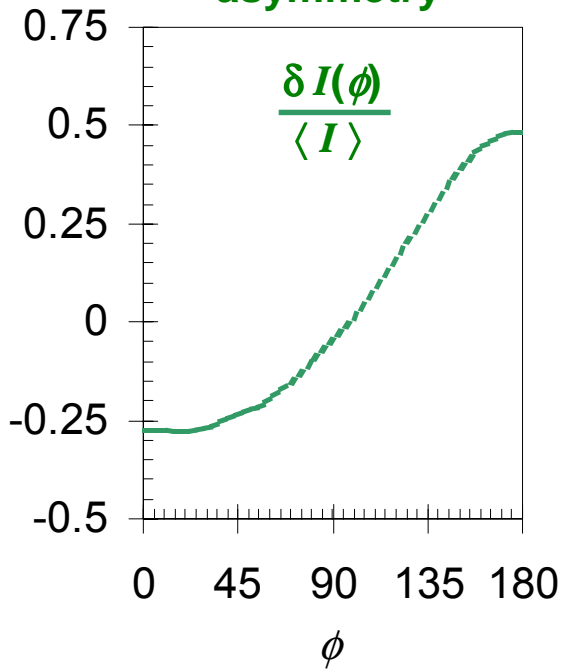


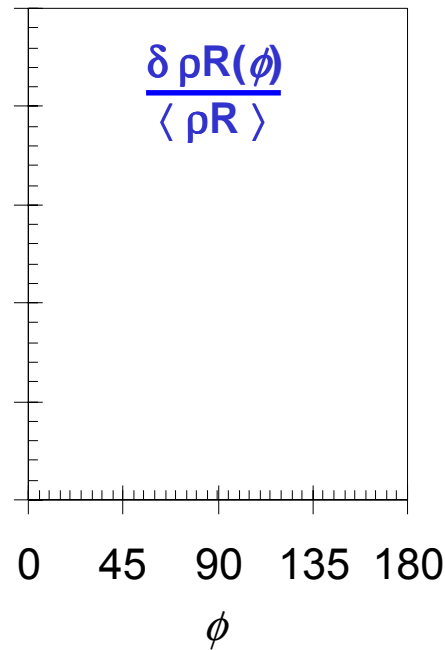
# Time evolution of $\rho R$ asymmetries in OMEGA direct-drive implosions

Laser drive asymmetry

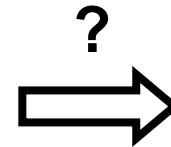
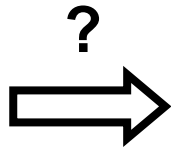
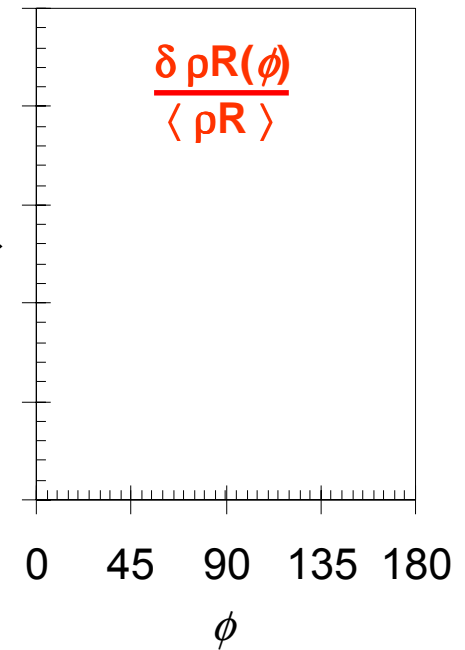


Angle wrt offset direction

Areal density at shock time (~1.7 ns)

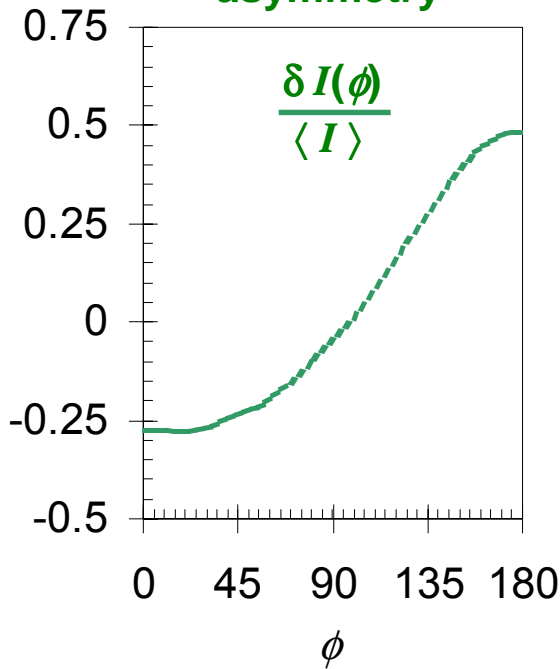


Areal density at bang time (~2.1 ns)

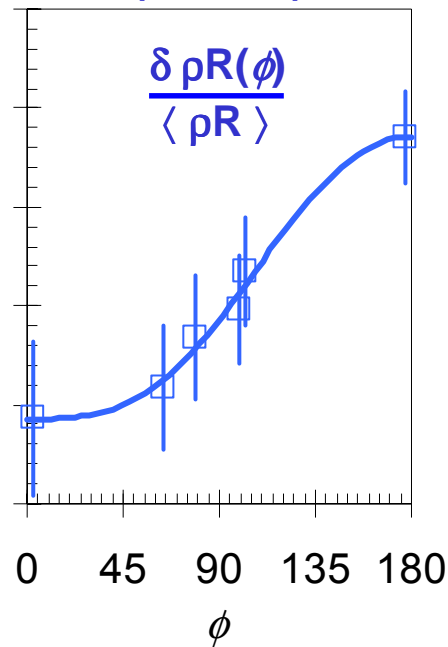


# Time evolution of $\rho R$ asymmetries in OMEGA direct-drive implosions

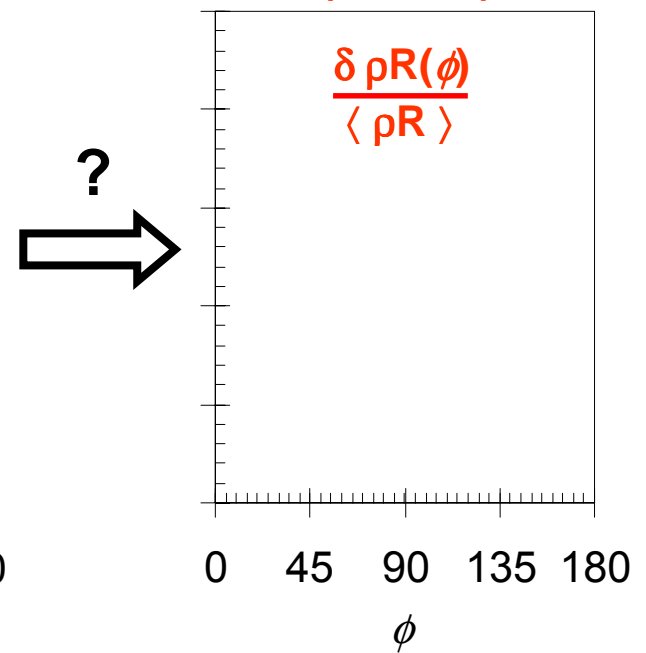
Laser drive asymmetry



Areal density at shock time (~1.7 ns)

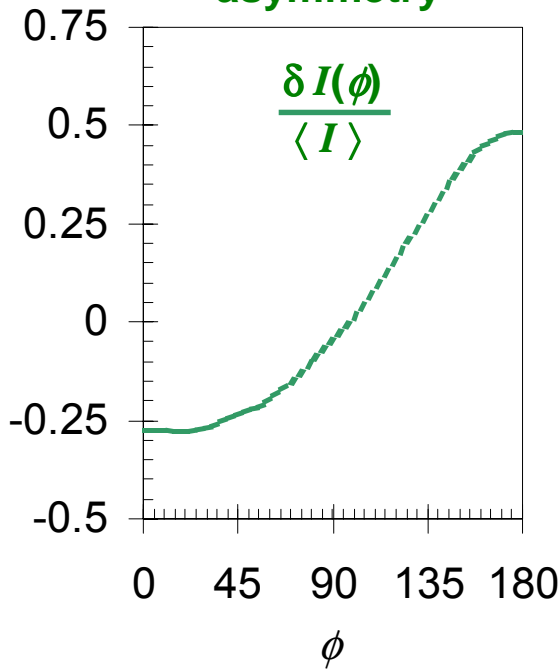


Areal density at bang time (~2.1 ns)

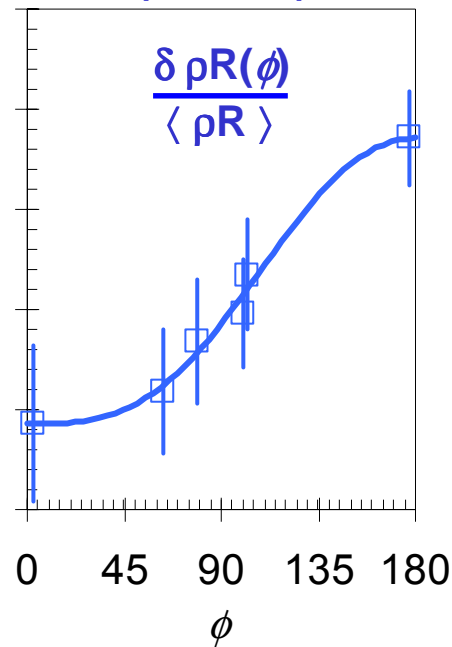


# Time evolution of $\rho R$ asymmetries in OMEGA direct-drive implosions

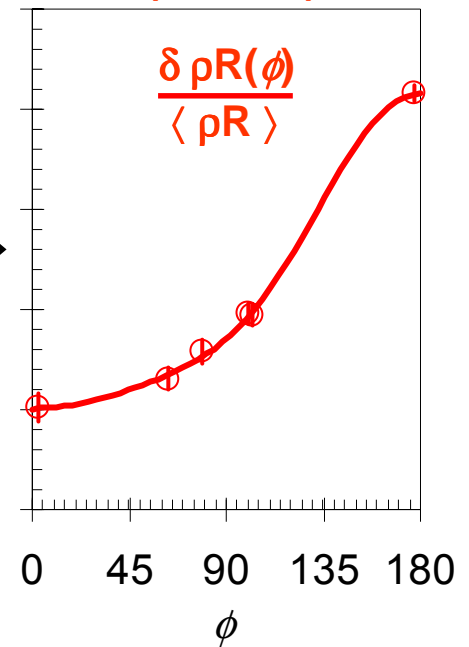
Laser drive asymmetry



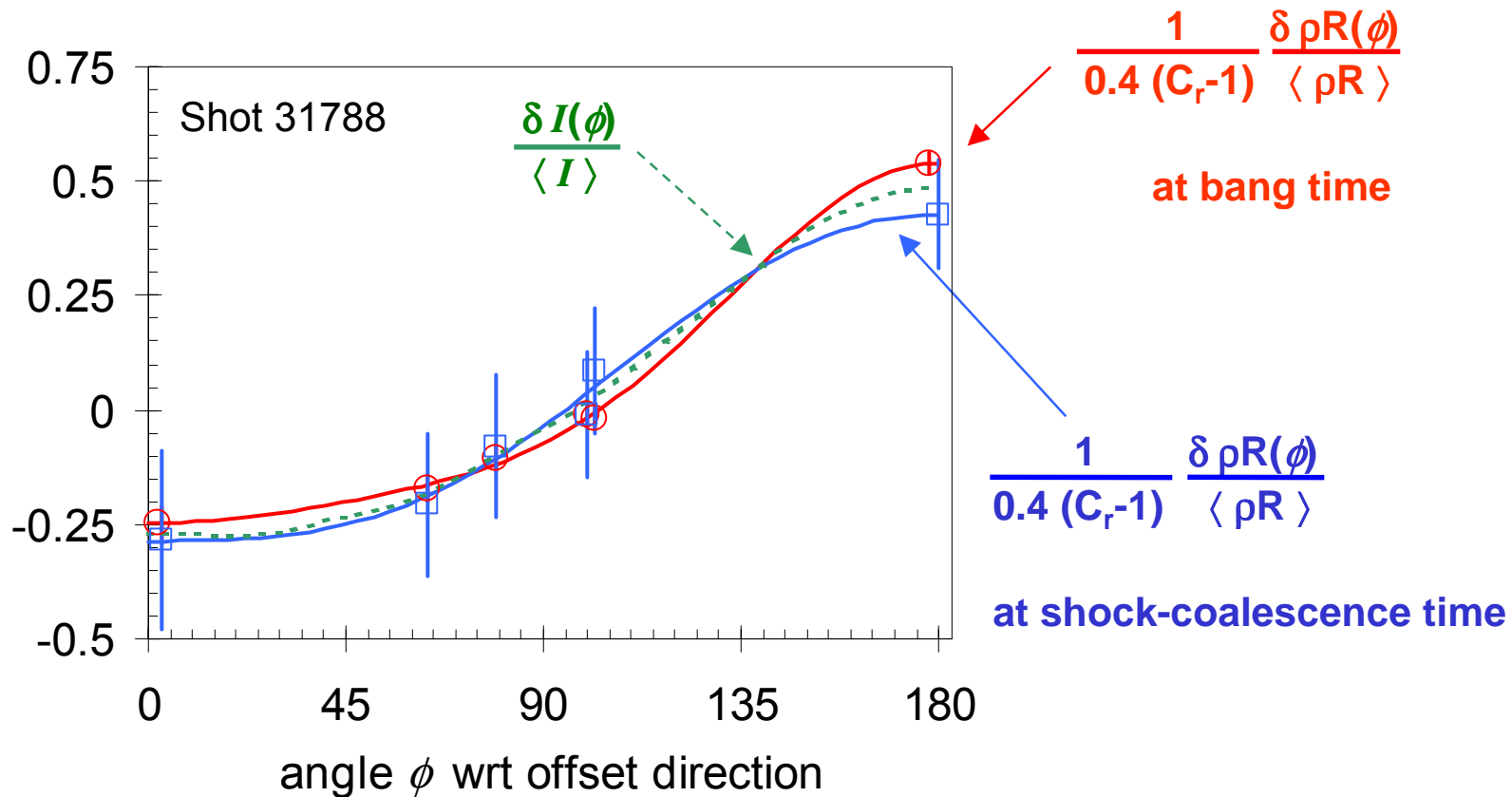
Areal density at shock time (~1.7 ns)



Areal density at bang time (~2.1 ns)



# Time evolution of $\rho R$ asymmetries in OMEGA direct-drive implosions



# Summary

---

For asymmetric OMEGA laser drive  $I(\phi)$  dominated by mode numbers  $\lesssim 3$ , applied to room-temperature capsules with thick CH shells,

- $\delta I(\phi) = I(\phi) - \langle I \rangle$  produces  $\delta \rho R(\phi) = \rho R(\phi) - \langle \rho R \rangle$  with the same shape;
- $\delta \rho R(\phi)$  maintains that shape throughout the implosion ...
- with amplitude depending primarily on the radial convergence ratio  $C_r$ :

$$\frac{\delta \rho R(\phi)}{\langle \rho R \rangle} \approx 0.4 (C_r - 1) \frac{\delta I(\phi)}{\langle I \rangle}$$

- Modes 1 and 2 grow at the same rate, with no phase inversions.

# Collaborators

---

*M.I.T.  
Plasma Science  
and Fusion Center*

J.R. Rygg  
J.A. Frenje  
C.K. Li  
R.D. Petrasso\*

*University of Rochester  
Laboratory for  
Laser Energetics*

J.A. Delettrez  
J.M. Soures  
V.Yu. Glebov  
V. Goncharov  
J. Knauer  
D.D. Meyerhofer  
T.C. Sangster  
R.L. Keck  
P.W. McKenty  
F.J. Marshall  
V. Smalyuk

*Lawrence Livermore  
National Laboratory*

S. Hatchett

\*Visiting Scientist, LLE

# This work is a logical extension of previous work that showed correlations between drive asymmetry and $\rho R$ asymmetry

---

- Small changes in  $I(\phi)$  result in changes in  $\rho R(\phi)^*$
- $\delta \rho R$  growth due to Bell-Plesset-like convergence effects should lead at bang time to\*\*

$$\frac{\langle \delta \rho R \rangle_{rms}}{\langle \rho R \rangle} = K (C_r - 1) \frac{\langle \delta I \rangle_{rms}}{\langle I \rangle}$$

- Data for low modes involving room-temperature and cryo  $D_2$  capsules at OMEGA were roughly consistent with this growth\*\* if

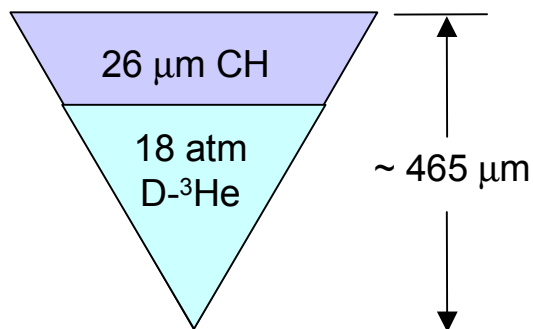
$$K \sim \frac{1}{2}$$

\*F.H. Séguin *et al.*, Phys. Plasmas 9, 3558 (2002).

\*\*C.K. Li *et al.*, submitted to Phys. Rev. Lett.

# New experiments have been performed for controlled drive asymmetries and accurate study of $\rho R(\phi)$ at different times

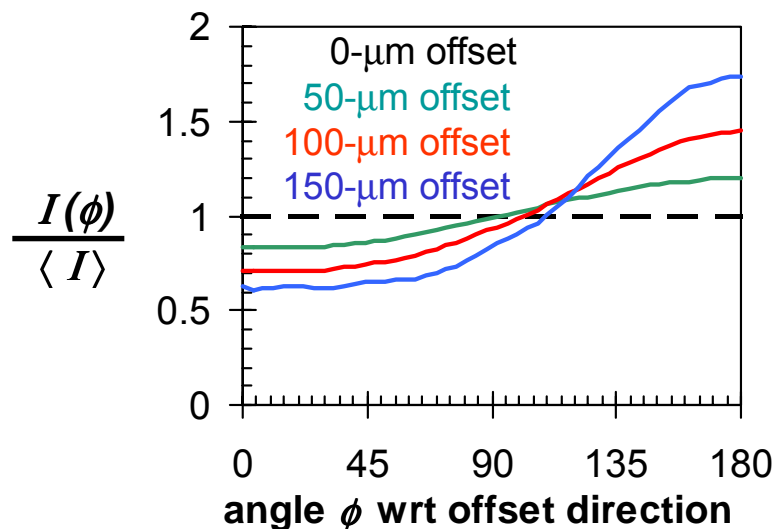
Room temperature capsules:



60-Beam OMEGA laser:

Pulse shape: 1-ns square  
Beam smoothing: 2D-SSD + PS  
On-target energy:  $\sim 23$  kJ

Different  $I(\phi)$  were generated by offsetting capsules from Target Chamber Center:

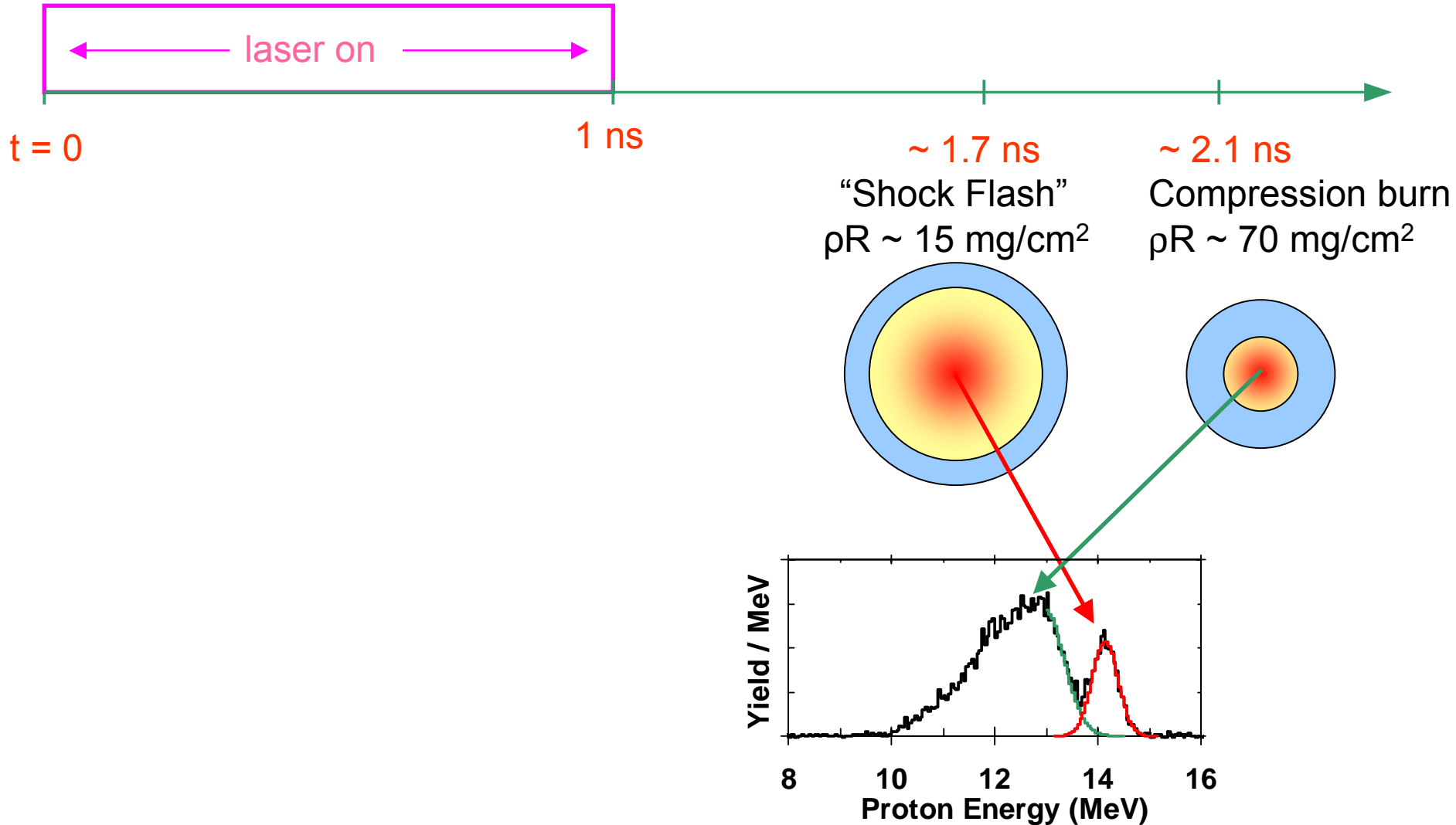


$\rho R(\phi)$  was inferred from the energy lost by  $\text{D}^3\text{He}$  protons leaving at different  $\phi$ :

Proton spectra were measured by 6 WRF proton spectrometers



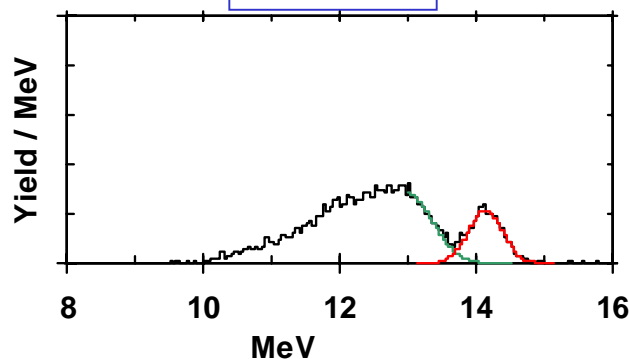
# There are two distinct time intervals during a D<sup>3</sup>He-capsule implosion when D<sup>3</sup>He protons are generated\*



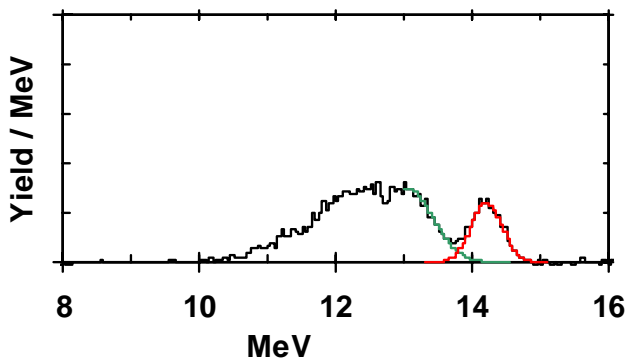
\*Petrasso *et al.*, Phys. Rev. Lett. **90**, 095002-1 (2003).

With no offset,  
all spectra are similar and  $\rho R$  is nearly independent of angle

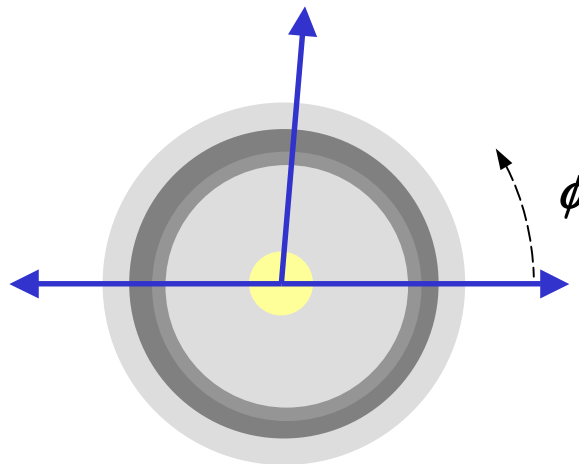
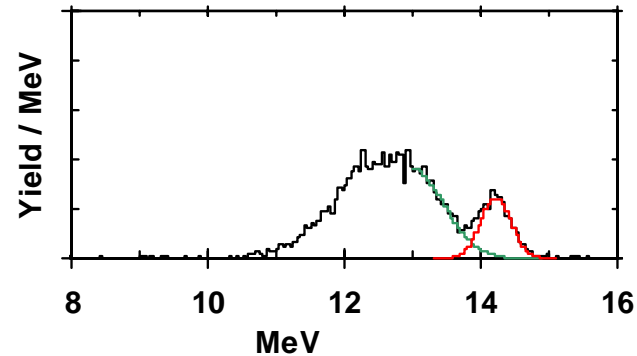
$\phi = 79^\circ$



$\phi = 180^\circ$

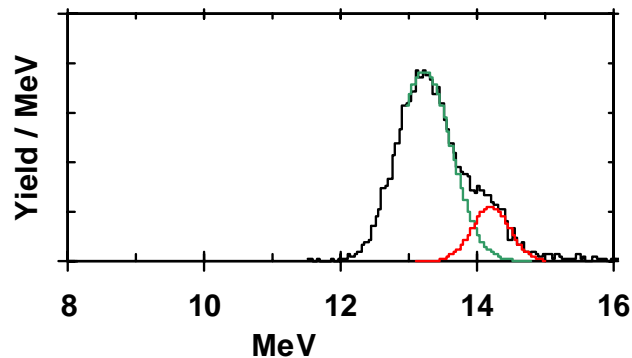


$\phi = 0^\circ$

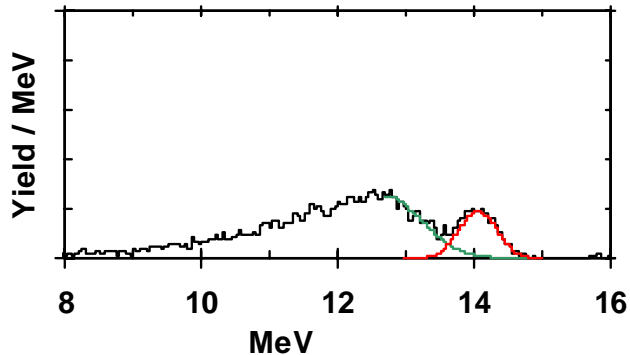


# With a 50- $\mu\text{m}$ offset, spectra (and $\rho R$ ) are different at different angles

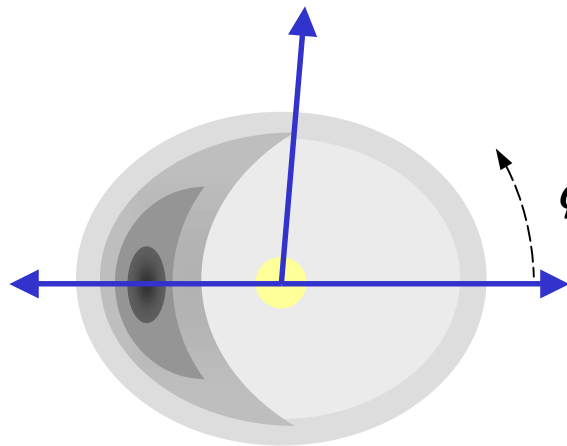
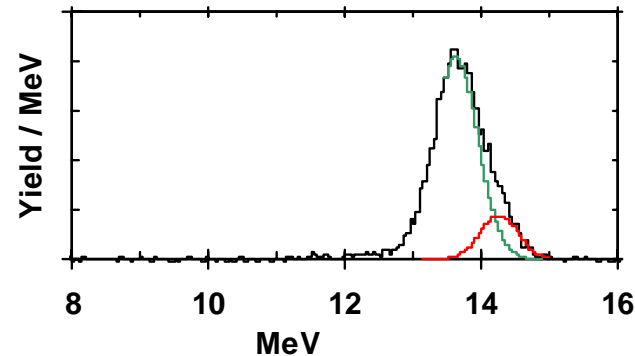
$\phi = 79^\circ$



$\phi = 180^\circ$



$\phi = 0^\circ$

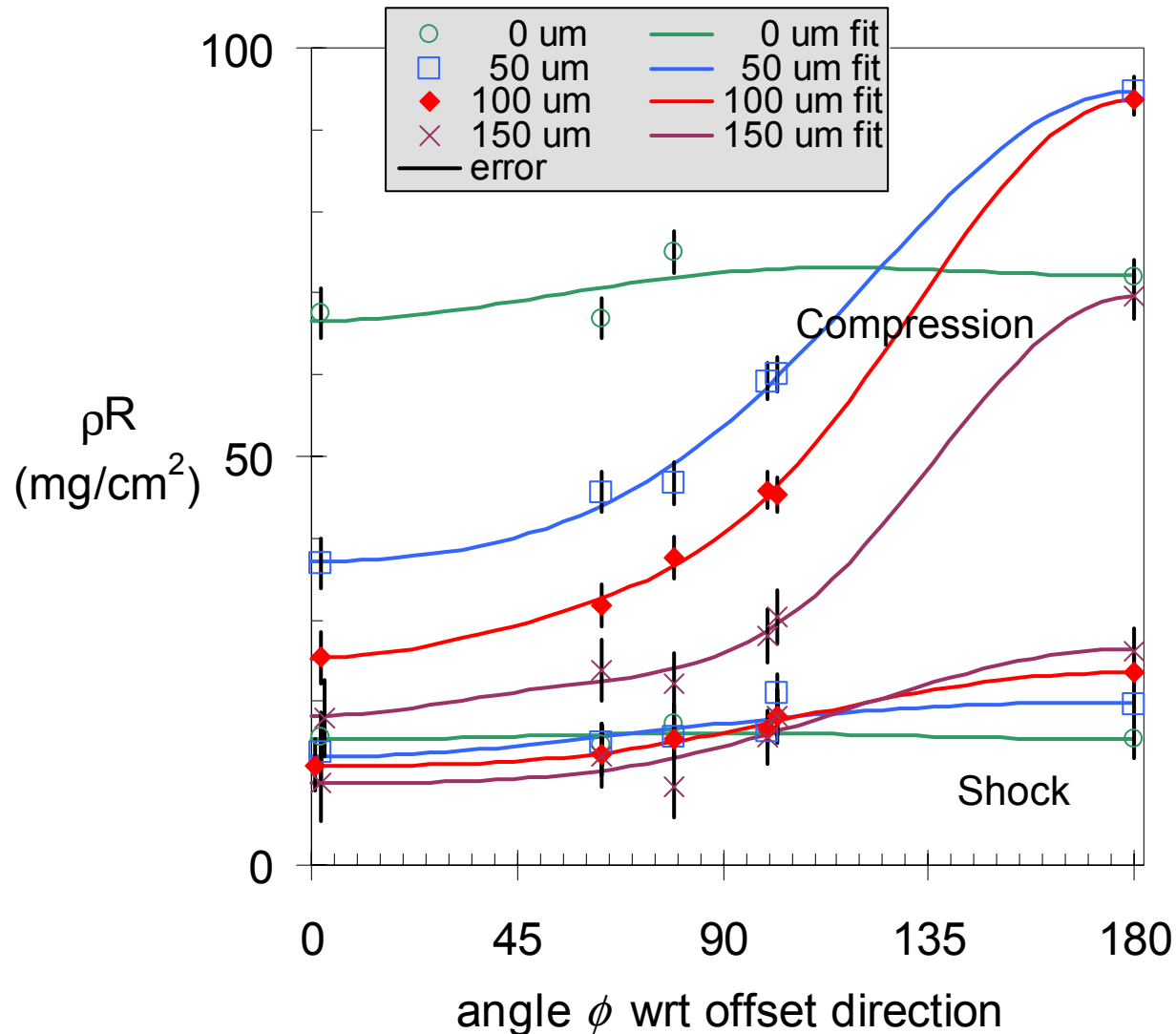


Direction of offset

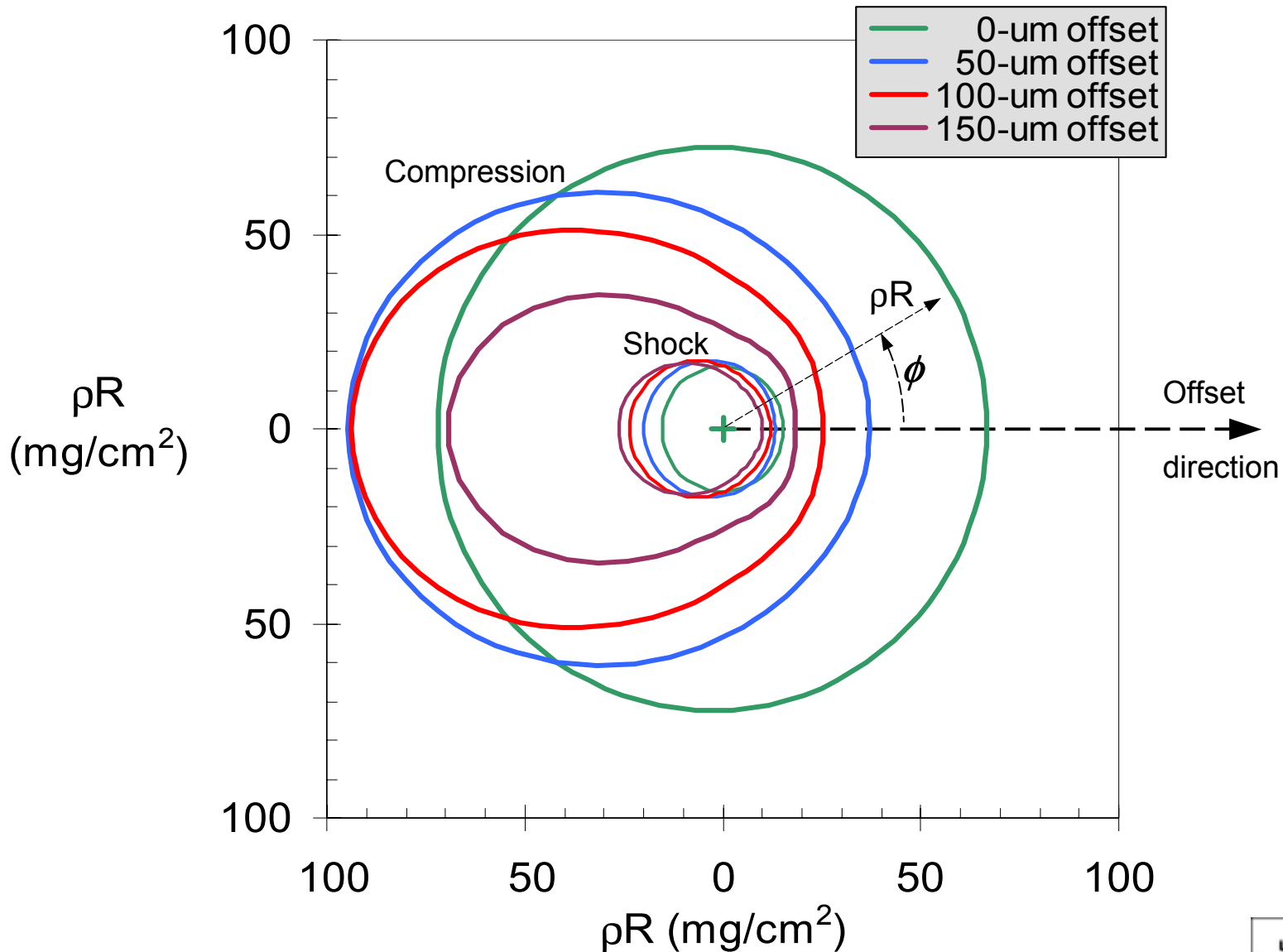
$\rho R$  measurements can be

plotted and fit with

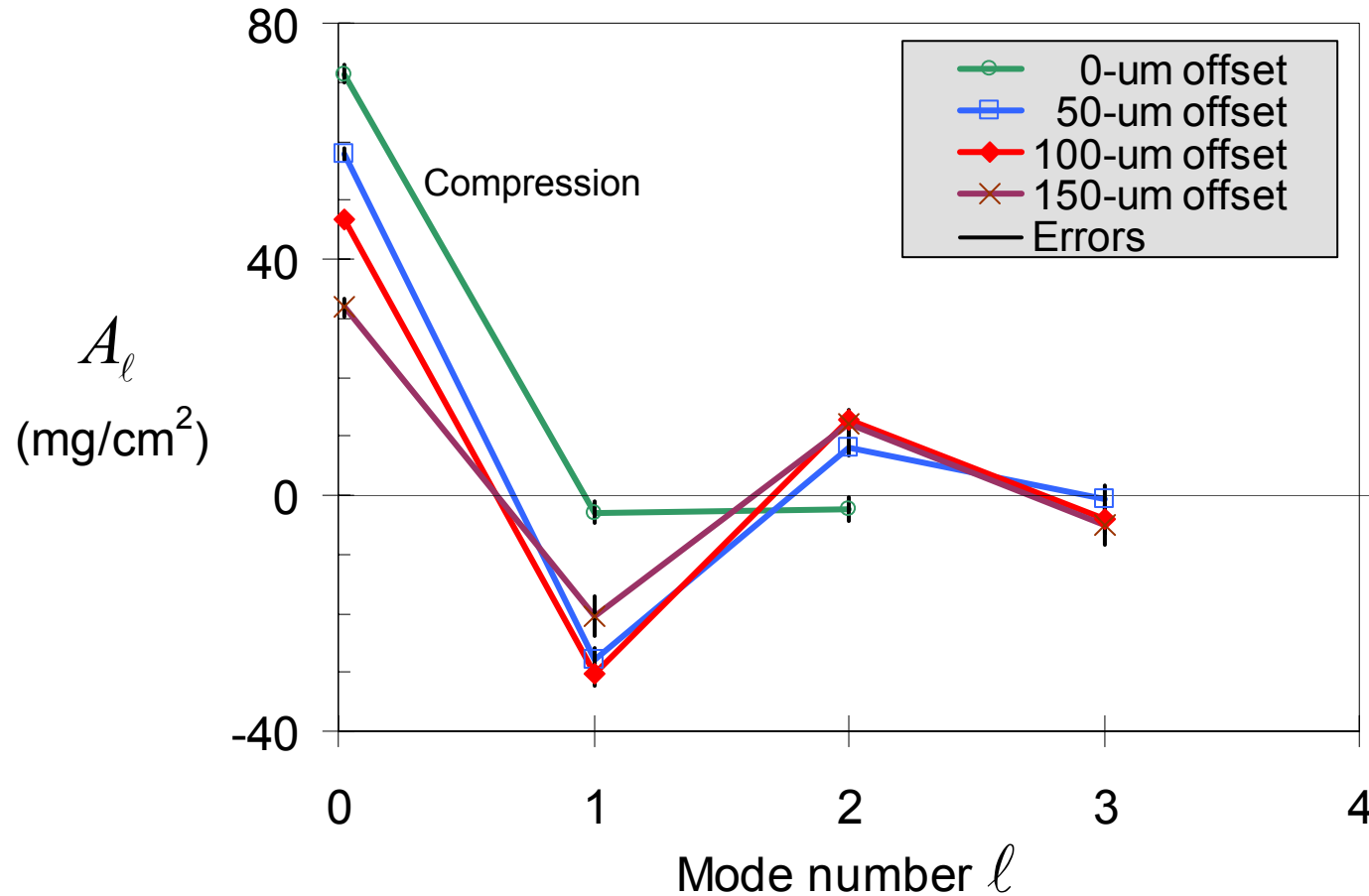
$$\rho R(\phi) = \sum_{\ell} A_{\ell} P_{\ell}(\cos(\phi))$$



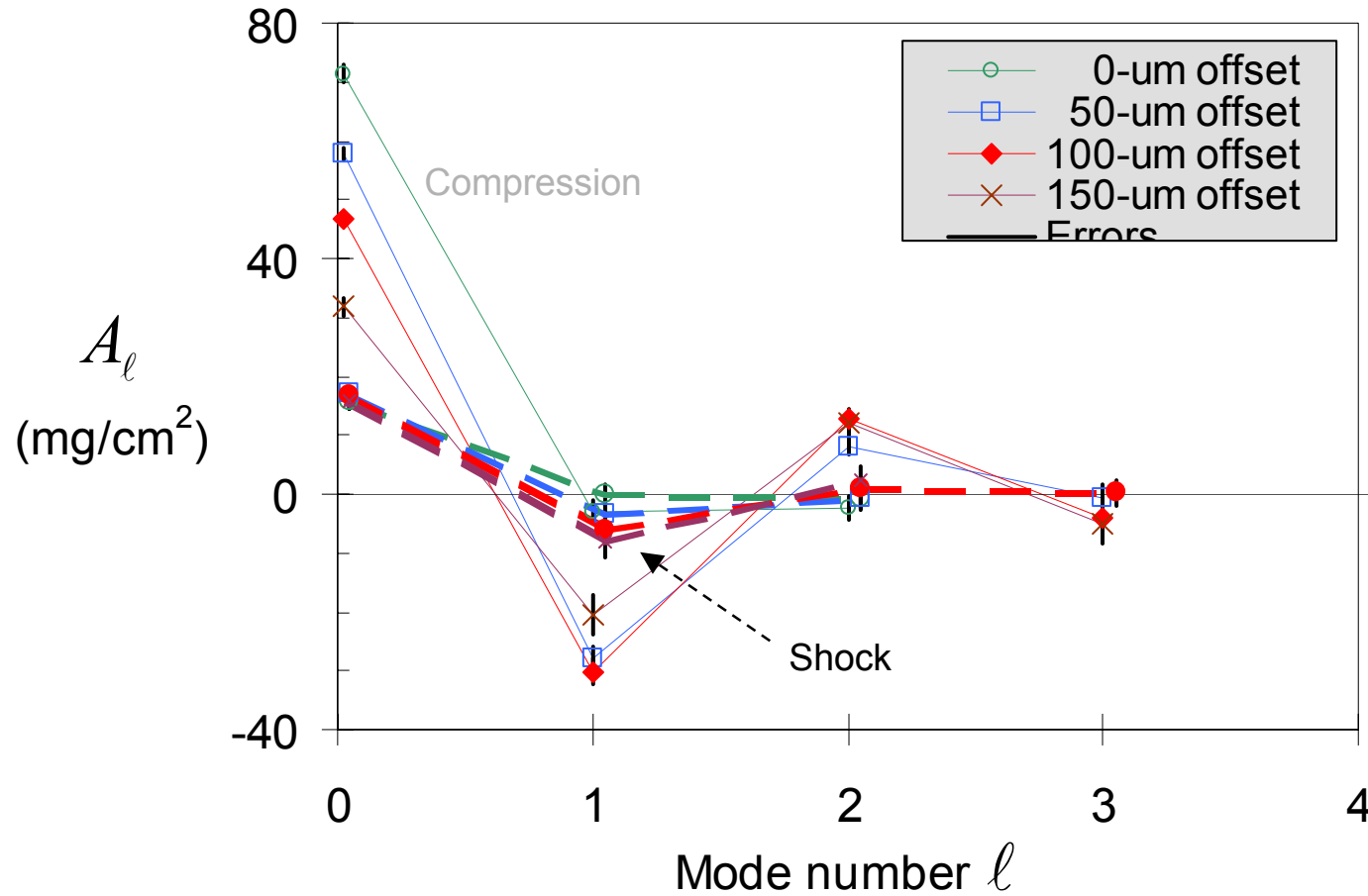
# Another way to show the deduced $\rho R$ vs $\phi$



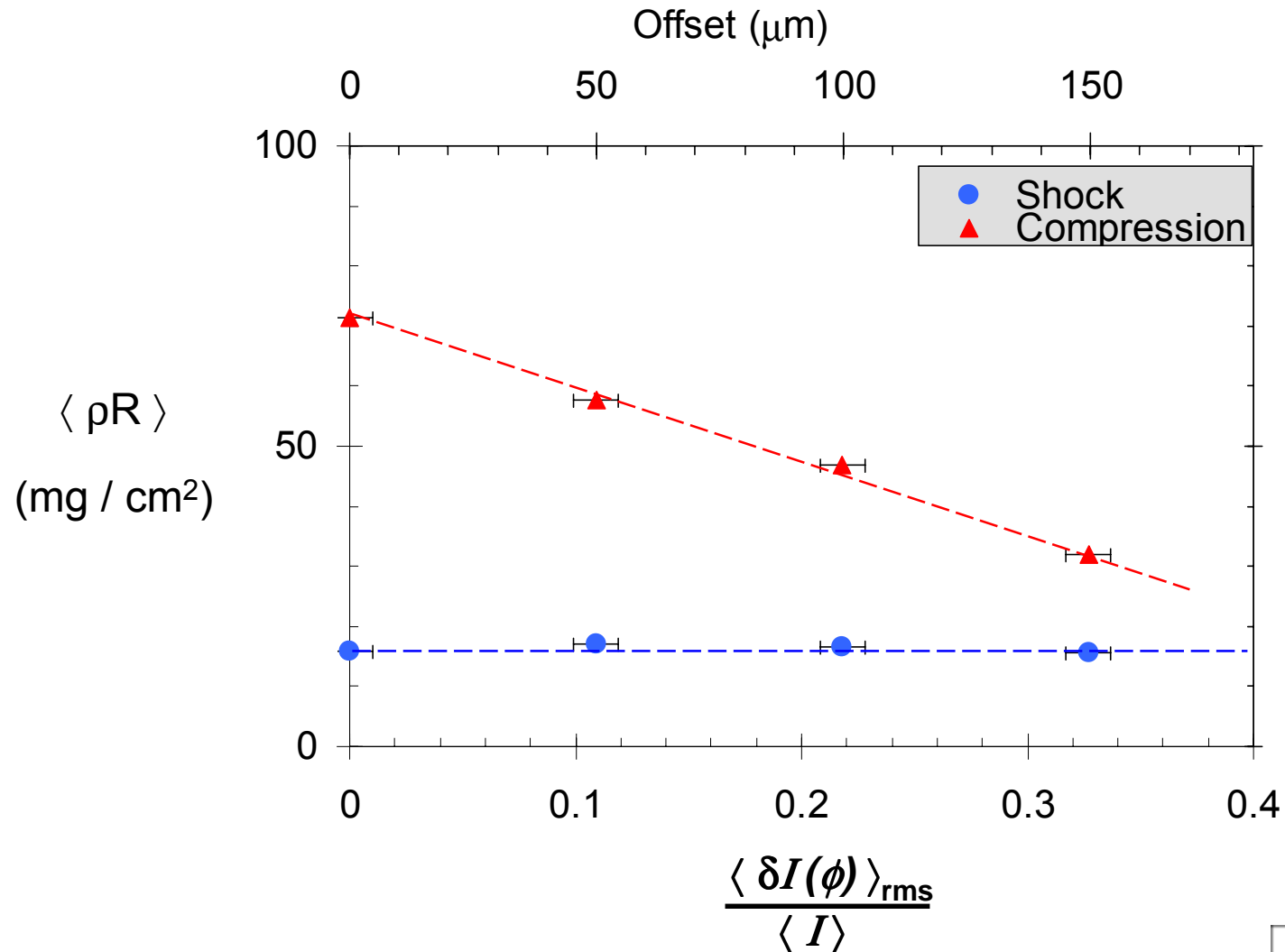
# Mode amplitudes for the fits $\rho R(\phi) = \sum_{\ell} A_{\ell} P_{\ell}(\cos(\phi))$



# Mode amplitudes for the fits $\rho R(\phi) = \sum_l A_l P_l(\cos(\phi))$



$\langle \rho R \rangle$  at shock time is independent of offset,  
while  $\langle \rho R \rangle$  at bang time decreases with offset

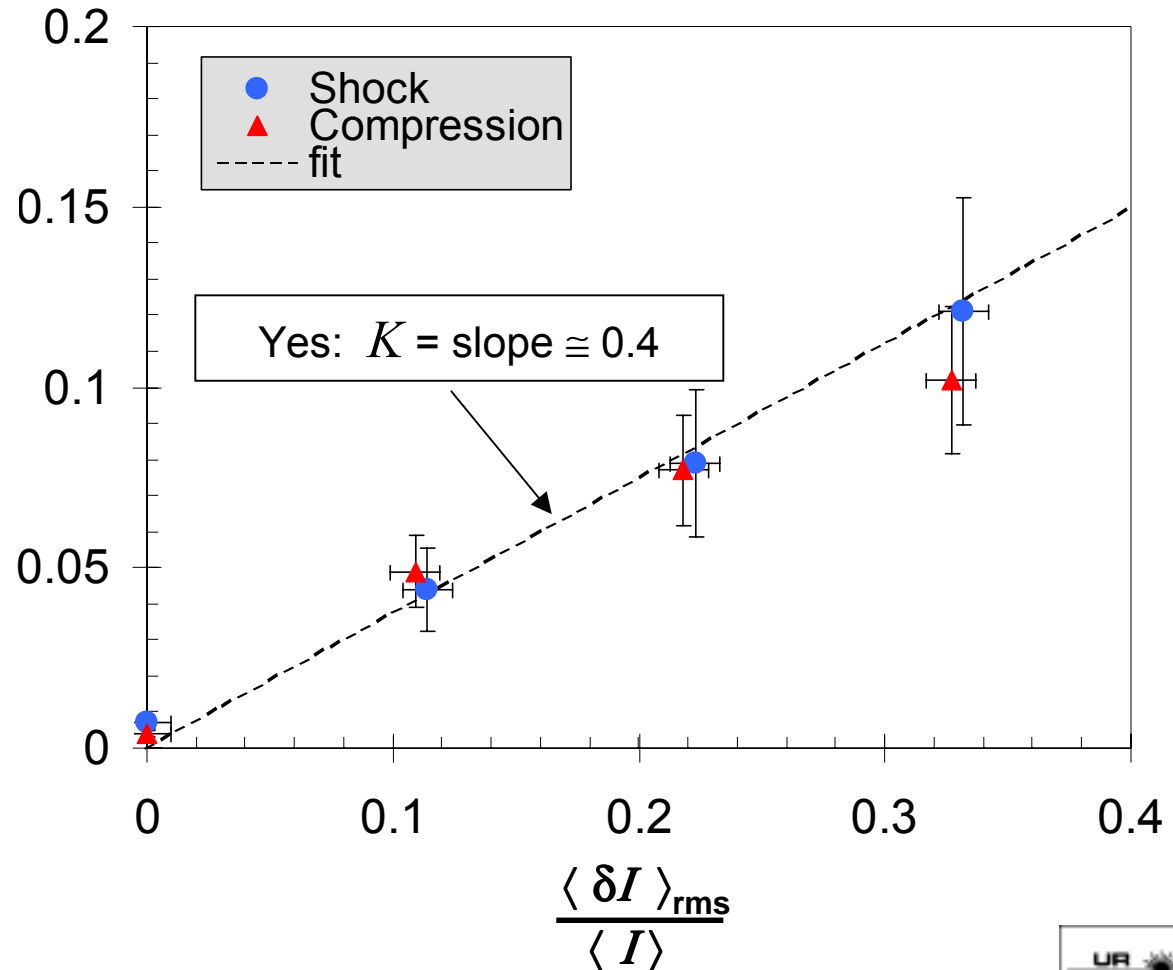




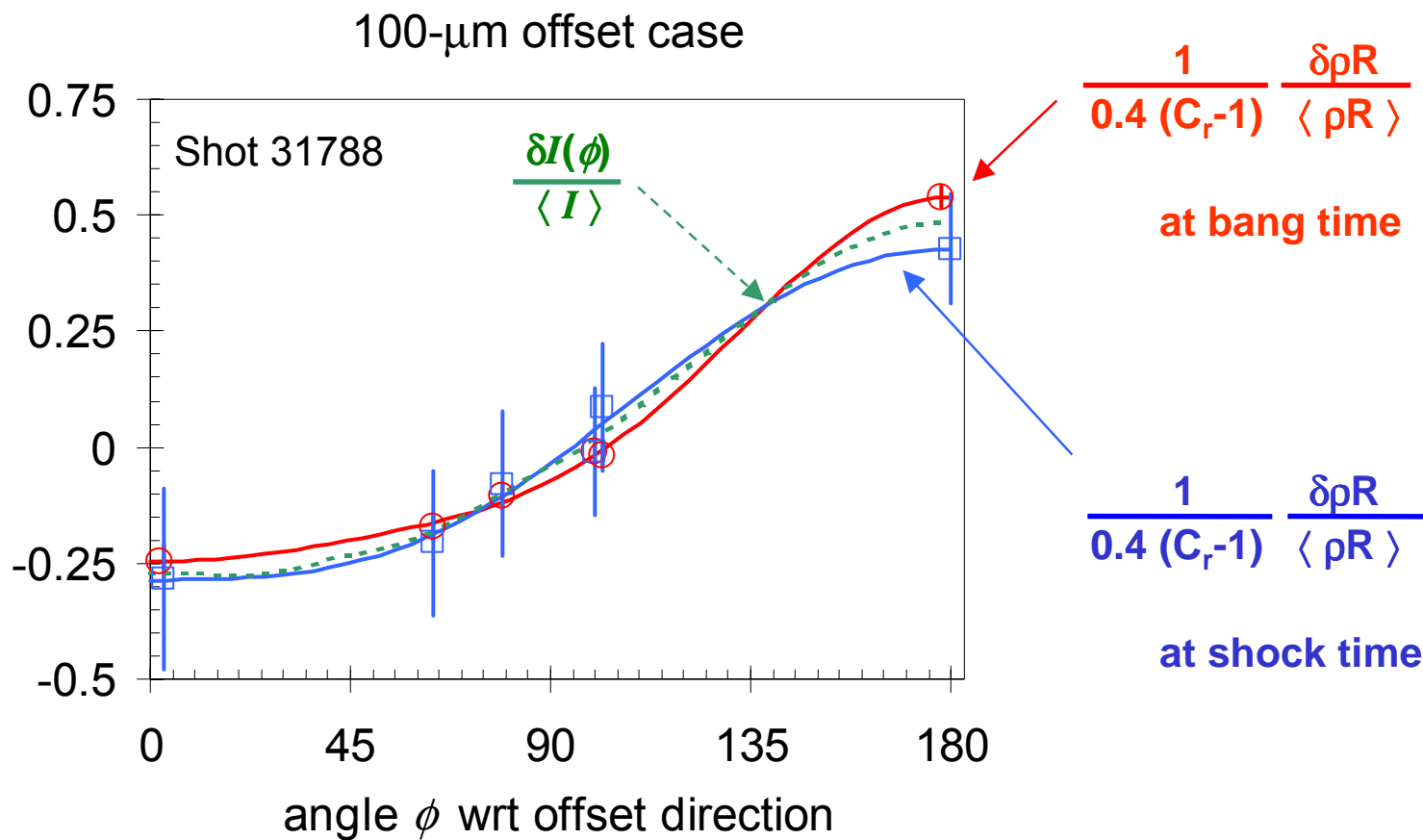
If we average  $\delta\rho R(\phi)$  and  $\delta I(\phi)$  over  $4\pi$ ,  
are they compatible with the previously-proposed scaling

$$\frac{\langle \delta \rho R \rangle_{rms}}{\langle \rho R \rangle} = K (C_r - 1) \frac{\langle \delta I \rangle_{rms}}{\langle I \rangle} \quad ?$$

$$\frac{1}{(C_r - 1)} \frac{\langle \delta \rho R \rangle_{rms}}{\langle \rho R \rangle}$$

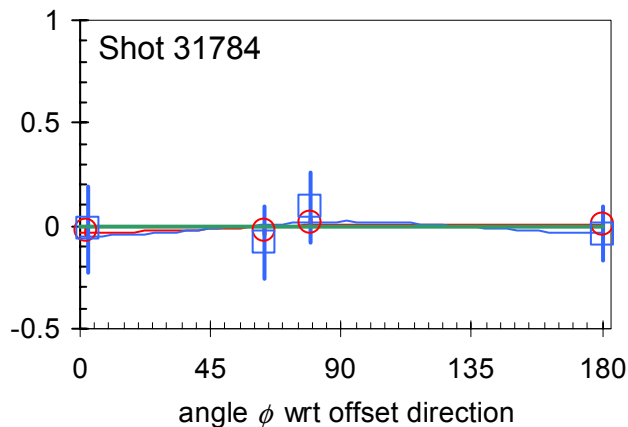


# The angular dependence of the laser drive, $I(\phi)$ , is carried through to $\rho R(\phi)$ at shock time and at bang time

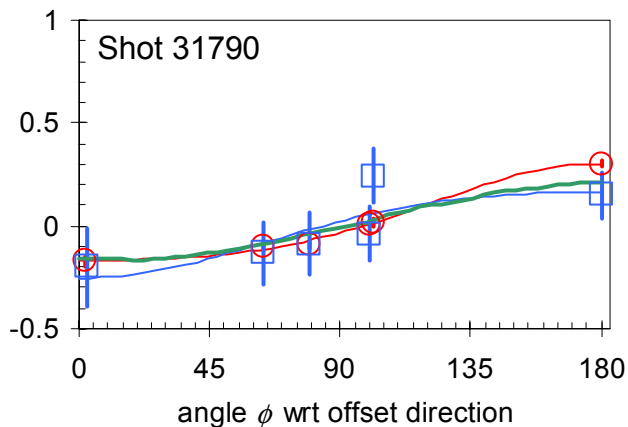


# The angular dependence of the laser drive, $I(\phi)$ , is carried through to $\rho R(\phi)$ at shock time and at bang time

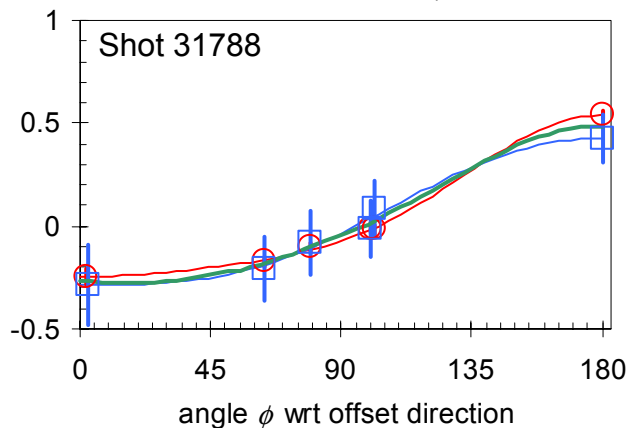
Offset = 0



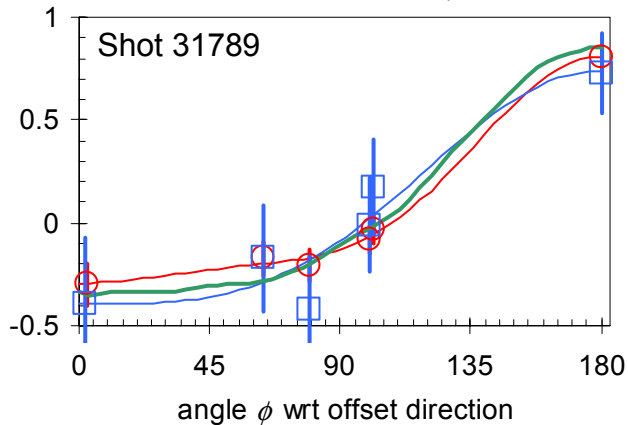
Offset = 50  $\mu\text{m}$



Offset = 100  $\mu\text{m}$



Offset = 150  $\mu\text{m}$



$$\frac{\delta I(\phi)}{\langle I \rangle}$$

$$\frac{1}{0.4 (C_r - 1)} \frac{\delta \rho R(\phi)}{\langle \rho R \rangle}$$

at shock time

$$\frac{1}{0.4 (C_r - 1)} \frac{\delta \rho R(\phi)}{\langle \rho R \rangle}$$

at bang time

# Summary

---

For asymmetric OMEGA laser drive  $I(\phi)$  dominated by mode numbers  $\lesssim 3$ , applied to room-temperature capsules with thick CH shells,

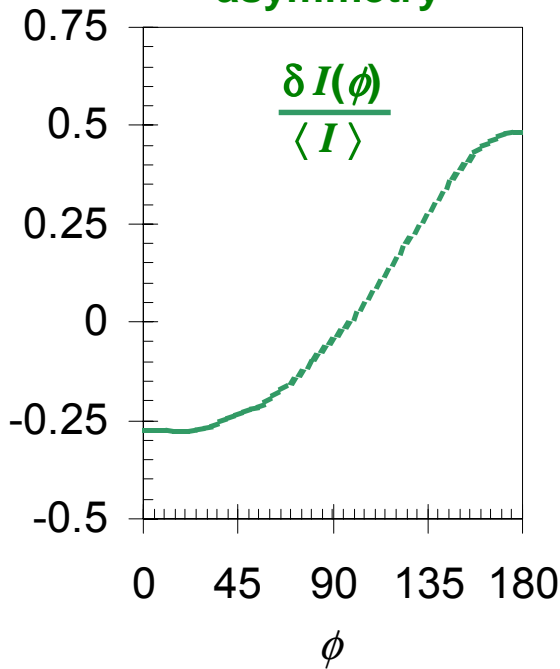
- $\delta I(\phi) = I(\phi) - \langle I \rangle$  produces  $\delta \rho R(\phi) = \rho R(\phi) - \langle \rho R \rangle$  with the same shape;
- $\delta \rho R(\phi)$  maintains that shape throughout the implosion ...
- with amplitude depending primarily on the radial convergence ratio  $C_r$ :

$$\frac{\delta \rho R(\phi)}{\langle \rho R \rangle} \approx 0.4 (C_r - 1) \frac{\delta I(\phi)}{\langle I \rangle}$$

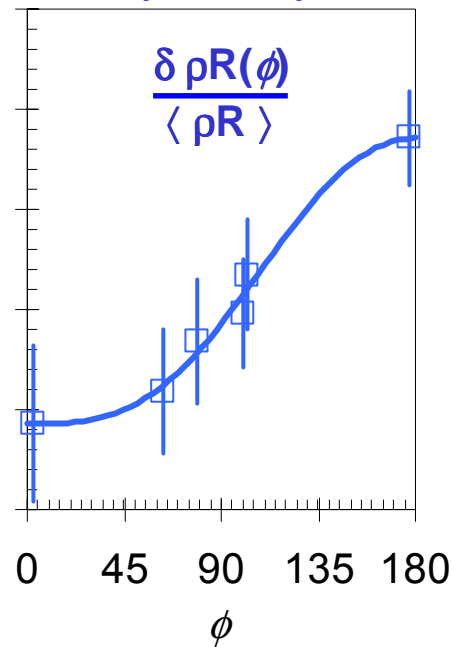
- Modes 1 and 2 grow at the same rate, with no phase inversions.

# Time evolution of $\rho R$ asymmetries in OMEGA direct-drive implosions

Laser drive asymmetry



Areal density at shock time (~1.7 ns)



Areal density at bang time (~2.1 ns)

