

# Recent Work to Improve the Omega Laser Facility's VISAR and Streaked Optical Pyrometer Diagnostics



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### VISAR\* and SOP\*\* use a common telescope to acquire data from laser-compressed samples

- 590- to 850-nm light from the shock front is imaged onto a streak camera
- Spatial and temporal data are collected simultaneously with a VISAR
- The brightness temperature is inferred from self-emission intensity using the absolute calibration

\*Velocity interferometer system for any reflector  
\*\*Streaked optical pyrometer

### Update on VISAR and SOP diagnostics

- In response to users' requests, VISAR and SOP will have periodic timing and calibration tests
  - a T-0 test will be performed twice a year
  - SOP calibration will be performed twice a year
- Telescopes and optical relays on OMEGA and OMEGA EP will be redesigned
  - improve SOP optical performance in 590- to 850-nm ranges
  - install baffles to reduce throughput of light from outside FOV\*
  - mount SOP calibration on telescope
- OMEGA EP TIM\*\*-12 upgrade
  - install TIM periscope to improve alignment for off-axis campaigns

\*FOV: field of view  
\*\*TIM: ten inch manipulator

### The target is driven on the surface facing the ASBO\* telescope and an etalon delay is inserted into one leg of the interferometer

- ASBO detects plasma emission at  $\lambda = 532$  nm
- No probe laser (no fringes)
- Long etalon delays  $\Delta t$  (234 ps; 311 ps)

\*ASBO: active shock breakout

### ASBO T-0 is defined as the relative timing between the 2% rise of emission from the laser pulse and the centroid of the first fiducial

### The ASBO signal is modeled as time-integrated laser energy modified by two exponential decays

$$\text{Signal} = \int_0^t e^{-c(t-t_1)^2} dt \left[ A e^{-b(t-t_1)} + B e^{-b(t-t_2)} \right] + R \int_0^t e^{-c(t-t_2)^2} dt \left[ A e^{-b(t-t_2)} + B e^{-b(t-t_3)} \right] + C$$

Drive term      Two-component decay      Delayed duplicate signal

\*FWHM: full width at half maximum

### The error in the T-0 measurement is within the timing resolution of the P820 streak camera

### An upgraded telescope will improve SOP imaging, calibration, and VISAR alignment

**Imaging**

- The current telescope is only optimized for 532-nm light, which causes the SOP to have imaging capabilities that are nearly 40x the diffraction limit

**Calibration**

- The current calibration source cannot be repeatedly installed and aligned
- The calibration source is only available on OMEGA

**Alignment**

- The design of the current telescope inhibits the ability to accurately point and center through the three primary optics
- Additional alignment aids will be developed to quantify and improve alignment

Upgrading the telescope gives an opportunity to make many improvements to the VISAR/SOP system as a whole.

### Achromat design performance over the 590- to 850-nm SOP spectral band is near-diffraction limited

rms: root mean square

### A VISAR telescope will meet all optical and mechanical requirements while being modular and easily accessible

- The new VISAR telescope is under design to replace the existing one. It must meet all optical and mechanical requirements (mainly the requirements defined in D-EA-G-286 Rev. B) while being modular and easily accessible

### The new telescope design is driven from past diagnostic performance and operator input

The front-end alignment system is designed to interface any Thorlabs 60-mm cage components

### The current telescope design allows light from outside the desired field of view to be relayed to the streak cameras

### Currently 2.2% of target chamber (TC) scattered light incident on the f/3.3 collection lens is able to exit the telescope

- Energy loss mechanisms
  - absorption/total internal reflection (TIR) by lenses
  - absorption by metal

### The baffled telescope allows 0.011% of TC scattered-light energy incident on the collection lens to exit the telescope

- Improved by a factor of 200

### Mounting the calibration source to the telescope will improve alignment repeatability, which will decrease absolute calibration error

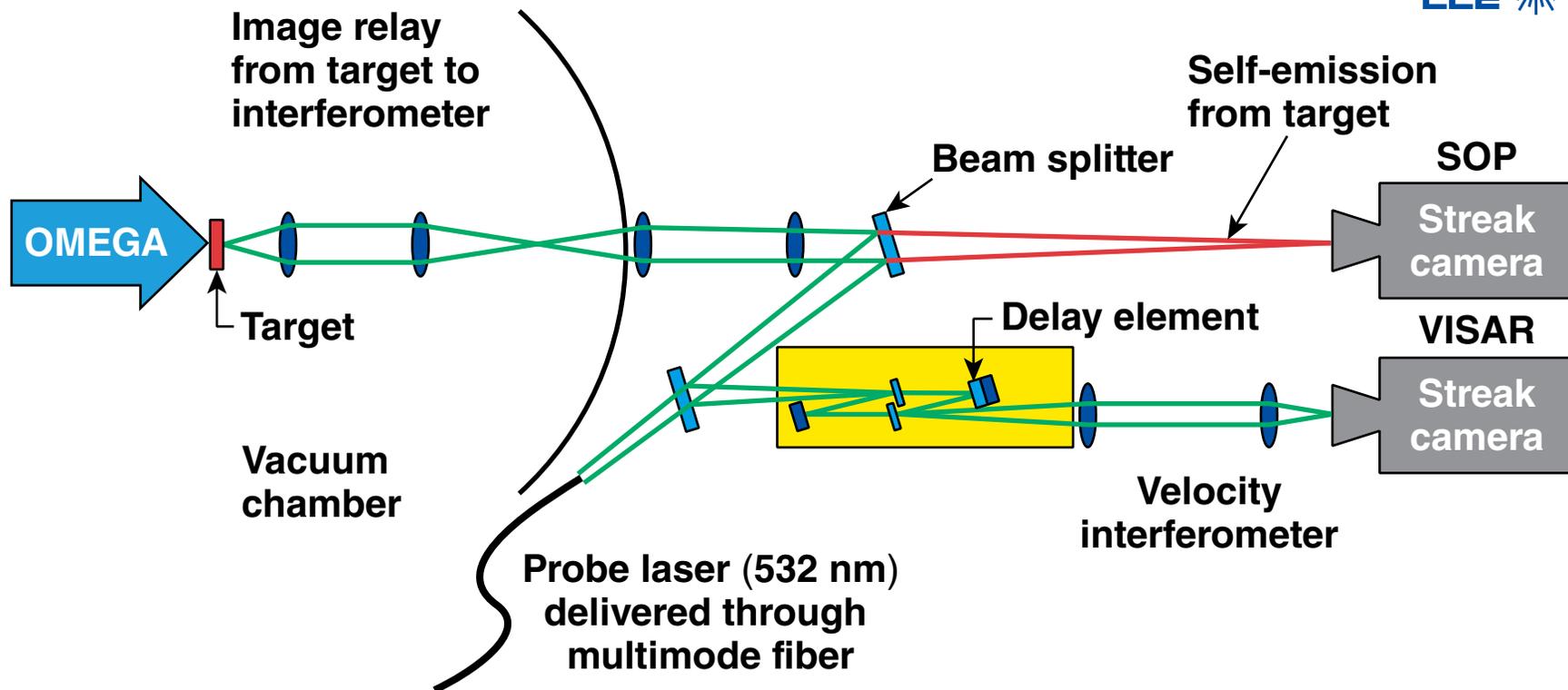
- The bulb-mount assembly clears the TIM flapper valve and all TC instrumentation provided the target positioners and P11 neutron temporal diagnostic (NTD) are retracted
- The bulb will be adjustable in x, y, z, and  $\theta$
- Off-line filament alignment will be staged in DEL\* ( $\pm 1$  mm in x, y, z, and  $\theta$ )
- SOP focus on the filament is fine tuned with the TIM insertion depth
- Cable strain relief points will be located on the socket, telescope tube, and telescope body

\*DEL: Diagnostic Evaluation Laboratory

### A TIM-12 periscope will allow for accurate pointing and centering down the telescope axis

This periscope is critical to telescope alignment, specifically for campaigns that are not at target chamber center (TCC).

# VISAR\* and SOP\*\* use a common telescope to acquire data from laser-compressed samples



- 590- to 850-nm light from the shock front is imaged onto a streak camera
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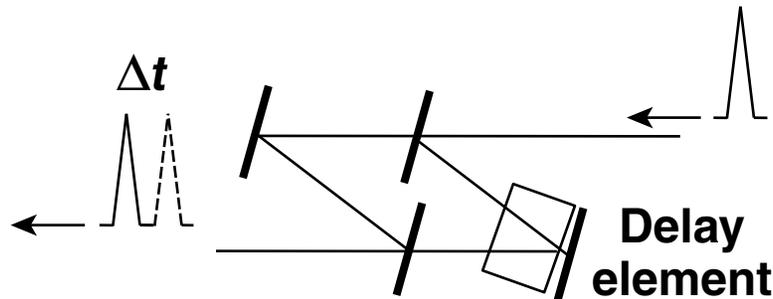
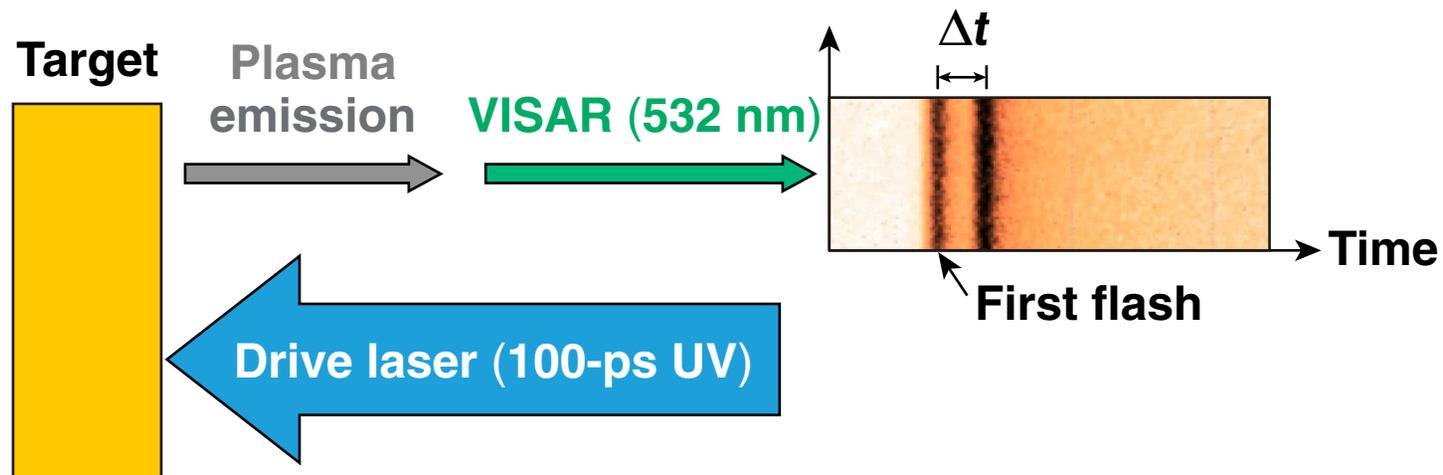
# Update on VISAR and SOP diagnostics

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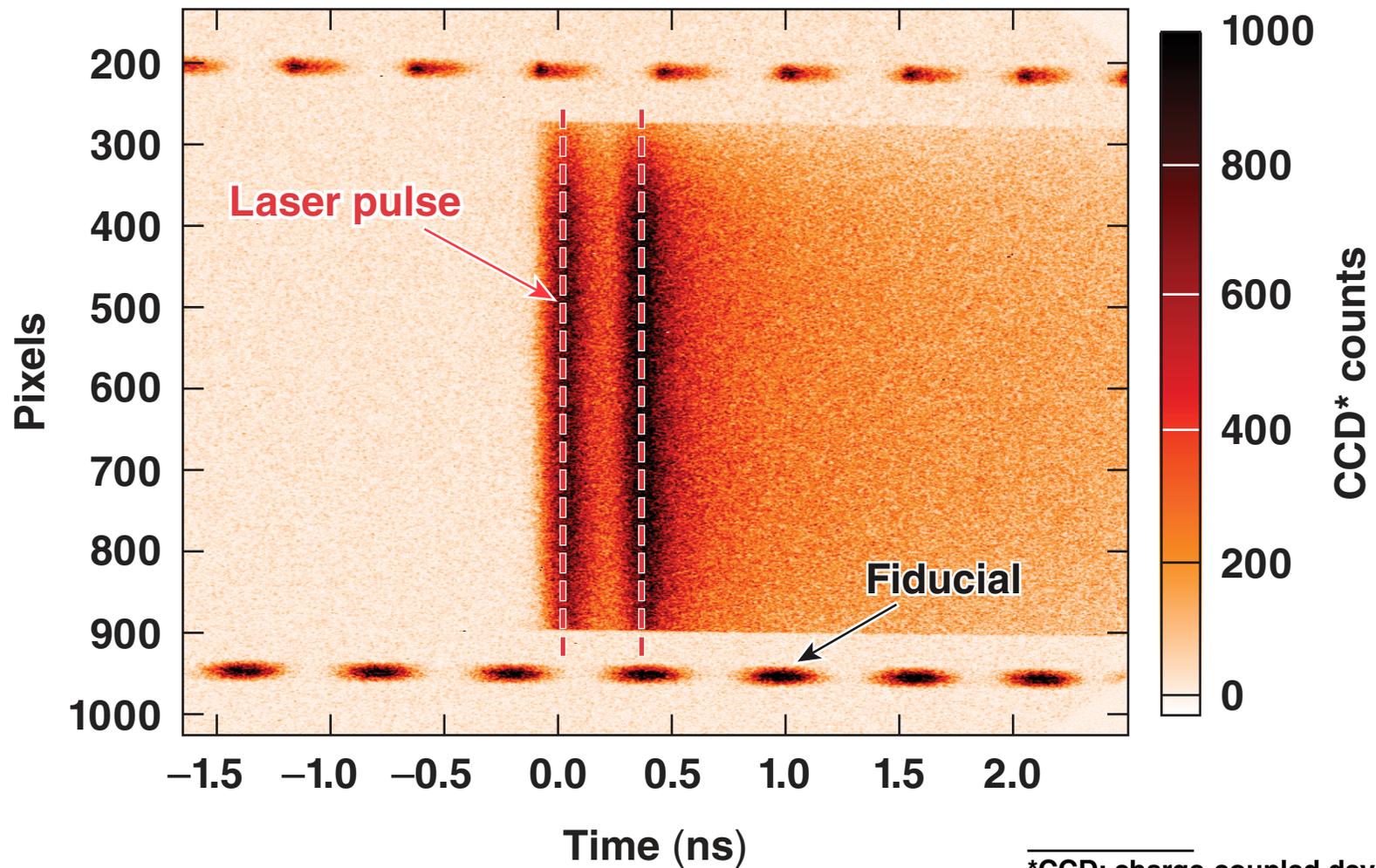
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- **Telescopes and optical relays on OMEGA and OMEGA EP will be redesigned**
  - improve SOP optical performance in 590- to 850-nm ranges
  - install baffles to reduce throughput of light from outside FOV\*
  - mount SOP calibration on telescope
- **OMEGA EP TIM\*\*-12 upgrade**
  - install TIM periscope to improve alignment for off-axis campaigns

The target is driven on the surface facing the ASBO\* telescope and an etalon delay is inserted into one leg of the interferometer

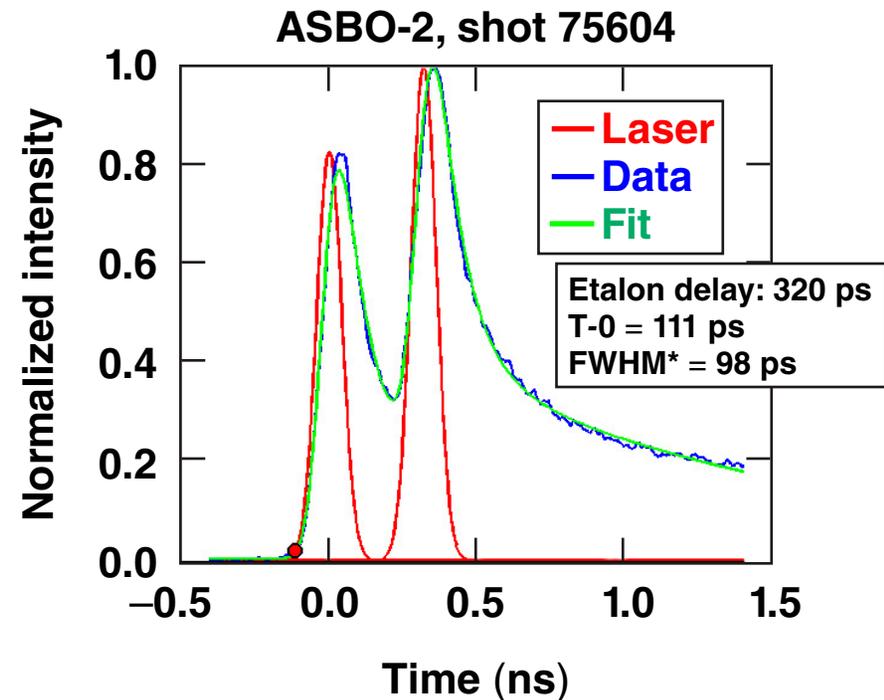
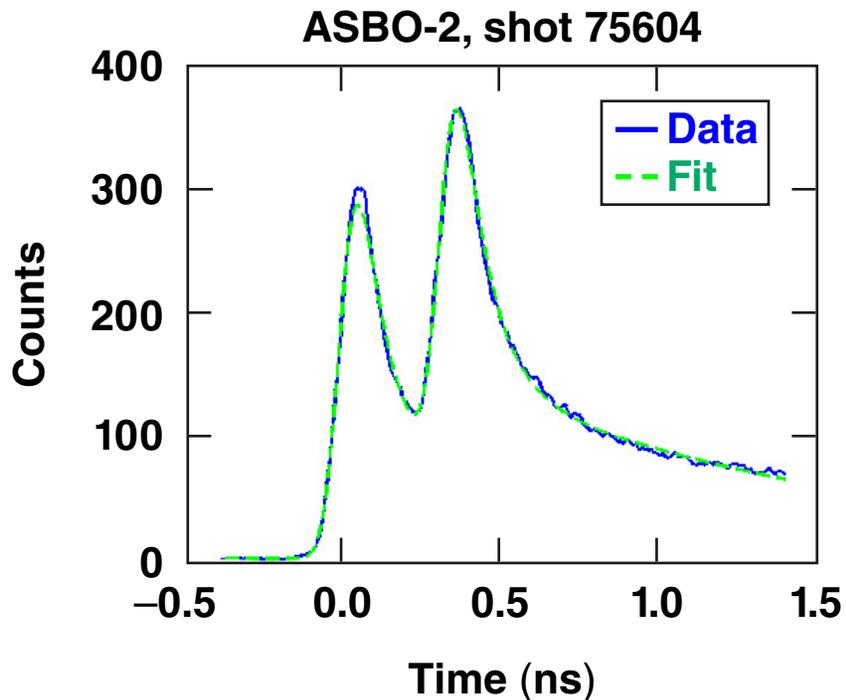


- ASBO detects plasma emission at  $\lambda = 532 \text{ nm}$
- No probe laser (no fringes)
- Long etalon delays  $\Delta t$  (234 ps; 311 ps)

**ASBO T-0 is defined as the relative timing between the 2% rise of emission from the laser pulse and the centroid of the first fiducial**



# The ASBO signal is modeled as time-integrated laser energy modified by two exponential decays



$$\text{Signal} = \underbrace{\int_0^t e^{-c(t-t_1)^2} dt}_{\text{Drive term}} \left[ \underbrace{Ae^{-a(t-t_1)} + Be^{-b(t-t_1)}}_{\text{Two-component decay}} \right] + R \underbrace{\int_0^t e^{-c(t-t_2)^2} dt}_{\text{Delayed duplicate signal}} \left[ Ae^{-a(t-t_2)} + Be^{-b(t-t_2)} \right] + C$$

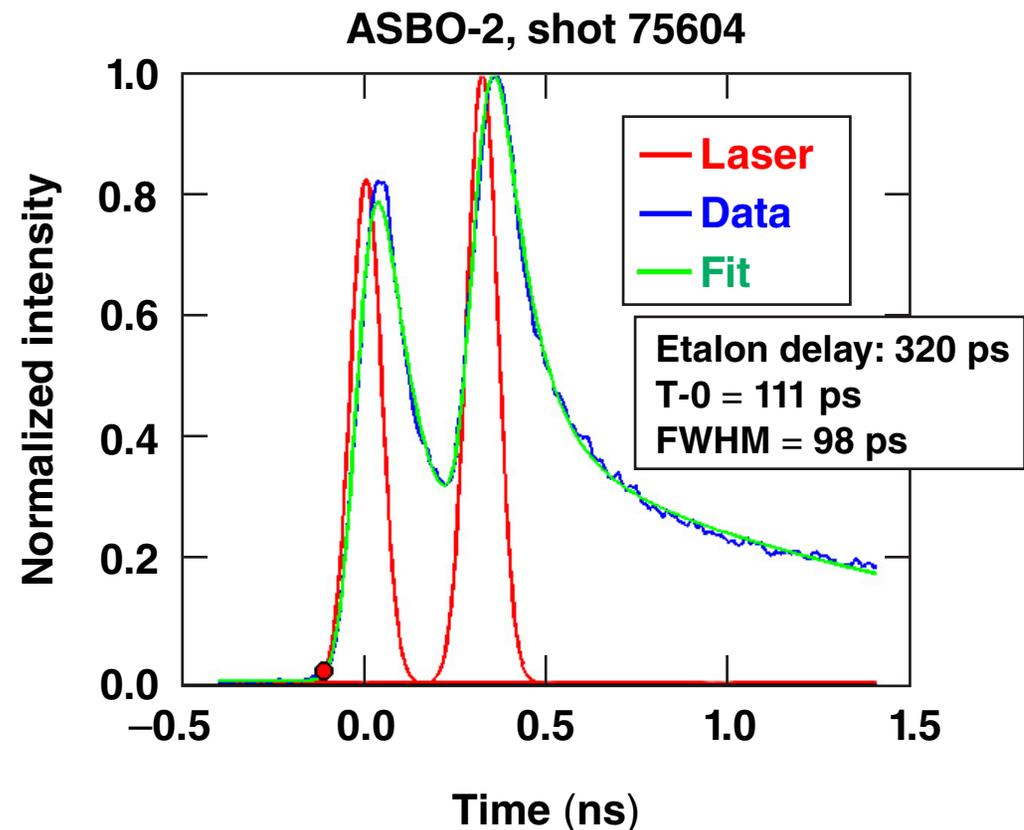
\*FWHM: full width at half maximum

# The error in the T-0 measurement is within the timing resolution of the P820 streak camera

- T-0 is defined as the relative timing between the 2% rise of the Gaussian fit of the emission and the centroid of the first fiducial
- Known etalon delay: 311 ps
- Measured etalon delay: 320 ps
- Timing resolution P820: 0.51% of sweep speed

Sweep speed: 5 ns  
Time resolution: 25 ps

$T-0 = -111 \pm 9$  ps



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## Imaging

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## Calibration

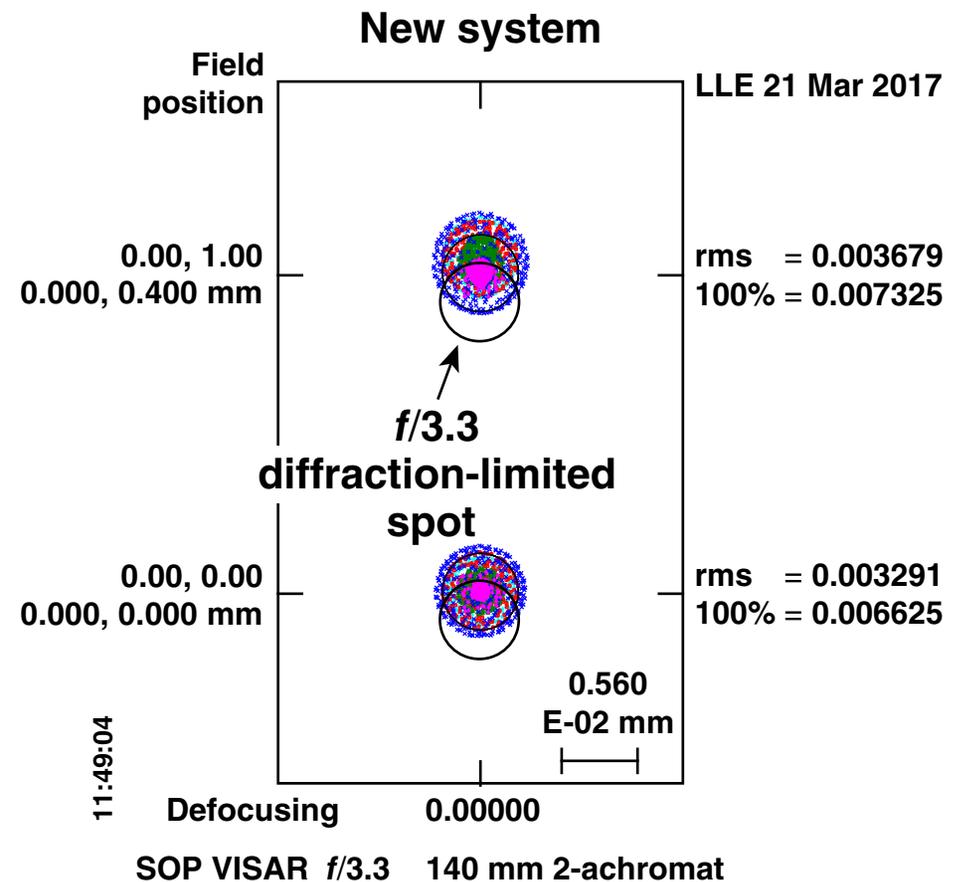
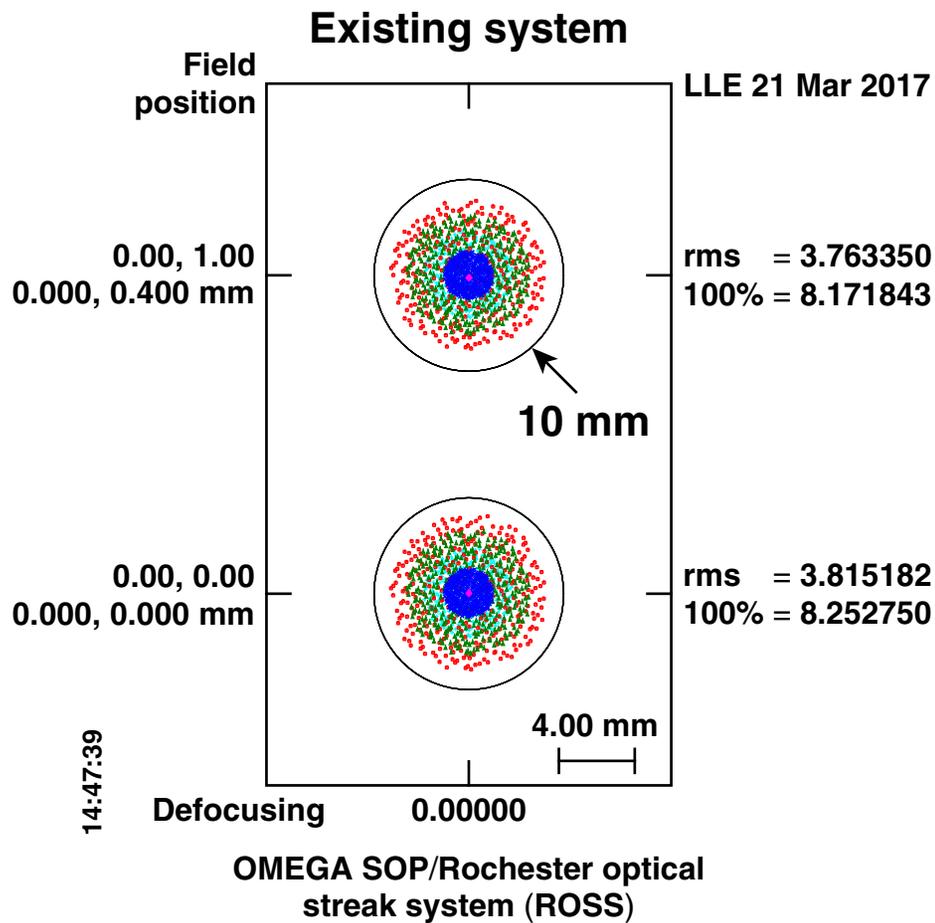
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## Alignment

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**Upgrading the telescope gives an opportunity to make many improvements to the VISAR/SOP system as a whole.**

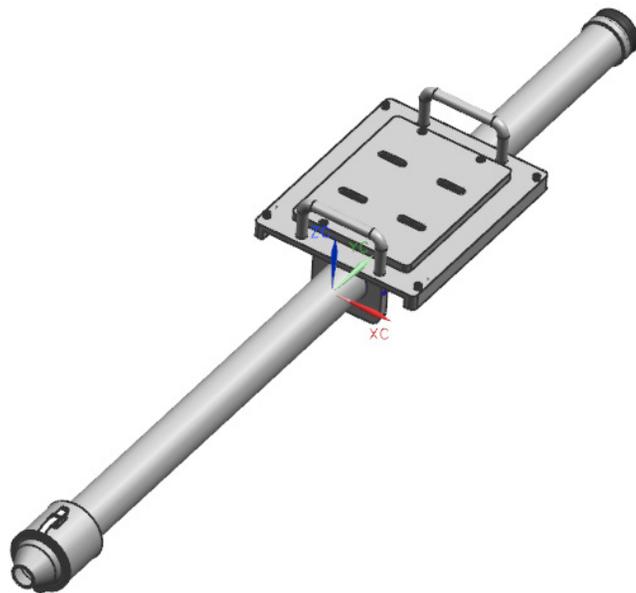
# Achromat design performance over the 590- to 850-nm SOP spectral band is near-diffraction limited



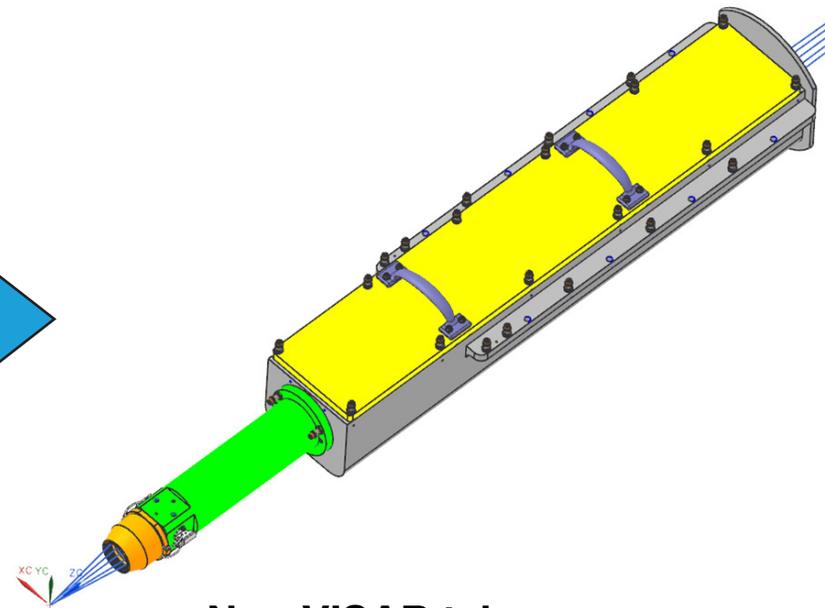
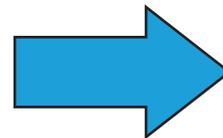
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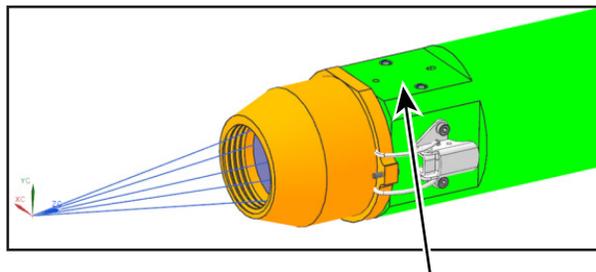
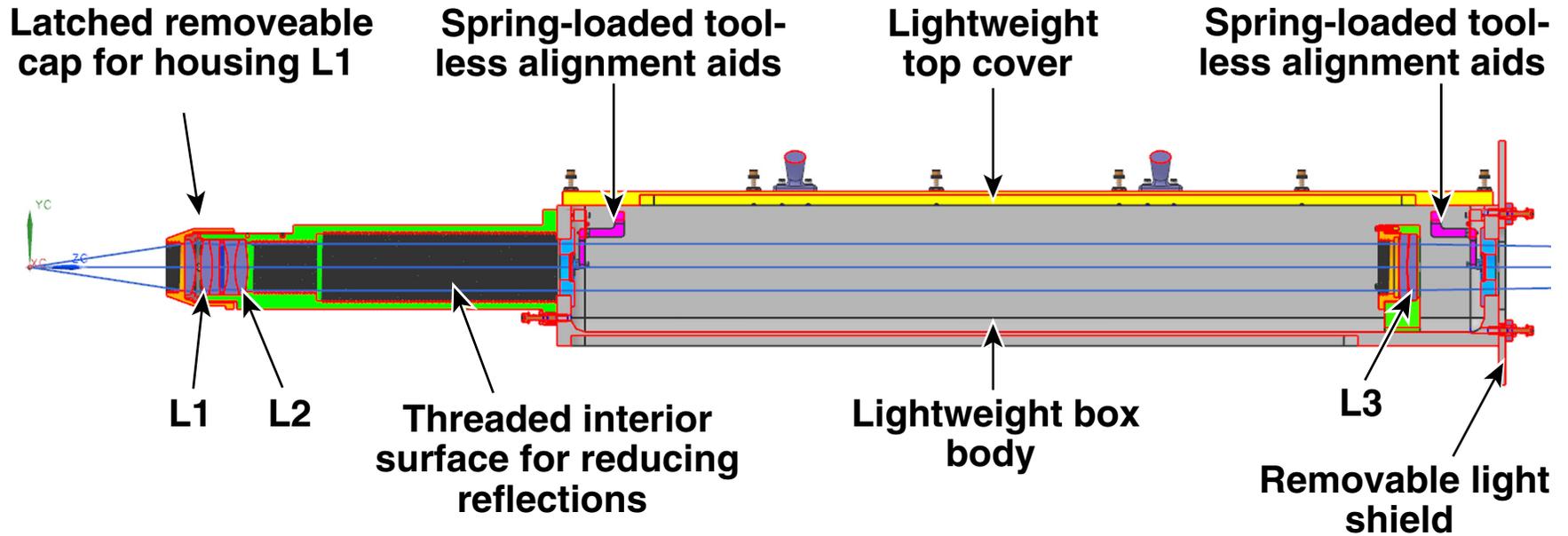


Existing VISAR telescope (D-AS-B-095)

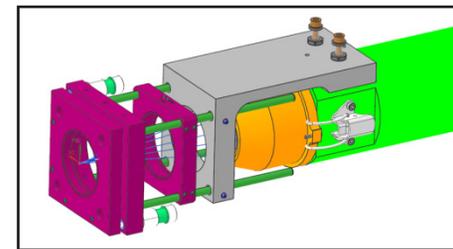


New VISAR telescope

# The new telescope design is driven from past diagnostic performance and operator input

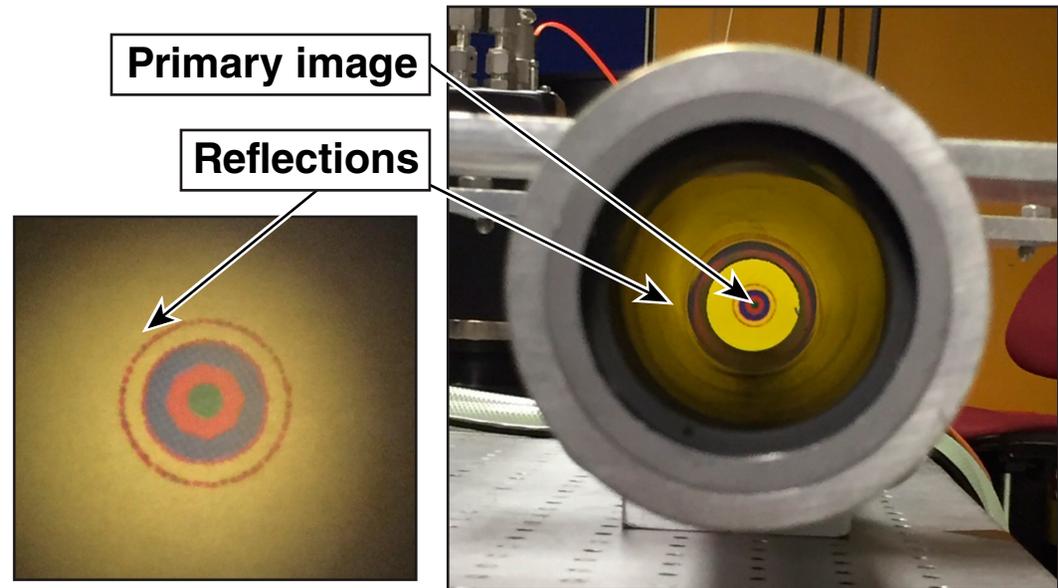
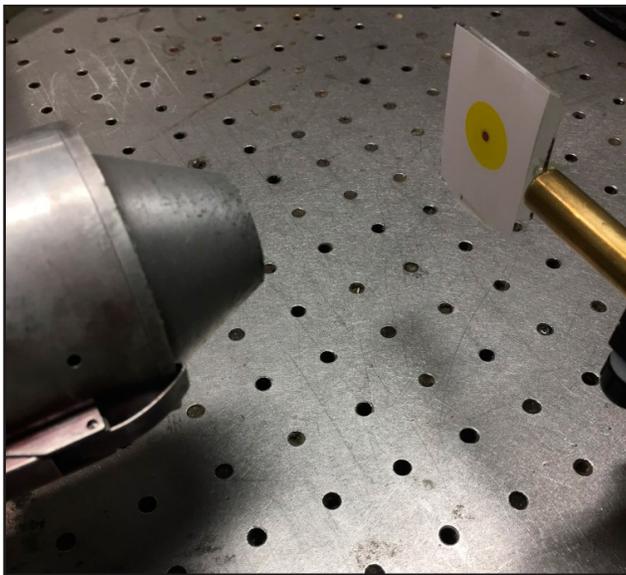
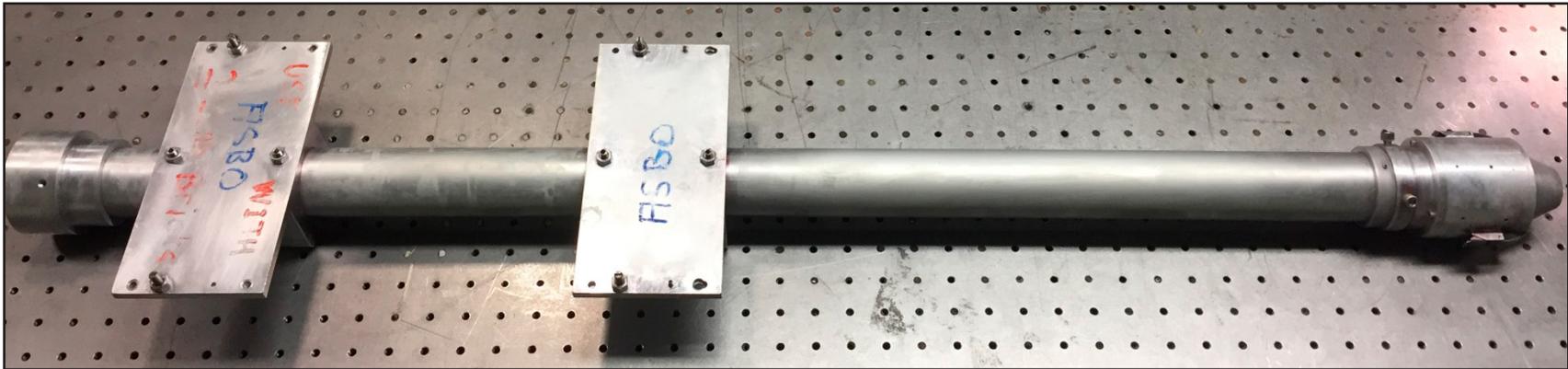


**Bolt pattern for the light-bulb fixture and the alignment tools**



**The front-end alignment system is designed to interface any Thorlabs 60-mm cage components**

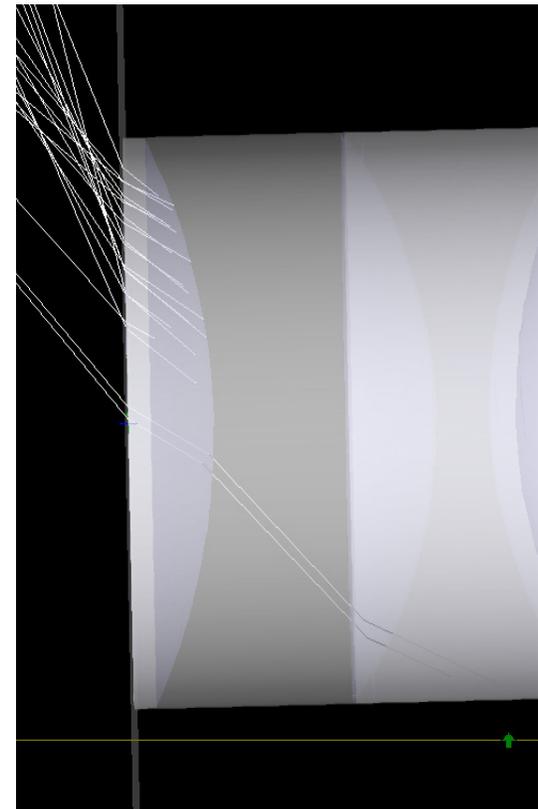
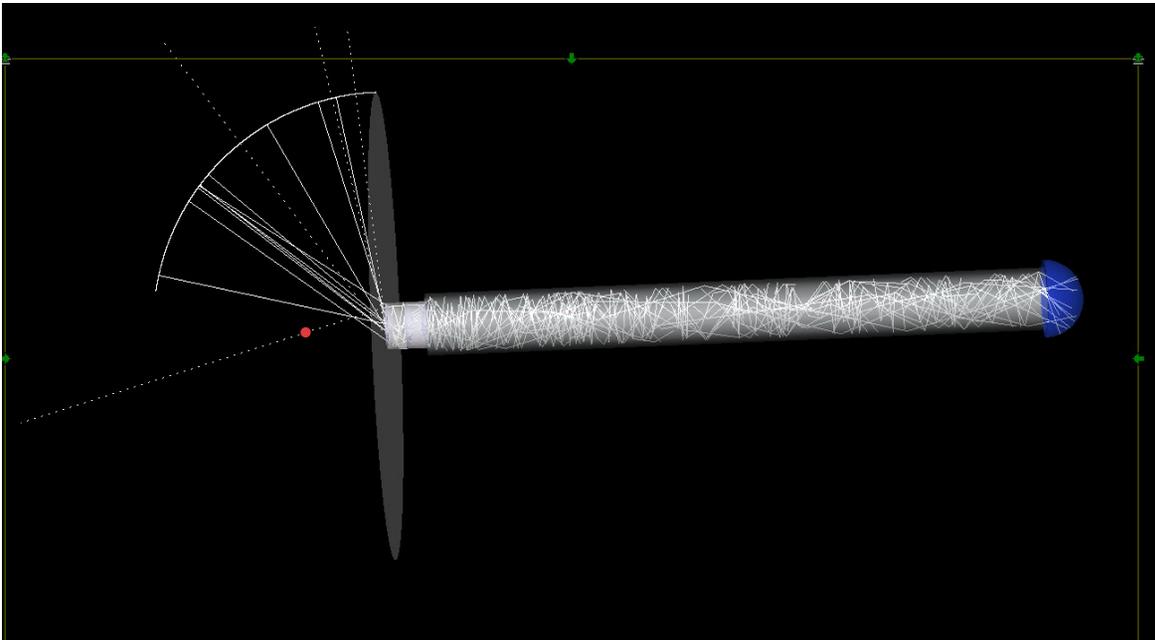
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First image plane

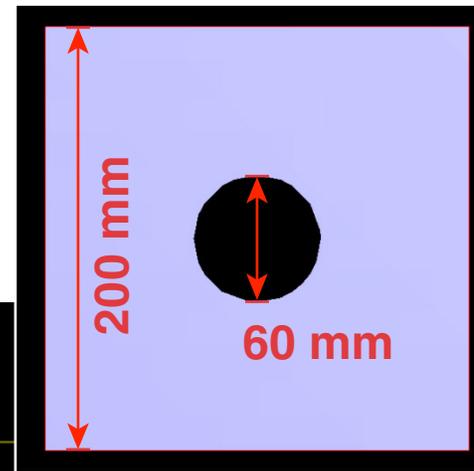
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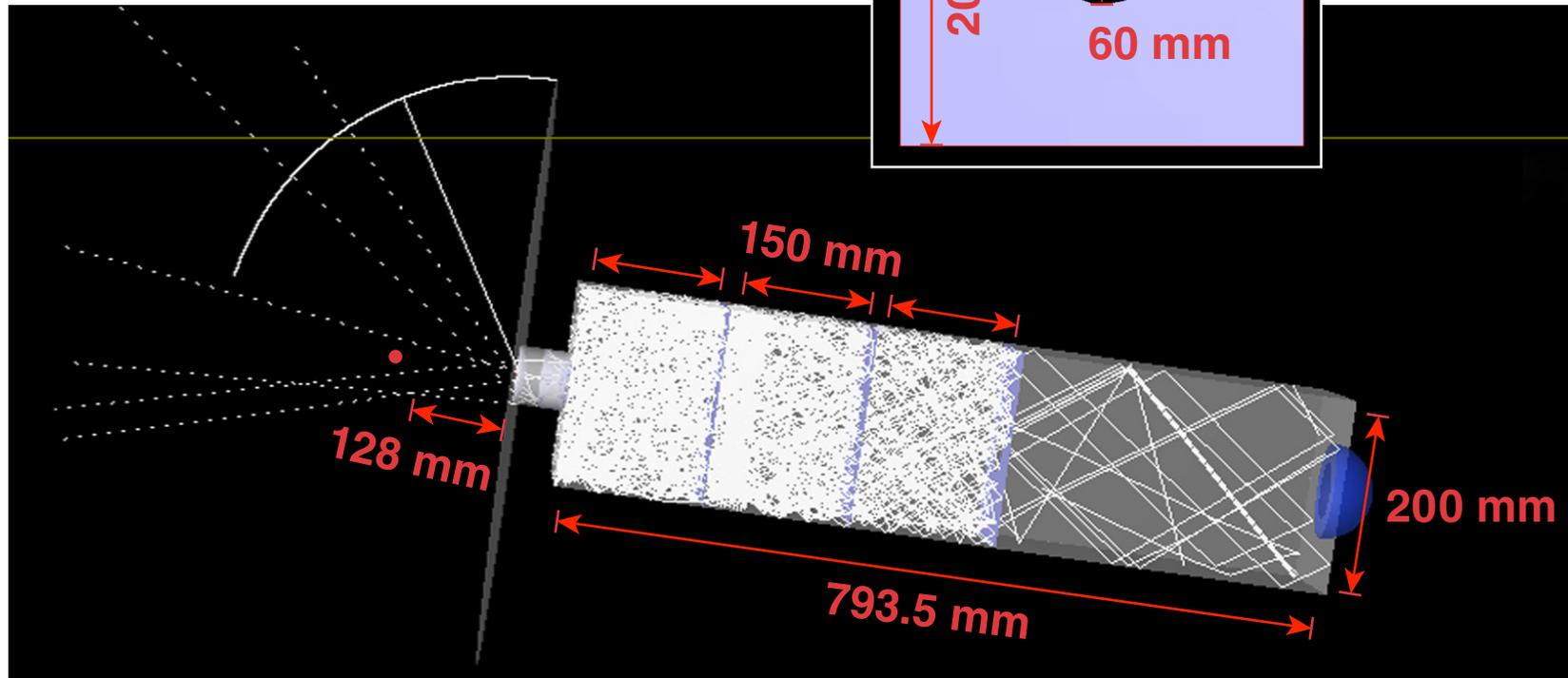


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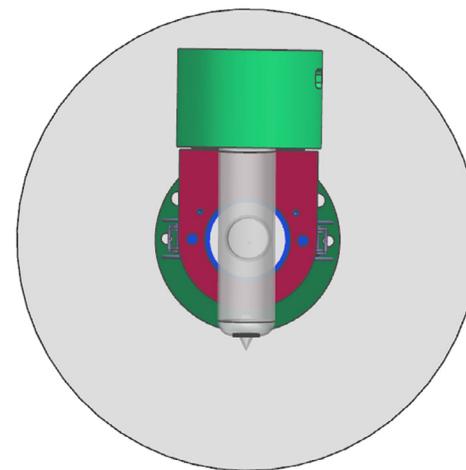
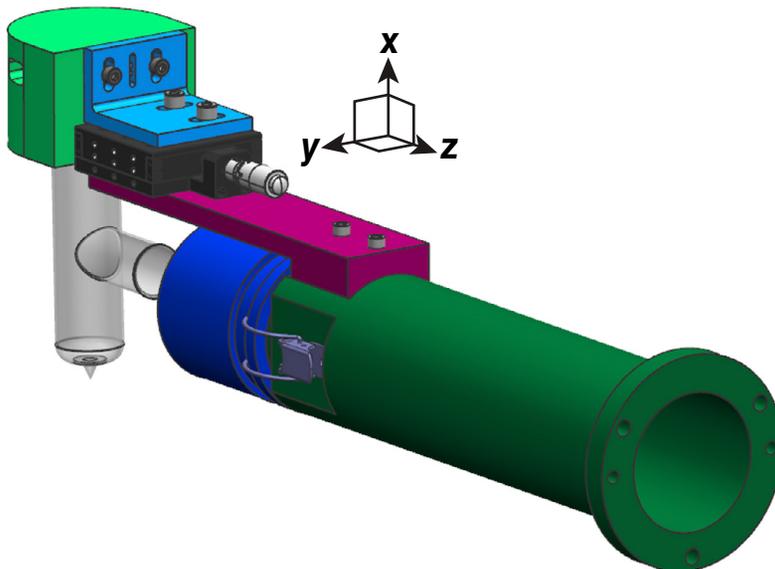
Flat  
baffle



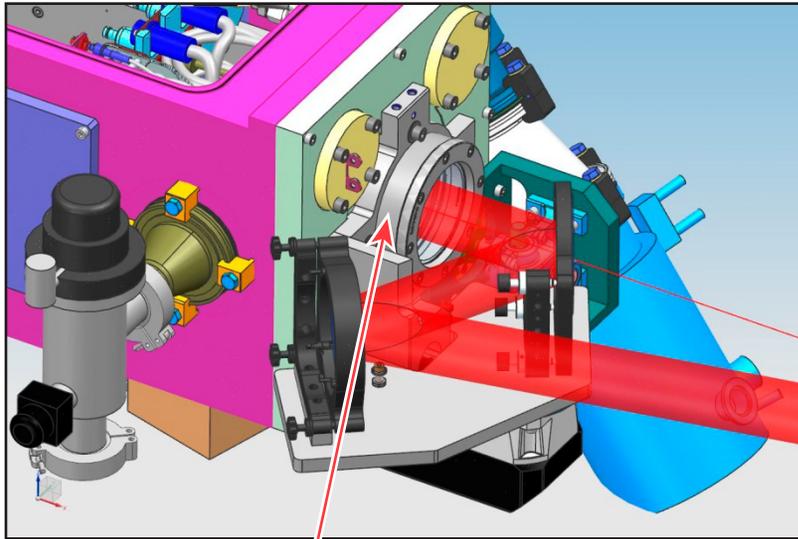
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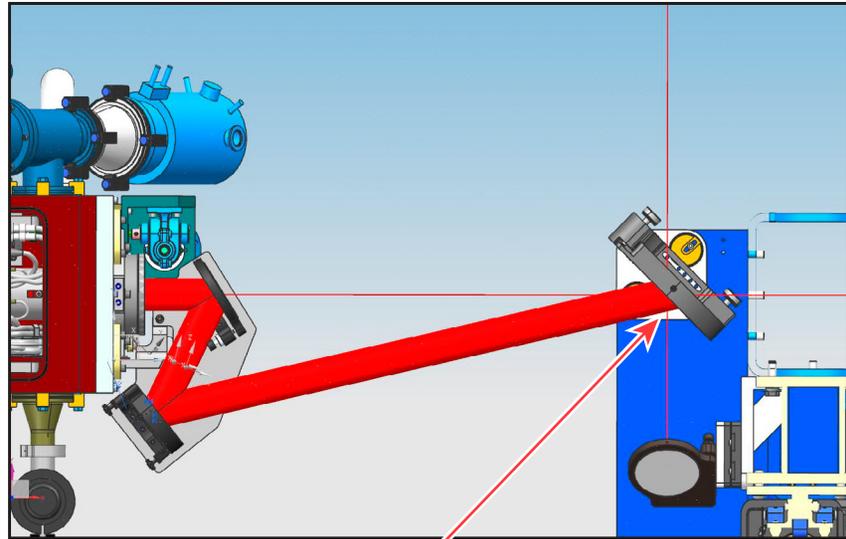
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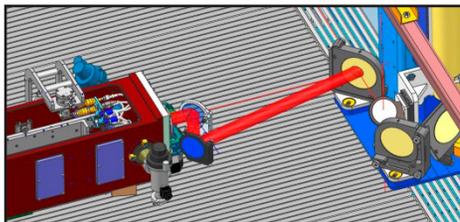
# A TIM-12 periscope will allow for accurate pointing and centering down the telescope axis



**The periscope mirror assembly will mount to the 4-in. rear TIM window**



**A new mounting plate for the 6-in. mirror will be designed to allow the mirror to be pointed at the periscope**



**This periscope is critical to telescope alignment, specifically for campaigns that are not at target chamber center (TCC).**