The South-Pole Bang-Time X-ray Diagnostic for the NIF



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Summary

The south-pole bang-time has been designed to provide an x-ray bang-time measurement with 30-ps accuracy on the NIF

- An accurate bang time for NIF implosions is essential for validating implosion modeling and tuning performance
- The south-pole bang-time diagnostic* has
 - a view straight up along the NIF's south pole into the LEH
 - four high-bandwidth channels with different filtration to cover a range of THD and DT shots
 - HOPG monochromators to maximize signal:background
- The SPBT analysis combines deconvolution and forward fitting to determine the bang-time signal
- The SPBT is scheduled to be deployed in December 2010

NIF

^{*}A. G. MacPhee *et al.*, "A Diamond Detector for Quasi Monochromatic ICF X-Ray Bang-time Measurement," submitted to Journal of Instrumentation.

The south-pole bang-time (SPBT) diagnostic is a collaborative effort between LLE and NIF/LLNL

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The SPBT diagnostic is needed to measure the bang time of NIF implosions with ~30-ps accuracy

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- NIF x-ray bang times from GXD vary consistently ~300 ps from DANTE x-ray-drive—based predictions
- The discrepancy is significantly more than the uncertainties
- Is this
 - a hohlraum low-drive issue
 - a capsule hydrodynamics issue
 - or a measurement issue?

Need to verify the bang-time measurements within 30 ps using another diagnostic.

The SPBT is designed to have a high-bandwidth, low-background signal to achieve 30-ps accuracy

- SPBT is not DIM-mounted and has a permanent dedicated fixture
- SPBT views straight up the south pole and sees little of the hohlraum wall
 - should have low background
- High-bandwidth scopes, cables, and connectors are used
- Optical fibers will be used to connect to NIF mezzanine to reduce cable effects



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The signal-to-background levels are maximized using HOPG monochromators to select an appropriate energy band



- grade ZYB: ~1- μ m crystallites, parallel to within ~0.8°
- Reflects x-rays that satisfy the Bragg condition: 11 keV
 - predictions indicate that the capsule bang-time x-ray spectrum peaks in the 9- to 12-keV range
- Four PCD channels with different filtration on each channel to cover a range of THD and DT shots



The expected SPBT signal is modeled by applying instrument response functions (IRF's) to the predicted x-ray source



• X-ray power *P*_{x-ray}(*t*) includes

- detector solid angle
- spectral IRF
- modeled x-ray spectral fluence
- integrated over spectrum
- Detector signal S(t) includes
 - convolution of temporal IRF with $P_{x-ray}(t)$
 - nonlinearity of bias circuit
 - attenuators
 - simulated noise

The bang-time analysis uses a forward fit of test P(t) to the observed signal.



The x-ray background signal during the laser pulse is fit by deconvolution with Tikhonov regularization*

- Tikhonov regularization with a derivative operator enforces smoothness
- This allows one to determine the background signal driven by the laser pulse (LPI, etc.) without any assumptions as to how it varies with time



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*A. N. Tikhonov and V. Y. Arsenin, *Solutions of Ill-Posed Problems* (Winston, Washington, DC, 1977).

The bang-time signal is determined by forward fitting with a decaying background and a skewed Gaussian peak

- Combination with deconvolution results provides a good reconstruction of both the x-ray background and the bang-time peak even in the presence of noise
 - 5- to 10-ps accuracy in the data fitting
 - other instrumental timing issues give
 ~30-ps total accuracy



Summary/Conclusions

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