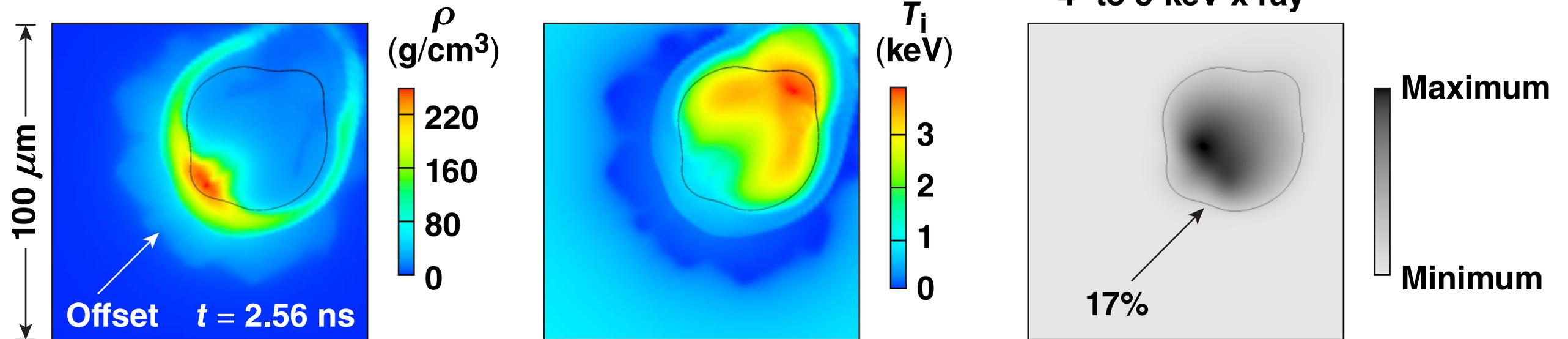


# Numerical Study of Large-Scale, Laser-Induced Nonuniformities in Cryogenic OMEGA Implosions

3-D *ASTER* simulations of shot 78378



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## Summary

# Large-scale, laser-induced nonuniformities can explain performance degradation of cryogenic implosions on OMEGA



- The 3-D hydrodynamic code *ASTER* was developed to study direct-drive implosions
- Simulations consider the OMEGA laser's configuration and include typical target offsets and measured beam imbalance, mispointing, and mistiming
- The effects of the stalk mount were mimicked by applying surrogate perturbations
- Simulations suggest that the implosions are mostly affected by target offset and beam imbalance

# Collaborators

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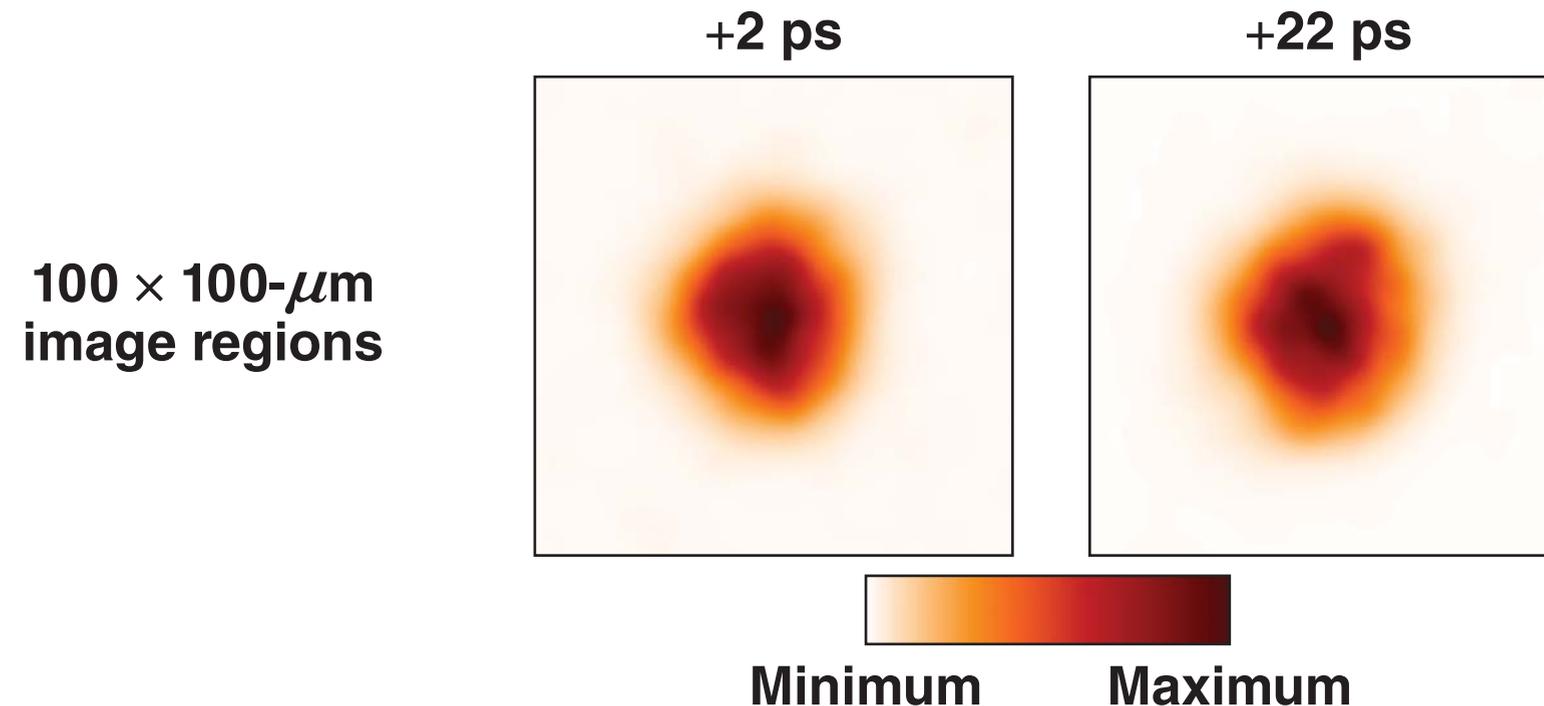
**V. N. Goncharov, F. J. Marshall, K. Silverstein,  
J. P. Knauer, D. H. Froula, and S. P. Regan**

**University of Rochester  
Laboratory for Laser Energetics**

# Short-scale perturbations alone cannot explain the performance of cryogenic implosions on OMEGA

- Implosions with  $\alpha \gtrsim 3.5$  and in-flight aspect ratio (IFAR)  $\lesssim 22$  underperform
- Images of implosion cores evidence the evolution of low- $\ell$ -mode structures

X-ray images in the 4- to 8-keV range (shown in shot 77064)\*



- $T_i$  measurements suggest significant bulk motions in the hot spot\*\*

\*F. J. Marshall *et al.*, UO4.00004, this conference.

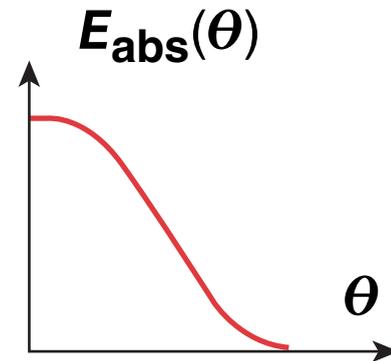
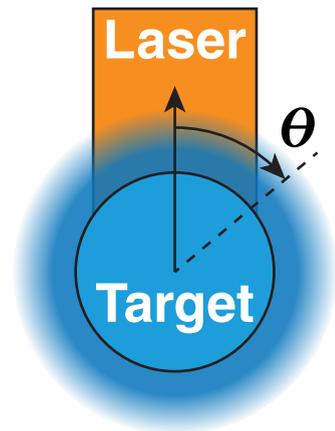
\*\*T. J. Murphy, *Phys. Plasmas* **21**, 072701 (2014).

# The 3-D code *ASTER* studies the effects of large-scale nonuniformities in OMEGA implosions

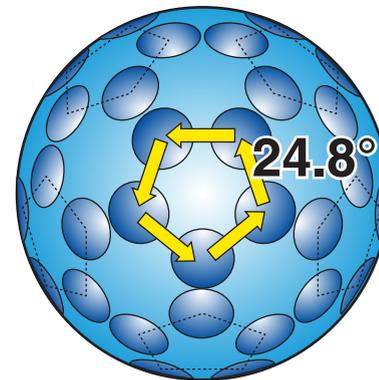
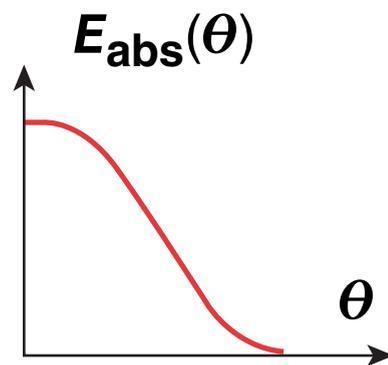


- Based on the Eulerian piecewise-parabolic method (PPM)\*
- Utilizes a two-temperature (ion and electron) fluid model of plasma, which can consist of multiple materials (DT, CH, etc.)
- Implemented on a 3-D orthogonal spherical grid ( $R, \theta, \varphi$ )
- Parallelized using the domain-decomposition approach
- A nominal-resolution ( $700 \times 60 \times 120$ ) implosion simulation takes less than 48 h
- Current physical options
  - tabulated  $Z$  and equation of states (EOS's)
  - Spitzer thermal conduction (with optional flux limitation)
  - simplified 3-D laser-deposition model

# ASTER uses a simplified 3-D laser deposition model



- The assumption of a spherical corona
- Adopts a ray-tracing routine with cross-beam energy transfer (CBET)\* from the 1-D code LILAC\*\*



Perturbations  
of individual  
beams



3-D laser  
deposition

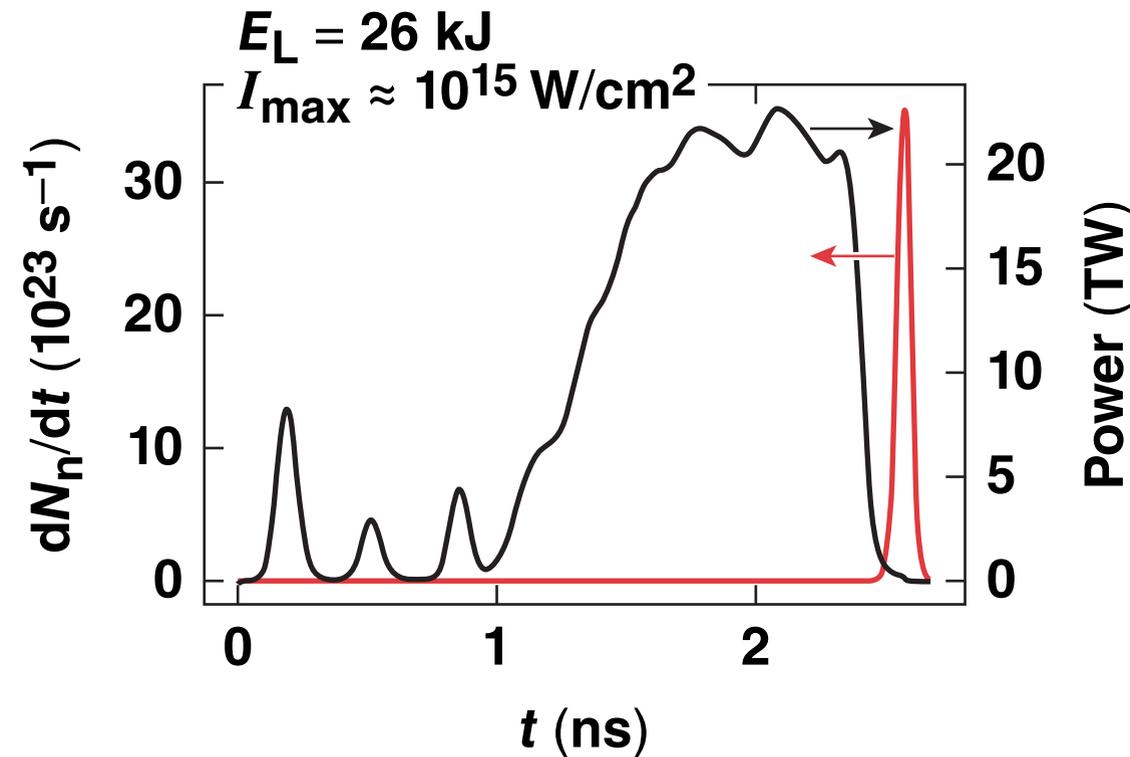
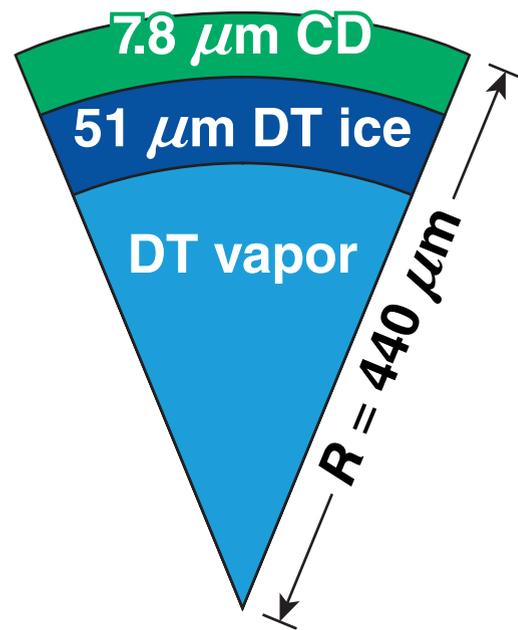
- Simulated effects:
- OMEGA beam overlap
  - beam-energy imbalance
  - beam mispointing
  - beam mistiming
  - target offset

\*I. V. Igumenshchev *et al.*, Phys. Plasmas **17**, 122708 (2010).

\*\*J. Delettrez *et al.*, Phys. Rev. A **36**, 3926 (1987).

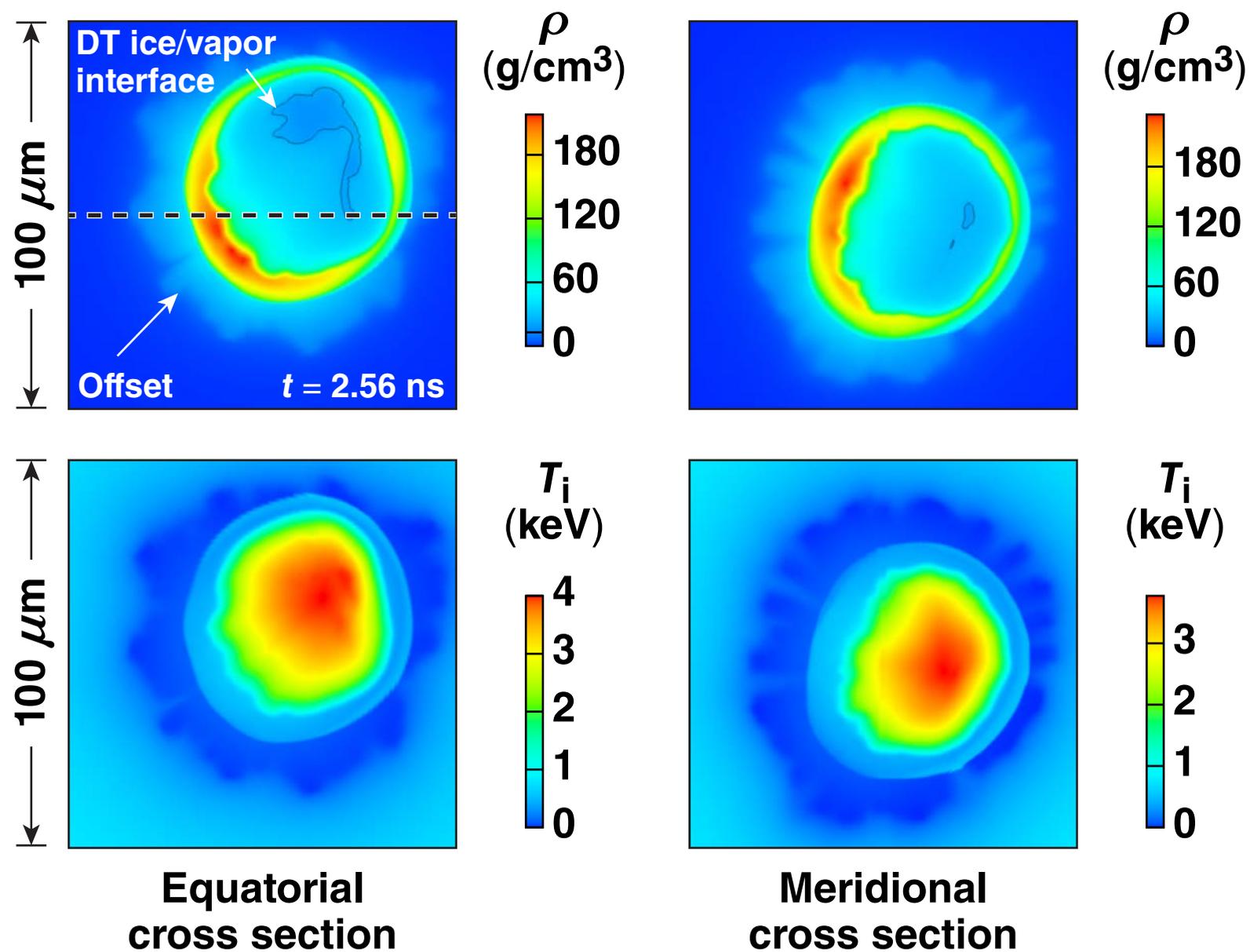
# OMEGA shot 78378 was simulated assuming measured 3-D laser-imposed perturbations

$\alpha = 4$   
IFAR = 22



- Nominal model:
  - Beam overlap
  - 10%  $\sigma_{\text{rms}}$  imbalance
  - 10- $\mu\text{m}$   $\sigma_{\text{rms}}$  mispointing
  - 5-ps  $\sigma_{\text{rms}}$  mistiming
- Typical target offset  $\sim 10$  to  $20 \mu\text{m}$

# Simulations of a nominal model with a 10- $\mu\text{m}$ offset suggest that offset and beam imbalance have the largest effect



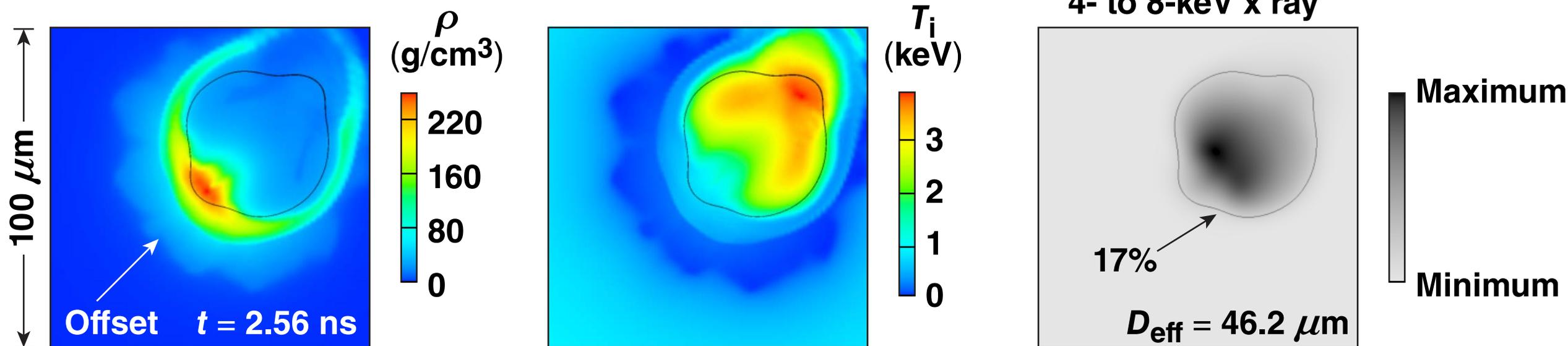
YOU = 0.51

- Larger perturbations are needed to explain OMEGA implosions (YOC ~ 30)

YOU: yield over uniform  
YOC: yield over clean

# A nominal model with a 20- $\mu\text{m}$ offset shows the typical performance of OMEGA $\alpha = 4$ cryogenic implosions

## 3-D ASTER simulations of shot 78378



YOU = 0.22

$P_n = 47 \text{ Gbar}$  (versus  $45 \pm 6 \text{ Gbar}$  for shot 78378)

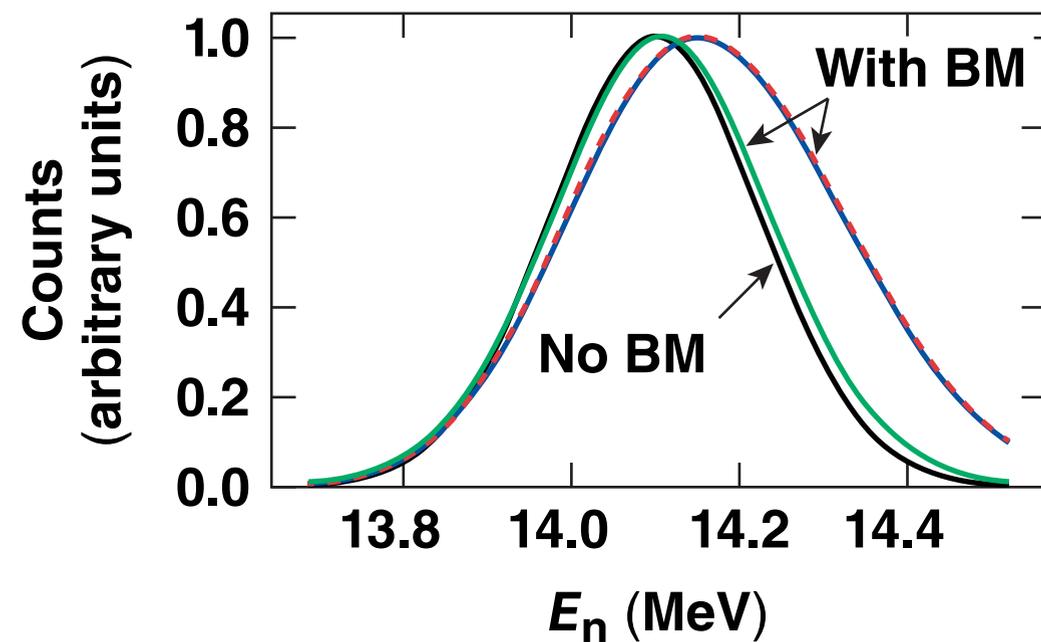
17% contour defines the hot spot size

- Self-emission x-ray images loosely reproduce the hot-spot shape
- The cold shell must be imaged

# Large-scale perturbations increase the variation of inferred $T_i$ , which is in good agreement with measurements

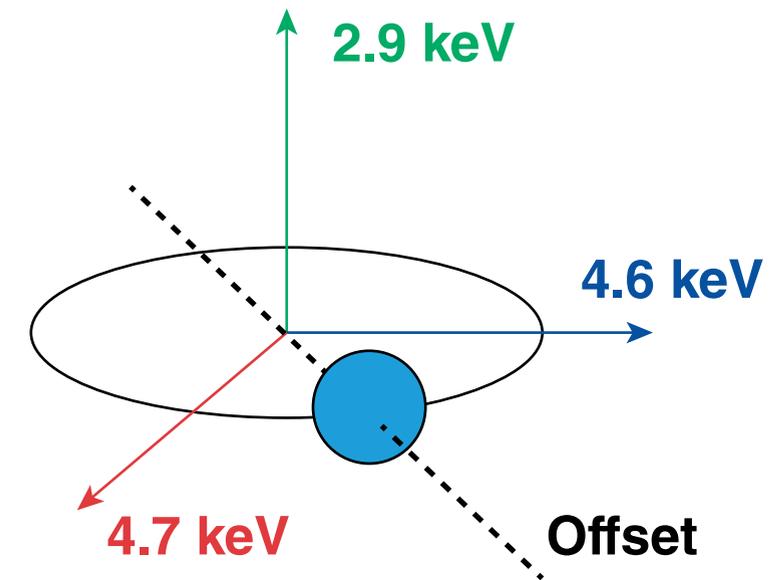
Nominal model with 20- $\mu\text{m}$  offset

Normalized neutron spectra\* with and without the effects of plasma bulk motion (BM)



$$\langle T_i \rangle_n = 2.73 \text{ keV}$$

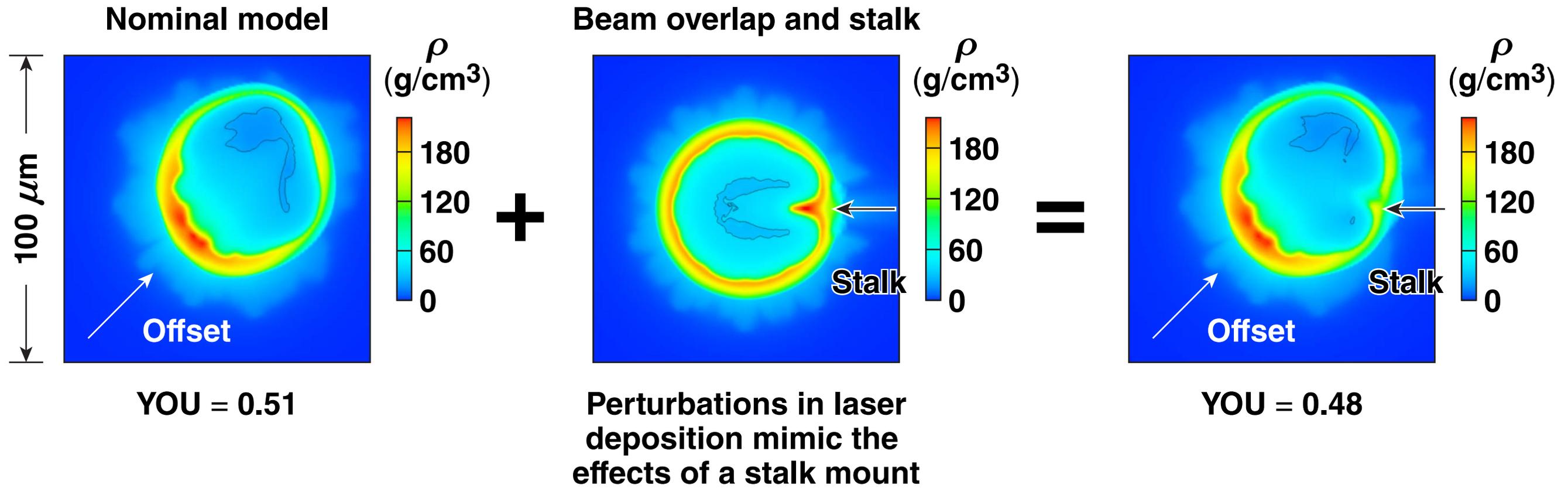
The directions of measurements and inferred  $T_i$



Shot 78378  
Inferred  $T_i$ : 3.6, 3.7, and 4.6 keV

# Perturbations from a mount stalk and other sources interact, reducing the implosion performance

Equatorial density map at neutron peak ( $t = 2.56$  ns)



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