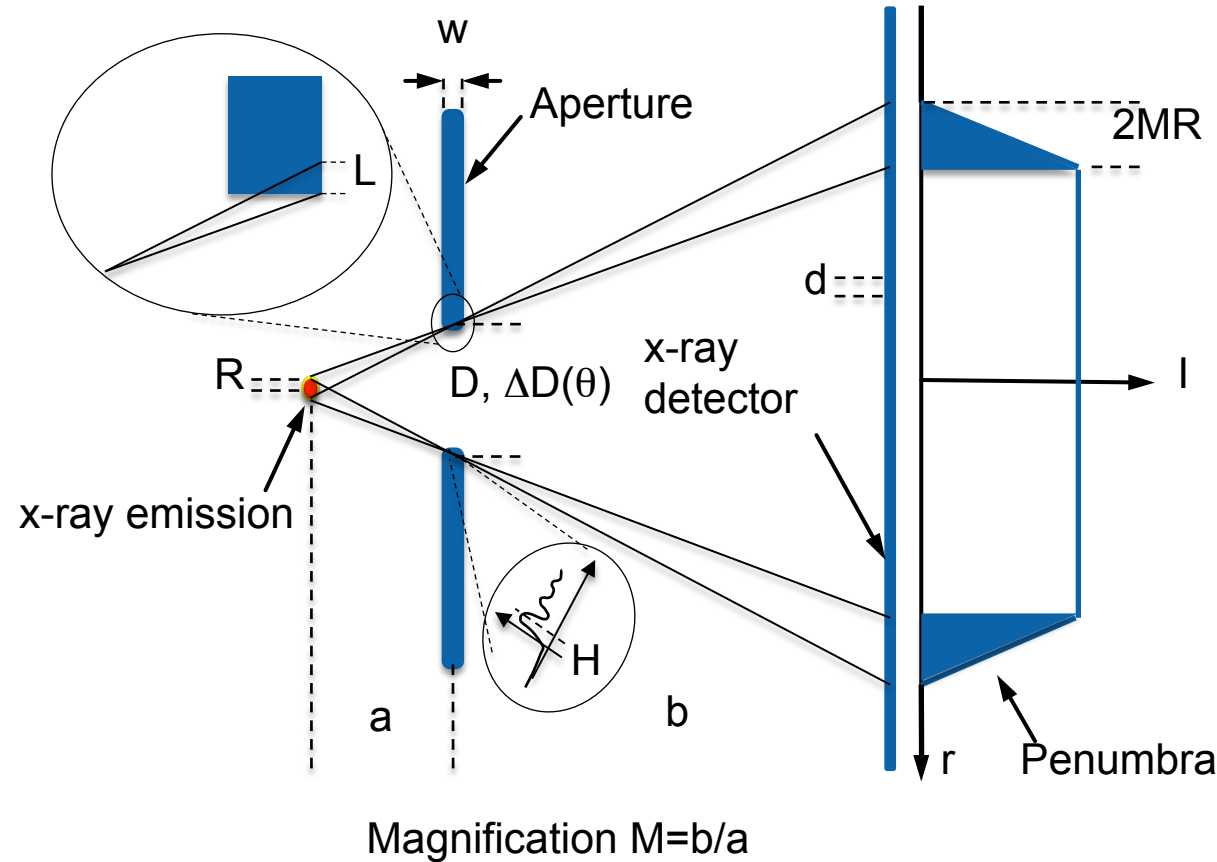


Penumbral Image



RECONSTRUCTED

Penumbral Imaging uses the edge of an aperture to encode the image



- Source radius R
- Imaged through aperture D
- Yielding Penumbra $2MR$ on detector

Physical limits:

- Fresnel Diffraction
- Effective optical density length scale L of aperture (thickness w)

Technical limits:

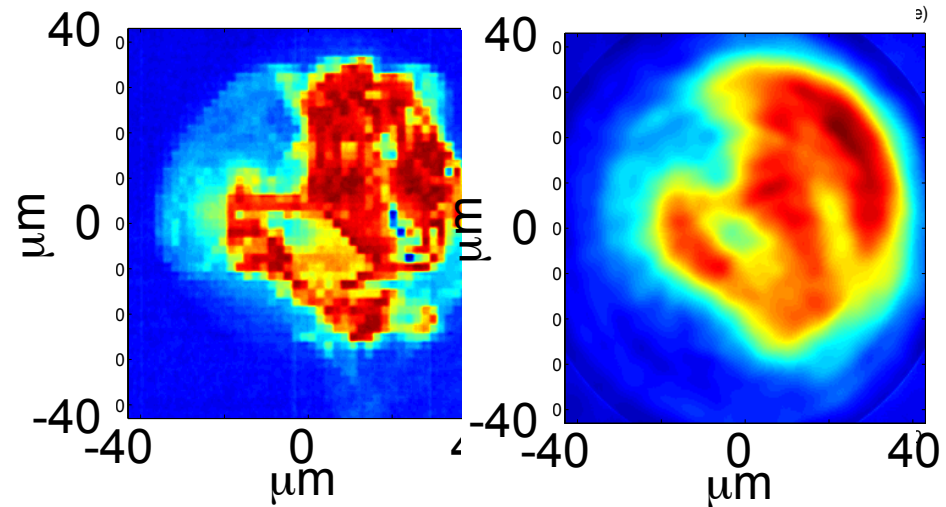
$$H \approx \sqrt{\lambda a}$$

- Det. Resolution d
- DQE
- SNR
- Aperture circularity $\Delta D(\theta)$

X-ray Penumbra Imaging can be a powerful technique to improve understanding of stagnated implosions

- Few μm resolution penumbral imaging has been successfully fielded on the NIF
- Manufacturing challenges of quality apertures have been overcome

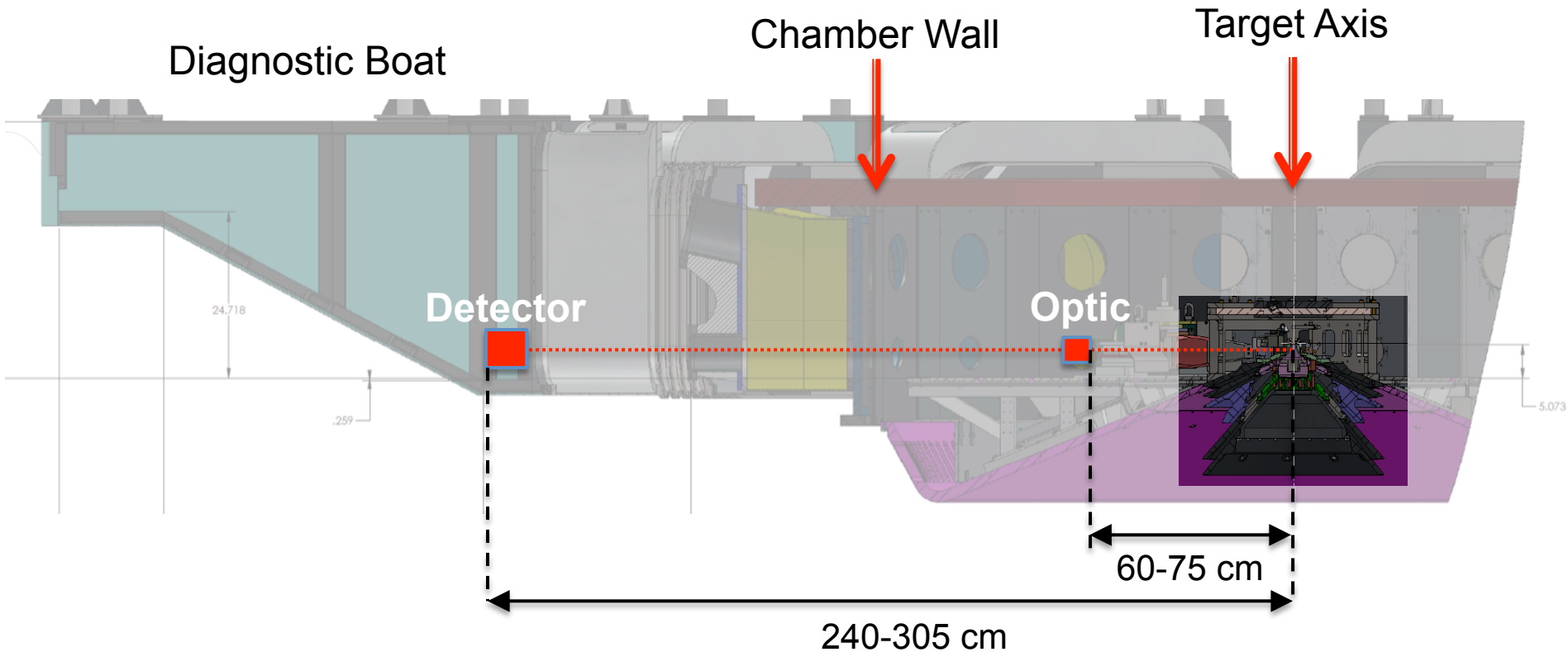
- Challenges of Penumbra imaging:
 - Aperture characterization and build for high energy
 - Energy selection
 - Fielding in NIF (close in?)



Wolters and multilayers

- Wolter optics promise large solid
 - Current technology achieves $\sim 30\mu\text{m}$ spatial resolution
 - Waiting on advancement of the technology achieve $\sim 5\mu\text{m}$ resolution
 - 3+years and a R&D project to do this, so there is risk
- Multilayer coatings that operate at grazing incidence allow the possibility of pseudo monochromatic or band-like energy response (implemented in NIF KBO1 at 10.2keV)
 - Multilayer with higher reflectivity are of great interest...

Wolter for Z at Sandia



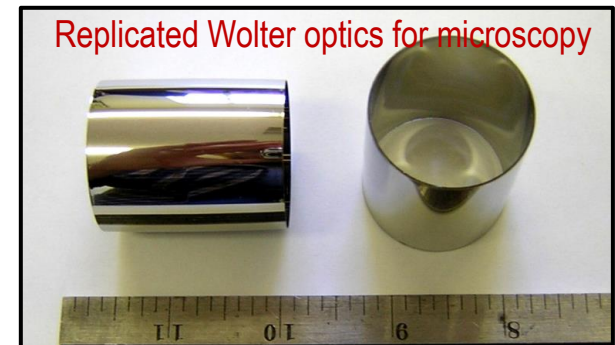
- Imaging K_{α} emission to diagnose stagnation conditions
- Wolter microscope ($M= 3.5$) on the Z target chamber with optic ~ 70 cm from the pinch
- Geometry similar to what has been previously demonstrated with multi-layer Wolters

Wolter for Z at Sandia

Need	Goals	Driver
Photon energies (K α 's)	Mo: 17.479 keV Ag: 22.163 keV W: 59.318 keV	Study K-shell radiators from Ag to W
Spectral window	~1 keV	Simultaneously view K α 1 & 2 from cold and low-ionization states
Field of View	2 cm in each direction	Collect all emission from 2 cm pinch. K α emission comes from large diameter
Spatial Resolution	~100 μ m	Resolve length-scale of structures emitting K α
Time resolution	Time integrated initially ~1 ns in 3-5 years	Resolve evolution over pulse
Sensitivity	100 J from ~cm ³ source with good S/N	Able to record 100 J over full source area

Other considerations

- Survivable (optic at >40 cm from source)
- Hard x-ray background (1-inch W in LOS)



Advancing Wolter optics to 5 μ m resolution for NIF

- Ability to focus is limited by low-spatial frequency errors (i.e. “figure”)
- Resolution degrades off-axis even for perfect mirrors (i.e. $\Delta\alpha=0$):
- To first order, the angular resolution of the optic is driven by angular resolution of the master mandrel

Main Challenge: Achieving small figure errors on small mandrels

Additional challenges: Minimizing electric-field edge effects in tuning the electro-plating processes for small-radius, “stubby” optics (those, in theory, ideal for NIF)

→ Results in difficulties to achieve the same figure quality in replicated optic as in mandrel

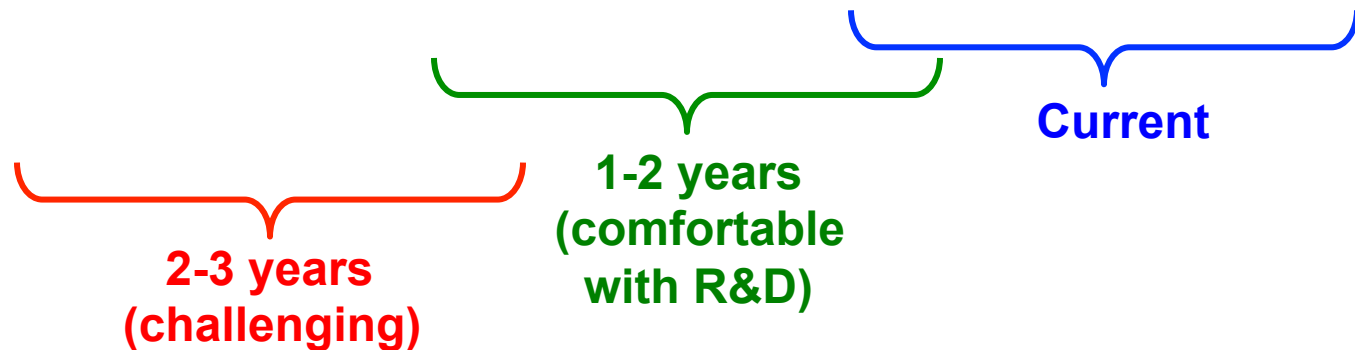
Current best performance for NASA MSFC replicated optics from optics with $L = 300 \text{ mm} \times 2$ and radius of 35–60 mm (larger than what is needed for NIF)

Path to high resolution Wolter optics for NIF

Preliminary NIF design: Total throw 8.0 m, grazing angle $\geq 0.4^\circ$

Spatial resolution, R (μm)

Mag $\Delta\alpha$	1"	2"	3"	4"	5"	6"
10	5	10	15	20	25	30
12	4	8	13	17	21	25



Comparison of hard X-ray microscope systems for hotspot imaging

	Multiple 5 μ m Pinholes mounted on target	Penumbra mounted on target	CBI	Wolter
Approx. Solid angle/through put	1e-7	1e-7*filter	2e-6	1e-6*R ² *efficiency
Number of images	10	1	1	1
	*5 μ m , 5mm from TCC	*100 μ m, 8cm from TCC, Mag 16	*60eV band,	

Summary: Hard x-ray imaging at NIF

- We are implementing several imaging systems on NIF to image (spatially and temporally resolved) hard X-rays
- Penumbra imaging and Wolter optics are the two technologies we have focused our efforts on
- Wolter optics with the desired resolution (5 μ m) are years away (2020)
- Can we apply a different technology in the mean time?

KBFRAMED results: LANL-HED 18-Feb-2015



OMEGA shot 76276

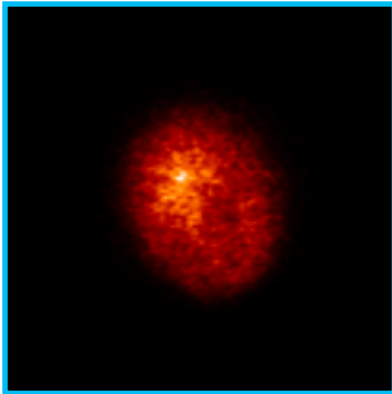


image 6

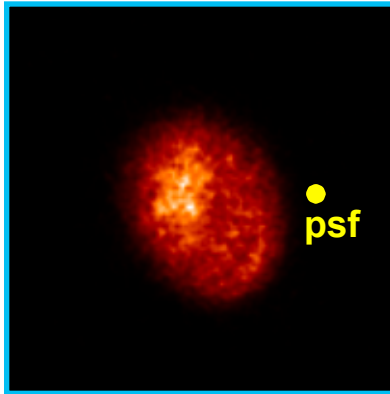


image 7

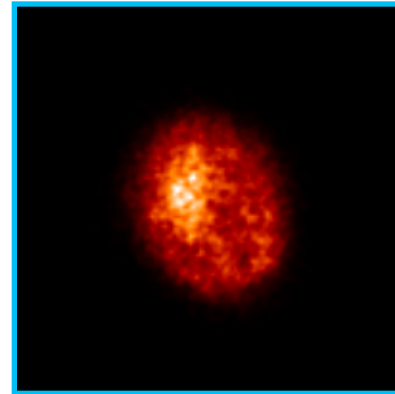


image 8

**KBFRAMED time integrated
images**

**Filter: 5 mils Be, 4 mils Al
200 x 200 μm region**

F. J. Marshall
26 Feb 2015

Crystal backlight imager

- Primarily designed for backlight application for “close in imaging” at $\sim 16\text{keV}$
- We could adapt the system for self emission
- A good amount of R&D is needed to test making and using crystals at 20keV
- Promises large solid angle, but very narrow energy band ($\sim 60\text{eV}$)