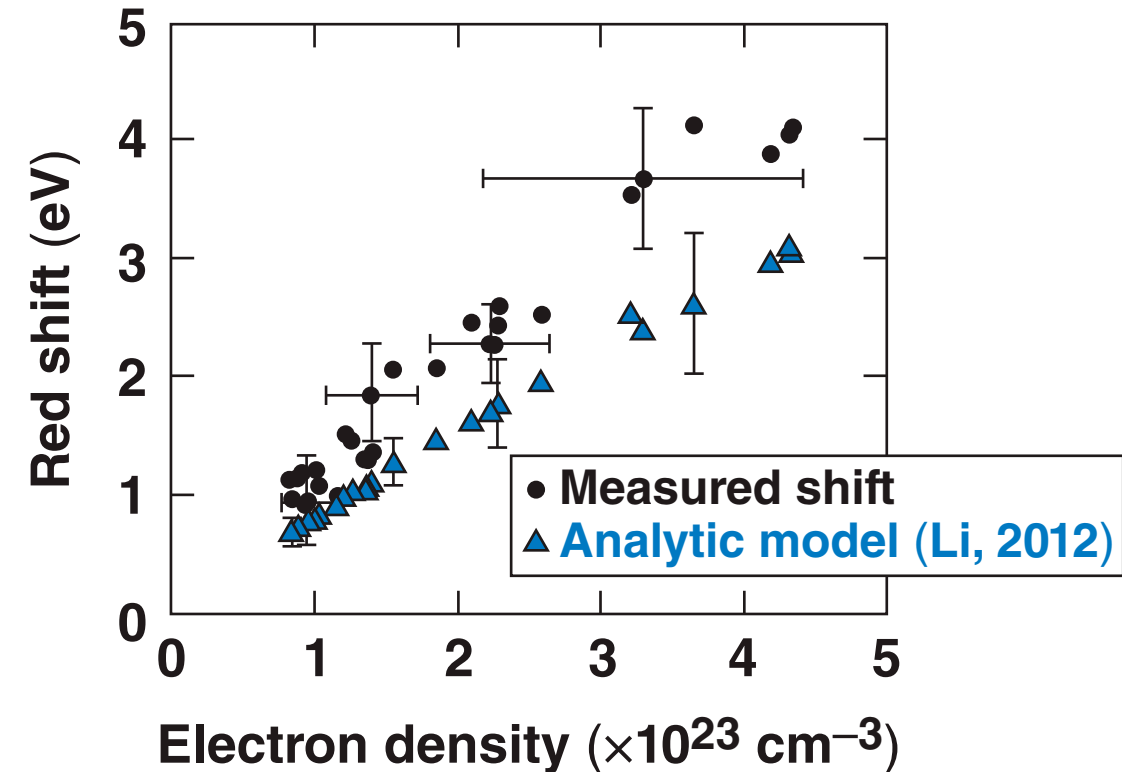
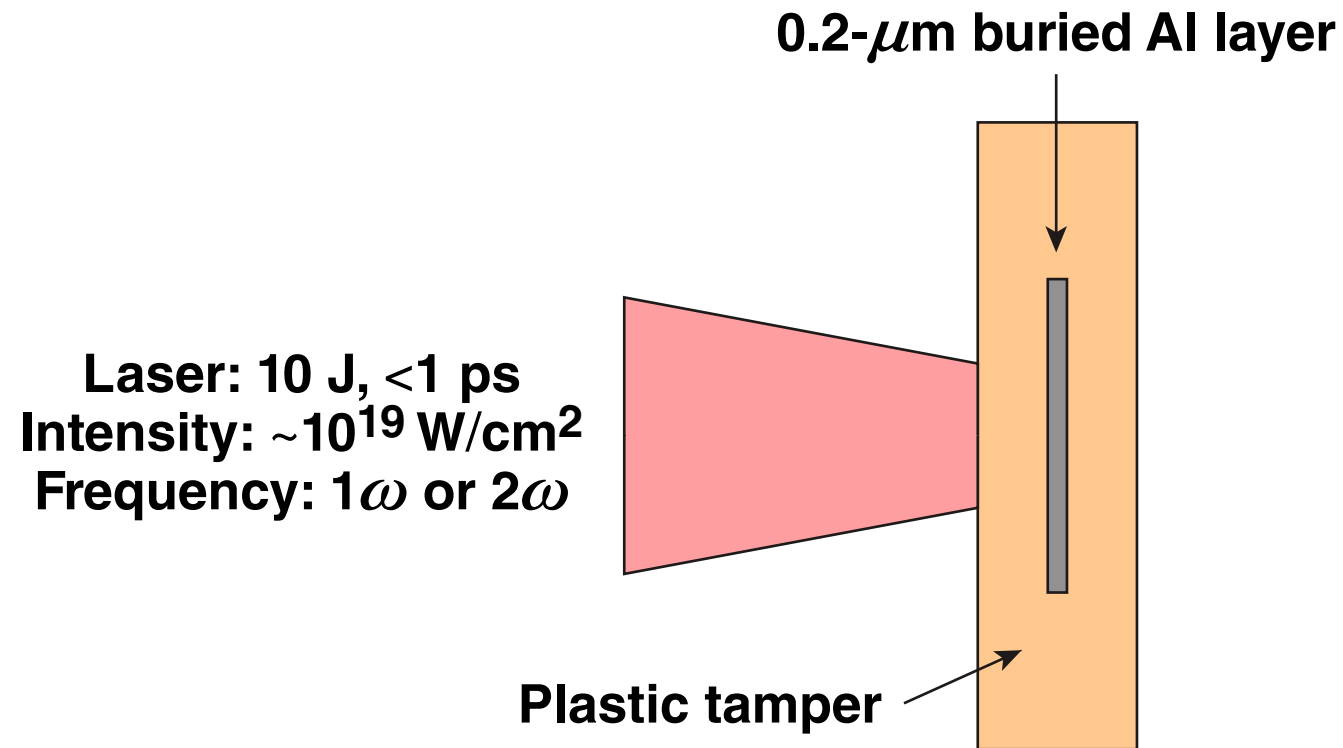


# Picosecond Streaked K-shell Spectroscopy of Near Solid-Density Aluminum Plasmas



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## Summary

A linear red shift of the  $1s2p-1s^2$  transition in He-like aluminum was observed for electron densities between 1 to  $5 \times 10^{23} \text{ cm}^{-3}$



- High-intensity, short-pulse laser interactions have been used to study dense plasma line shifts
- Picosecond x-ray spectroscopy was used to measure the thermal line emission from a buried aluminum tracer layer
- The plasma conditions were inferred from the thermal line width and satellite-intensity ratio using a nonlocal thermodynamic equilibrium (NLTE) collisional-radiative atomic physics model (*PrismSPECT*)\*

The observed line shifts are consistent with an analytic line-shift model\*\* based on numerical ion-sphere calculations for dense plasmas.

\* J. J. MacFarlane *et al.*, High Energy Density Phys. **3**, 181 (2007).

\*\* X. Li and F. B. Rosmej, Europhys. Lett. **99**, 33001 (2012).

# Collaborators

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**P. M. Nilson, S. Ivancic, C. Mileham,  
I. A. Begishev, and D. H. Froula**

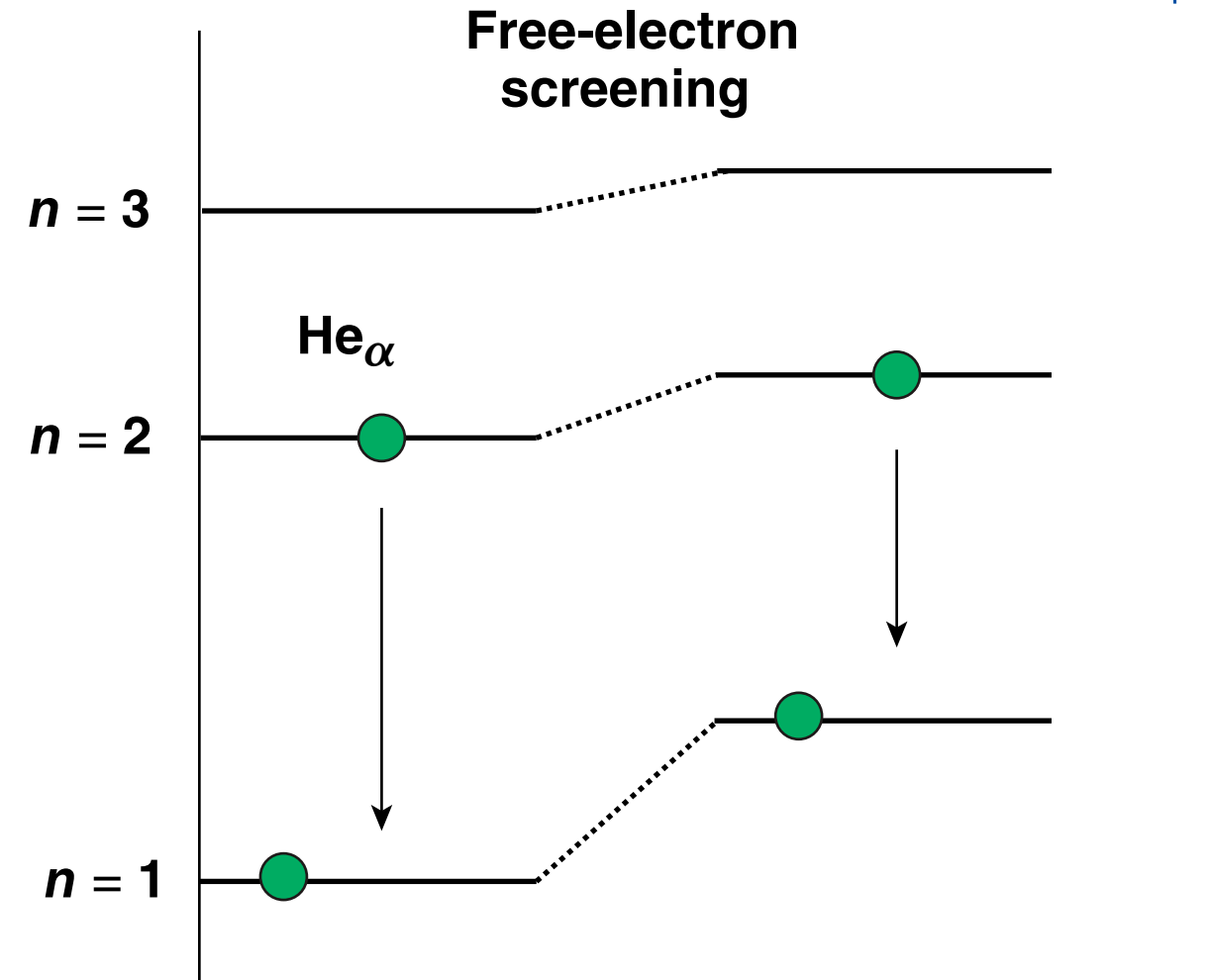
**University of Rochester  
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**I. E. Golovkin**

**Prism Computational Sciences**

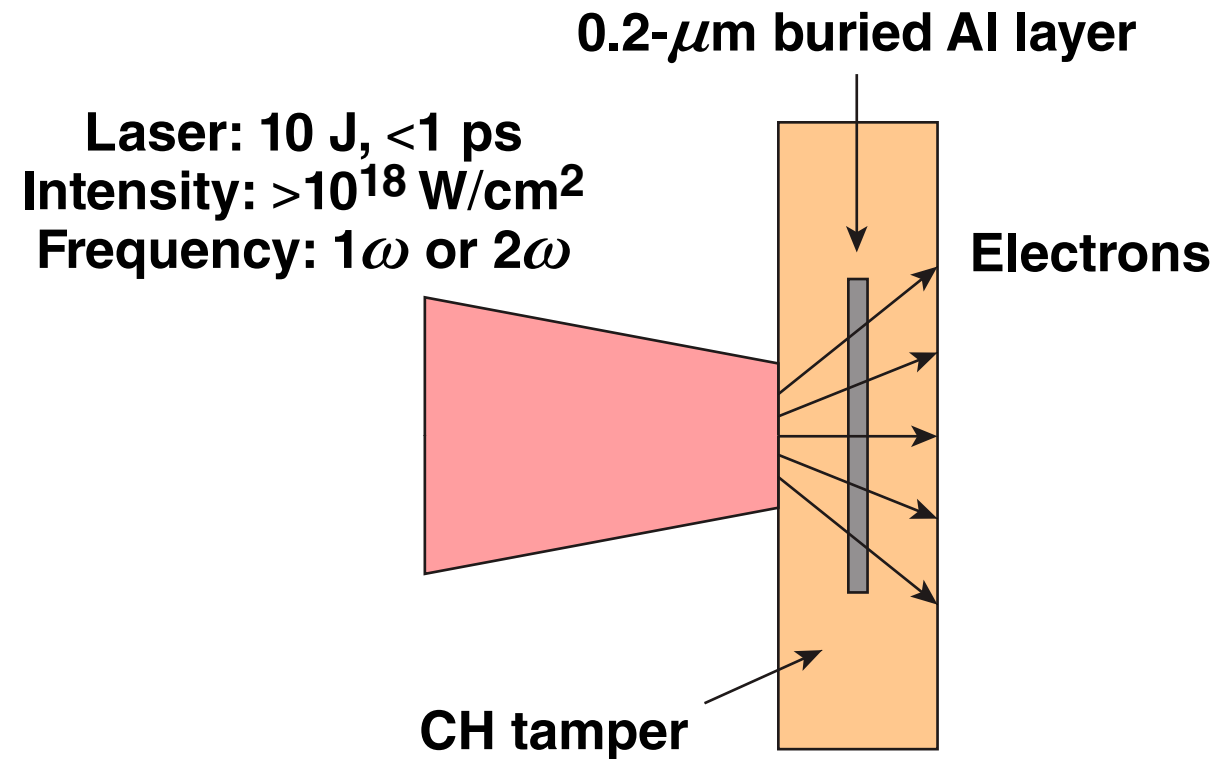
# Accurate descriptions of plasma-dependent atomic properties are required to understand high-energy-density systems

- The radiative and thermodynamic properties of a plasma are modified at high energy density<sup>\*,\*\*</sup>:
  - nonequilibrium equation of state
  - pressure ionization and continuum lowering
  - energy level and spectral line shifts
- Dense plasma line shifts originate from free-electron modification of the ionic potential<sup>\*\*,\*\*†</sup>:
  - free electrons screen the nuclear charge
  - bound energy levels shift toward the continuum
  - emission lines shift to lower photon energies



\*G. Chabrier, F. Douchin, and A. Y. Potekhin, *J. Phys., Condens. Matter* **14**, 9133 (2002).  
\*\*D. Salzmann, *Atomic Physics in Hot Plasmas*, International Series of Monographs on Physics, Vol. 97 (Oxford University Press, New York, 1998).  
†H. R. Griem, *Spectral Line Broadening by Plasmas* (Academic Press, New York, 1974).

# Experiments using buried-layer targets access the dense, high-temperature plasma regime

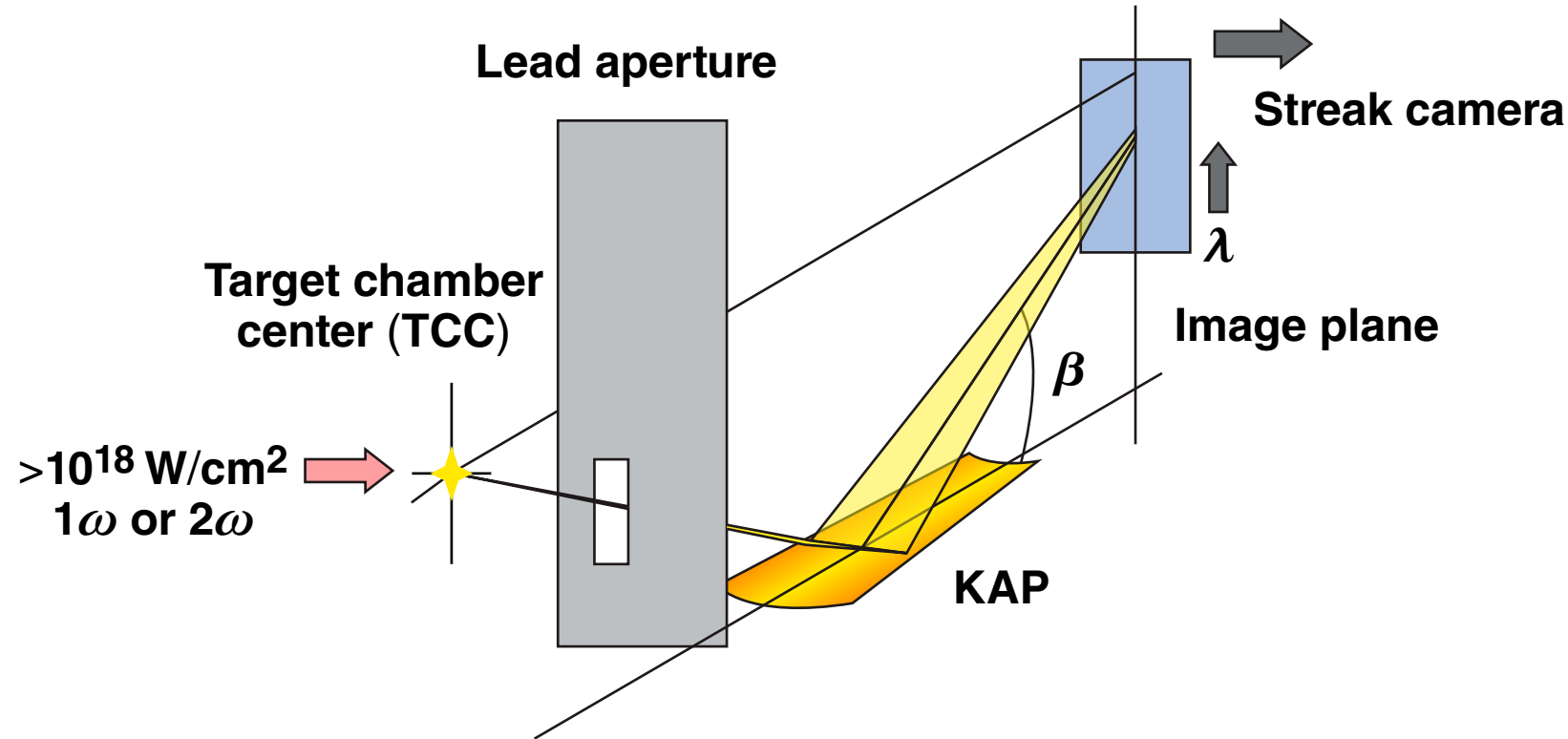


- The buried layer heats through collisional dissipation of a resistive return current
- Buried-layer emission is studied with an ultrafast streaked x-ray spectrometer

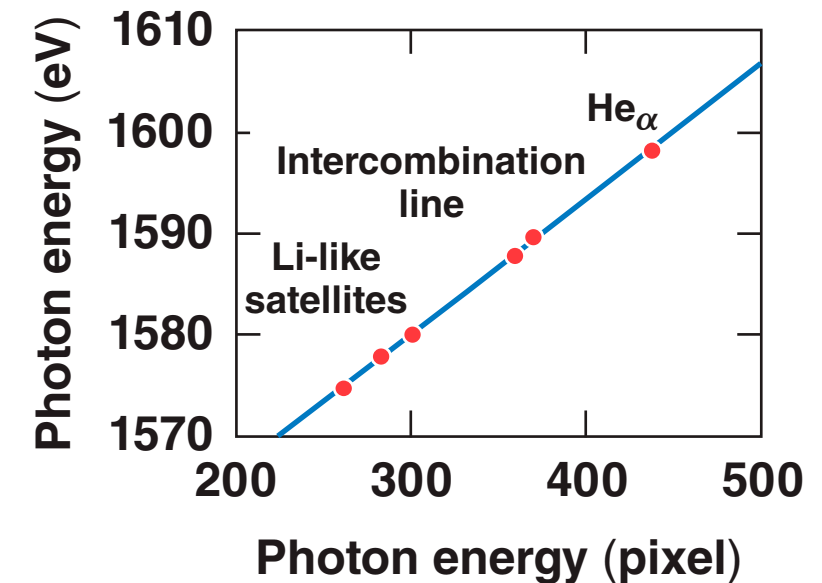
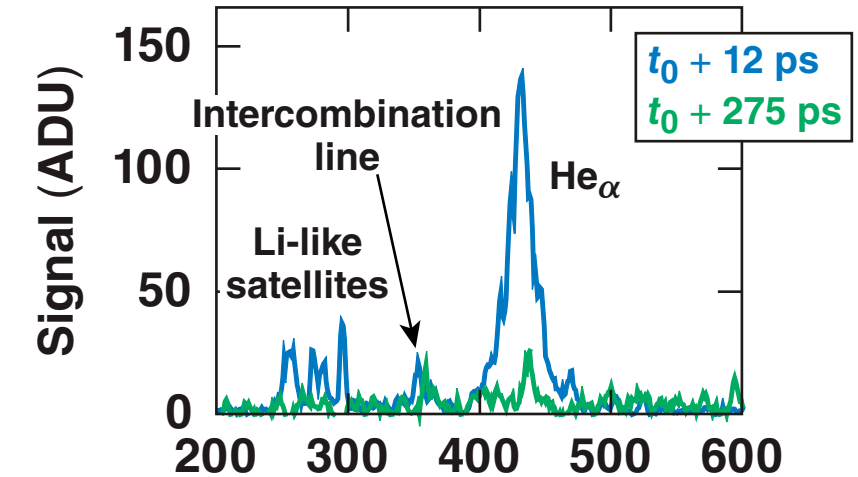
The data are compared to simulated spectra to infer the plasma conditions.

\* C. R. D. Brown *et al.*, Phys. Rev. Lett. **106**, 185003 (2011).  
\*\* D. J. Hoarty *et al.*, High Energy Density Phys. **9**, 661 (2013).

# A focusing, time-resolved Hall spectrometer measured $\text{He}_\alpha$ emission from a buried aluminum layer

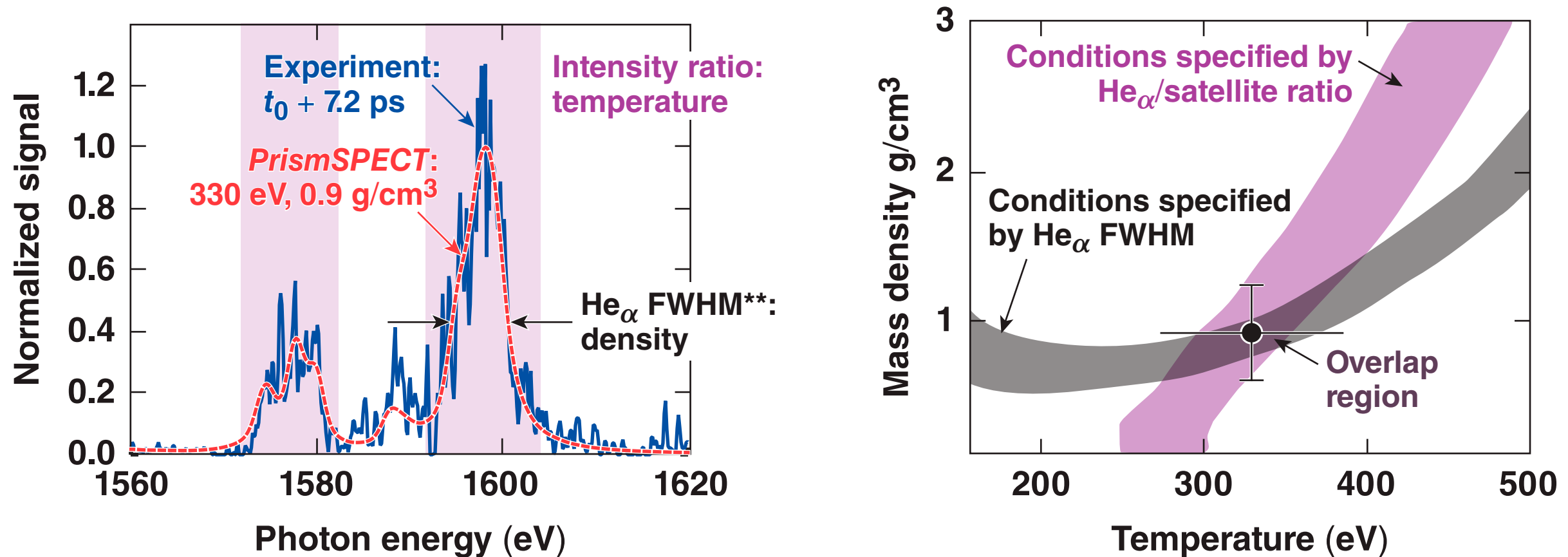


- Conically curved focusing KAP crystal
- Spectral range  $\pm 90 \text{ eV}$  around Al  $\text{He}_\alpha$
- Spectral resolution  $E/\Delta E \sim 800$
- Temporal resolution  $\sim 2 \text{ ps}$



\*T. A. Hall, J. Phys. E: Sci. Instrum. **17**, 110 (1984).

# The instantaneous temperature and density were inferred by comparing with a NLTE collisional-radiative atomic physics model\*

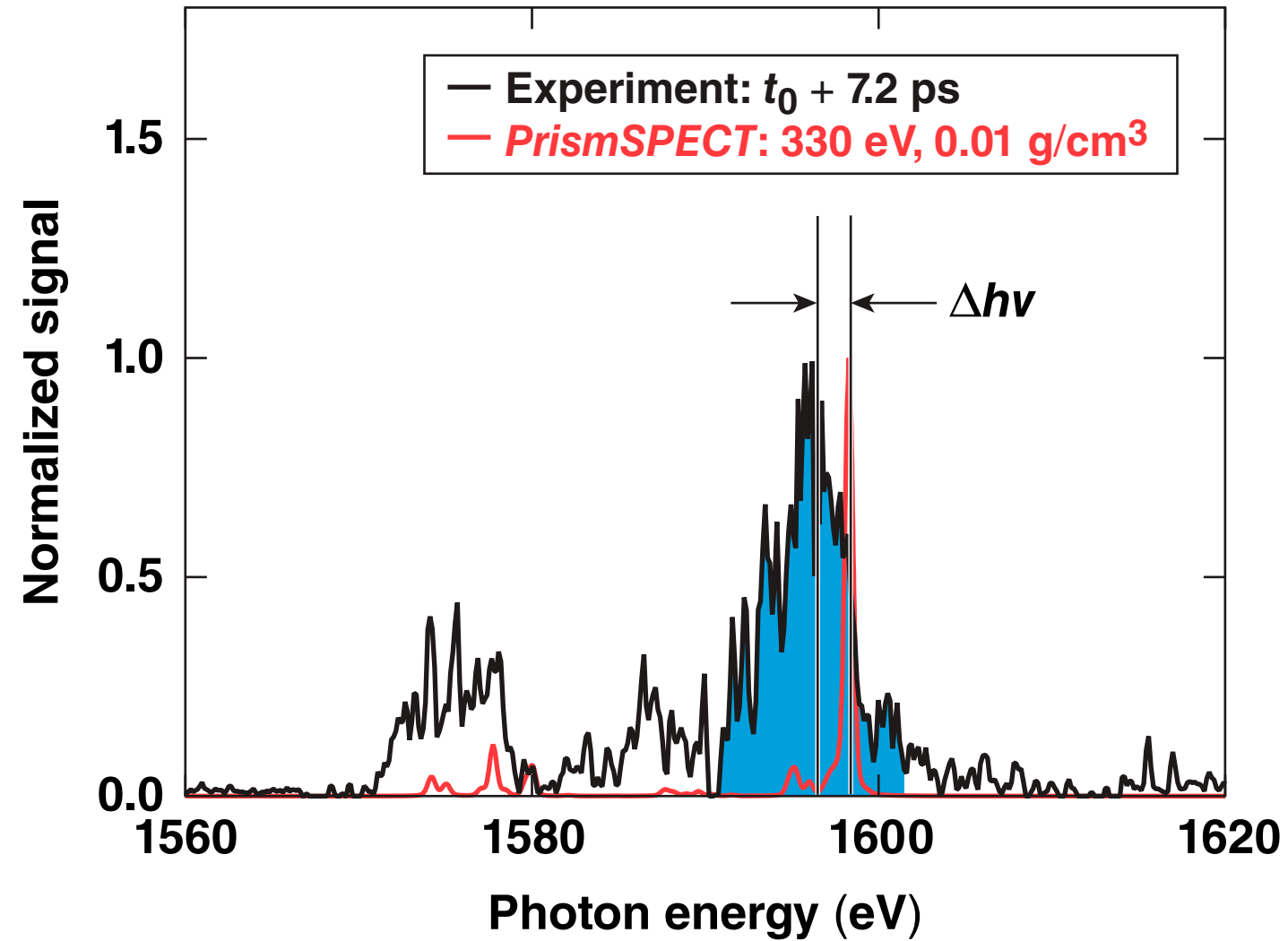


The calculation considers satellite production from Al IX–XIV ions with Doppler, Stark, natural, Auger, and opacity-broadening contributions.

\* J. J. MacFarlane *et al.*, High Energy Density Phys. **3**, 181 (2007).

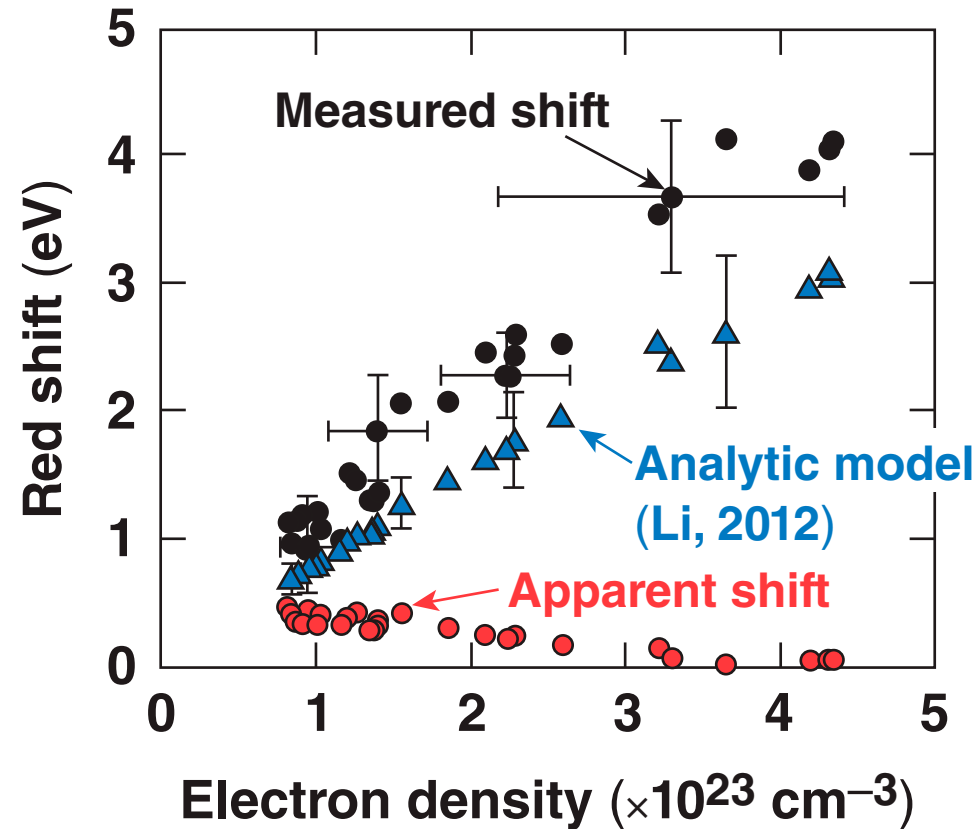
\*\* FWHM: full width at half maximum

# The spectral shifts were quantified by the first moment of the line shape





# A simplified line-shift model based on numerical ion-sphere calculations shows broad agreement with the experimental data



- The analytic model is a parameterization of the self-consistent ion-sphere model (SCFISM)<sup>\*,\*\*</sup> for dense, finite-temperature plasmas
- Apparent shifts calculated from the NLTE atomic model verify that the observed shifts are not spurious

**The line-shift model agrees well at lower densities; there is evidence for deviation at the most extreme conditions.**

<sup>\*</sup> F. J. Rogers, H. C. Graboske, and D. J. Harwood, Phys. Rev. A 1, 1577 (1970).

<sup>\*\*</sup> S. Ichimaru, Rev. Mod. Phys. 54, 1017 (1982).

## Summary/Conclusions

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